

# Neutrons and Fundamental Symmetries Experimental III: Other Symmetry Tests

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Slides from J. Barrow, A. Young, M. Snow, H. Shimizu, M. Hurber, H. Abele, ...



# Topics I will cover:

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## Lecture 1: beta-decay

- A brief history of the electroweak theory---the precursor to the Standard Model.
- Neutron decay to test the V-A theory & beyond the SM interactions
- Current status with neutron experiments on  $g_A$  & lifetime
- Physics is Symmetries

## Lecture 2: EDM

- CP violation
- Electric Dipole Moments: Highly sensitive low-energy probes of new Physics
- muon-  $g-2$

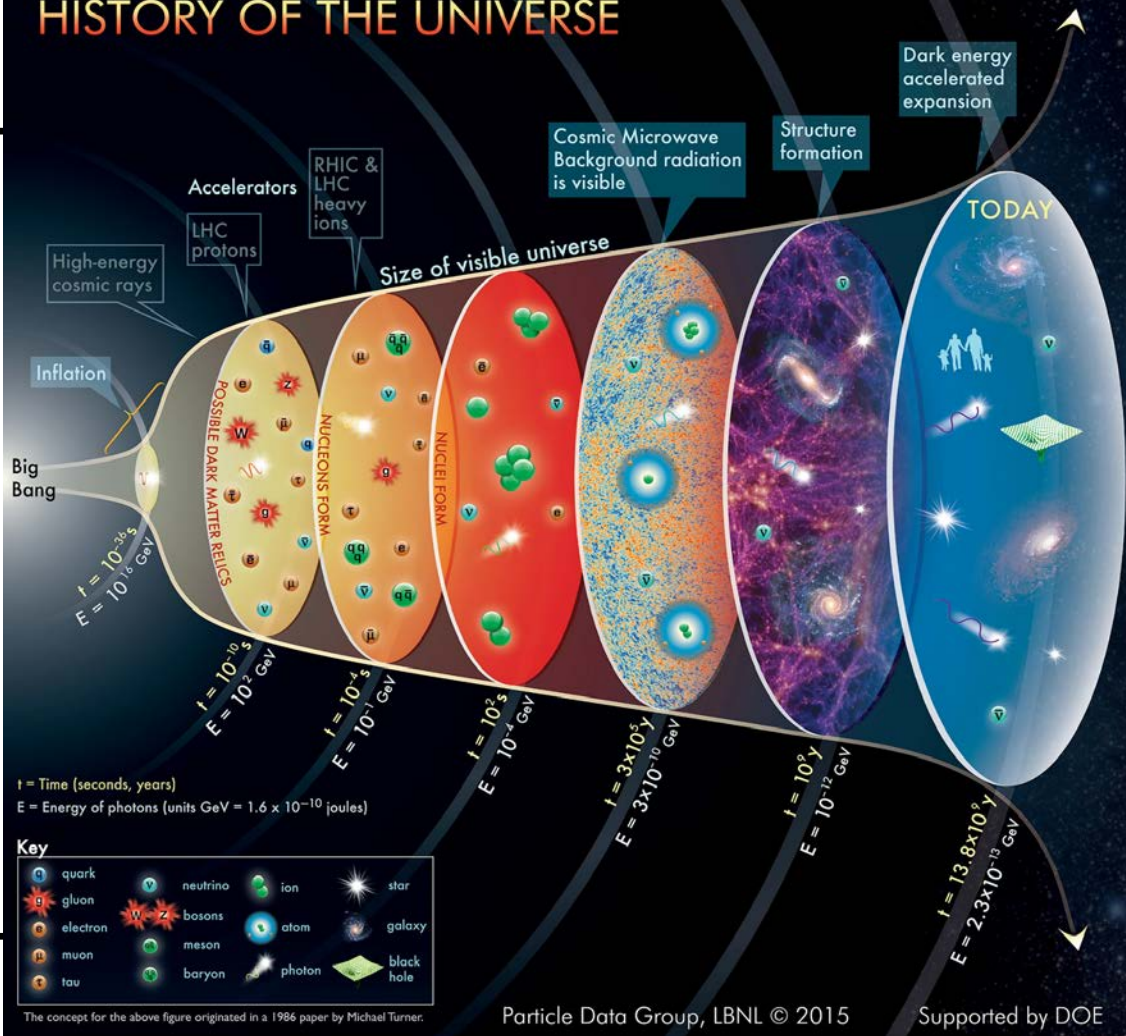
## Lecture 3: other symmetry violation measurements/tests

- Baryogenesis & symmetry violations
- $N_{\bar{b}}$  oscillation: B violation
- Hadronic weak interactions: P violation
- NOPTREX: T violation
- Neutron interferometry: Lorentz symmetry violation

Q: What are the cosmological consequences of symmetry breaking?



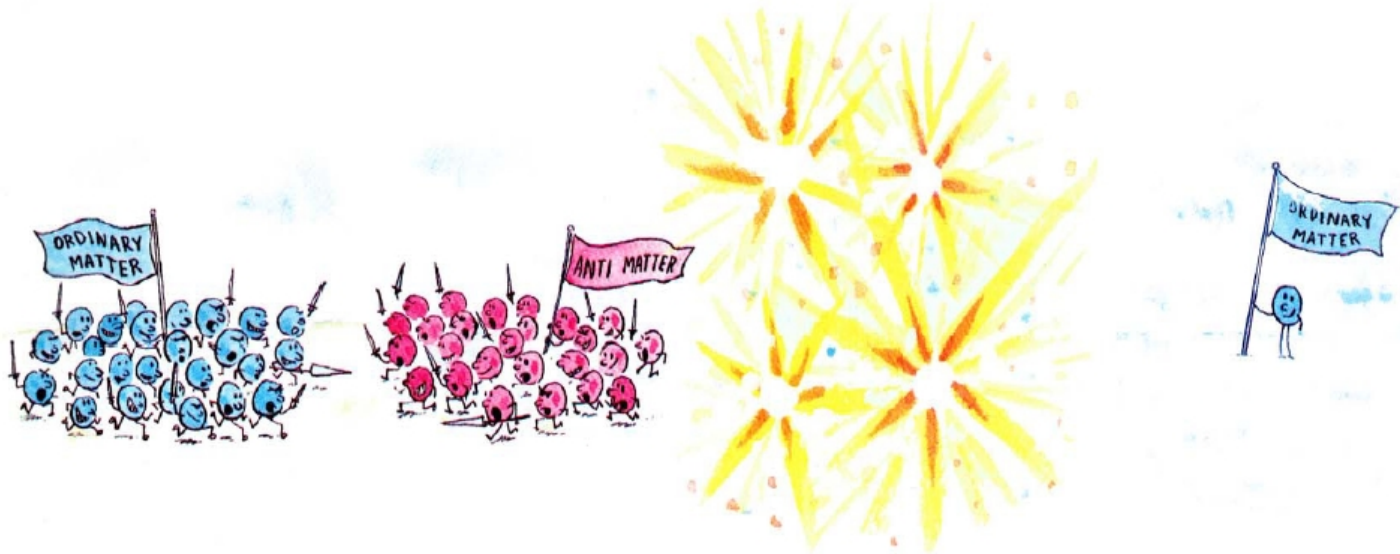
# HISTORY OF THE UNIVERSE



# Matter-Antimatter Asymmetry of the Universe

(or Baryon Asymmetry of the Universe, BAU)

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$$\eta \equiv \frac{\eta_B - \eta_{\bar{B}}}{\eta_\gamma}$$

$\eta$ : baryon asymmetry of the universe (BAU)

$$2.6 \times 10^{-10} < \eta < 6.2 \times 10^{-10}$$



# The Sakharov Conditions

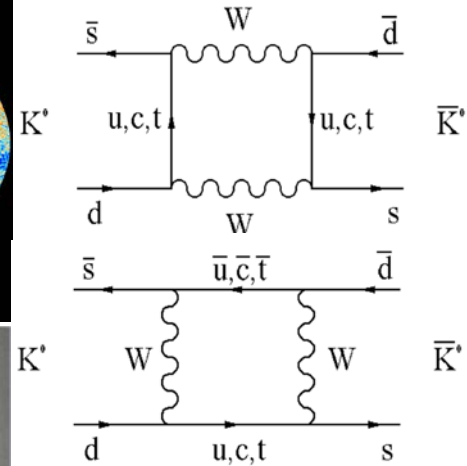
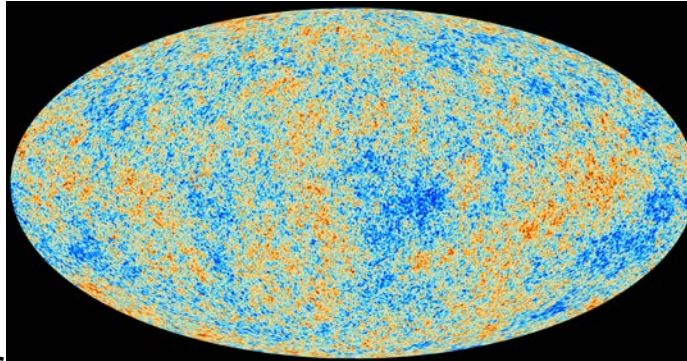
left-handed particle

under C  $\rightarrow$  left-handed antiparticle  
then P  $\rightarrow$  **right-handed antiparticle**

The ***baryon asymmetry of the universe*** motivates **three hypotheses put forward by Sakharov in 1967**

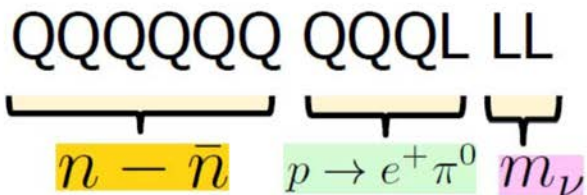
**Any model** attempting to explain our universe must satisfy the following:

- 1. CP violation** (& C non-conservation; different interactions of particles and antiparticles)
  - ***Exists in SM, but the degree of violation might be too small.***
- 2. Departure from thermal equilibrium** (provided by the expansion of the universe)
  - ***Demonstrated from astronomical observations***
- 3. Baryon number (charge)  $B$  violation**
  - ***Still never seen experimentally***

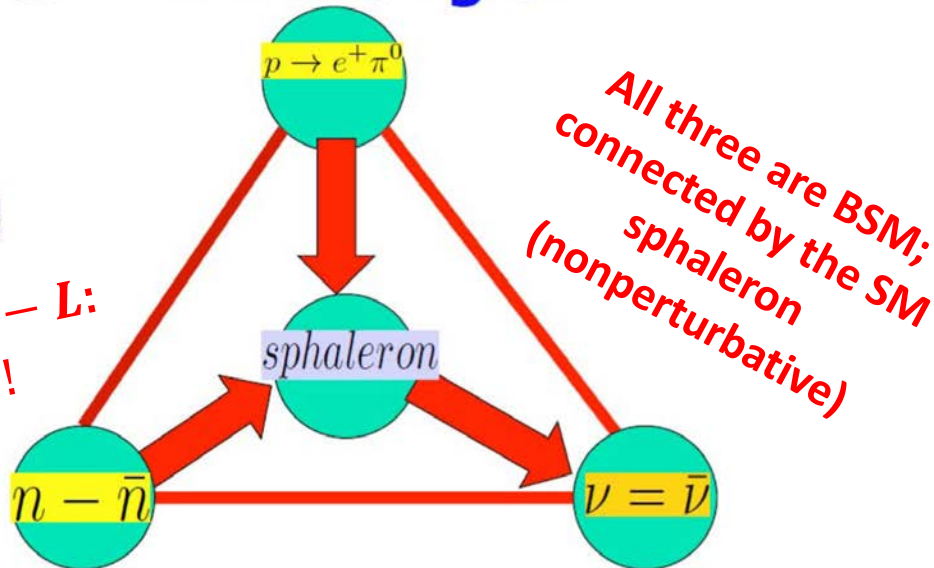


# From NNbar to Majorana neutrino via sphalerons

- Sphaleron Op. rewrite



## B-L Triangle:



The sphaleron conserves  $B - L$ :  
 IF  $\Delta L_{0\nu 2\beta} = 2 \Rightarrow \Delta B !!!$

Observe  $p$ -decay and  $n\bar{n} \rightarrow$  Neutrino Majorana



# $n \rightarrow \bar{n}$ Oscillation

Unkown mixing interaction

- Schrödinger equation

$$i\hbar \frac{\partial}{\partial t} \begin{pmatrix} \psi_n \\ \psi_{\bar{n}} \end{pmatrix} = \begin{pmatrix} E_n & \varepsilon \\ \varepsilon & E_{\bar{n}} \end{pmatrix} \begin{pmatrix} \psi_n \\ \psi_{\bar{n}} \end{pmatrix}, \quad \begin{pmatrix} \psi_n(0) \\ \psi_{\bar{n}}(0) \end{pmatrix} = \begin{pmatrix} 1 \\ 0 \end{pmatrix}$$

$$|\psi_{\bar{n}}(t)|^2 = \frac{4\varepsilon^2}{\omega^2 + 4\varepsilon^2} \sin^2\left(\frac{1}{2}\sqrt{\omega^2 + 4\varepsilon^2}t/\hbar\right)$$

$$\omega = (E_n - E_{\bar{n}}) = \left(m_n + \frac{p^2}{2m_n} + V_n\right) - \left(m_{\bar{n}} + \frac{p^2}{2m_{\bar{n}}} + V_{\bar{n}}\right)$$

- Transition Probability: (if  $\omega t \ll 1$ )

$$P_{n \rightarrow \bar{n}}(t) = |\psi_{\bar{n}}(t)|^2 = \varepsilon^2 \cdot (t/\hbar)^2 = \left(\frac{t}{\tau_{n\bar{n}}}\right)^2 \quad \tau_{n\bar{n}} = \frac{\hbar}{\varepsilon}$$

# Suppression of $n \rightarrow \bar{n}$ transition

$$|\psi_{\bar{n}}(t)|^2 = \frac{4\varepsilon^2}{\omega^2 + 4\varepsilon^2} \sin^2\left(\frac{1}{2}\sqrt{\omega^2 + 4\varepsilon^2}t/\hbar\right)$$

- Free neutron in a magnetic field

$$\omega = (E_n - E_{\bar{n}}) = \left(m_n + \frac{p^2}{2m_n} + V_n\right) - \left(m_{\bar{n}} + \frac{p^2}{2m_{\bar{n}}} + V_{\bar{n}}\right) = 2\mu \cdot B$$

- **Under earth field (0.5 gauss),  $2\mu HB = 6 \times 10^{-12} \text{ eV}$**

- $\varepsilon_{n\bar{n}} = \frac{\hbar}{\tau_{n\bar{n}}} < 10^{-23} \text{ eV}$  with  $\tau_{\text{free}} > 1.2 \times 10^8 \text{ s}$

$$|\psi_{\bar{n}}(t)|^2 = \frac{4\varepsilon^2}{\omega^2} \sin^2(t/\tau_{\text{Larmor}}) = 10^{-23} \sin^2(t/2 \times 10^{-4})$$

- To measure  $\tau_{\text{free}} > 1.2 \times 10^8 \text{ s}$ , the magnetic field has to be as small as  $0.5 \times 10^{-11} \text{ gauss!}$

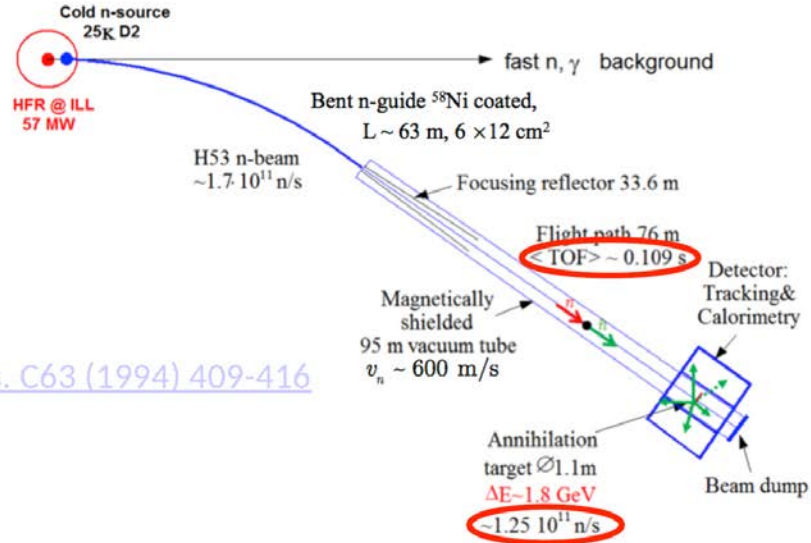
- For the neutron time-in-flight  $t = 0.1 \text{ s}$ ,

- $B < 5 \text{ mgauss.}$

$$\varepsilon_{n\bar{n}} = \frac{\hbar}{t} < 10^{-14} \text{ eV}$$



# Current Limits

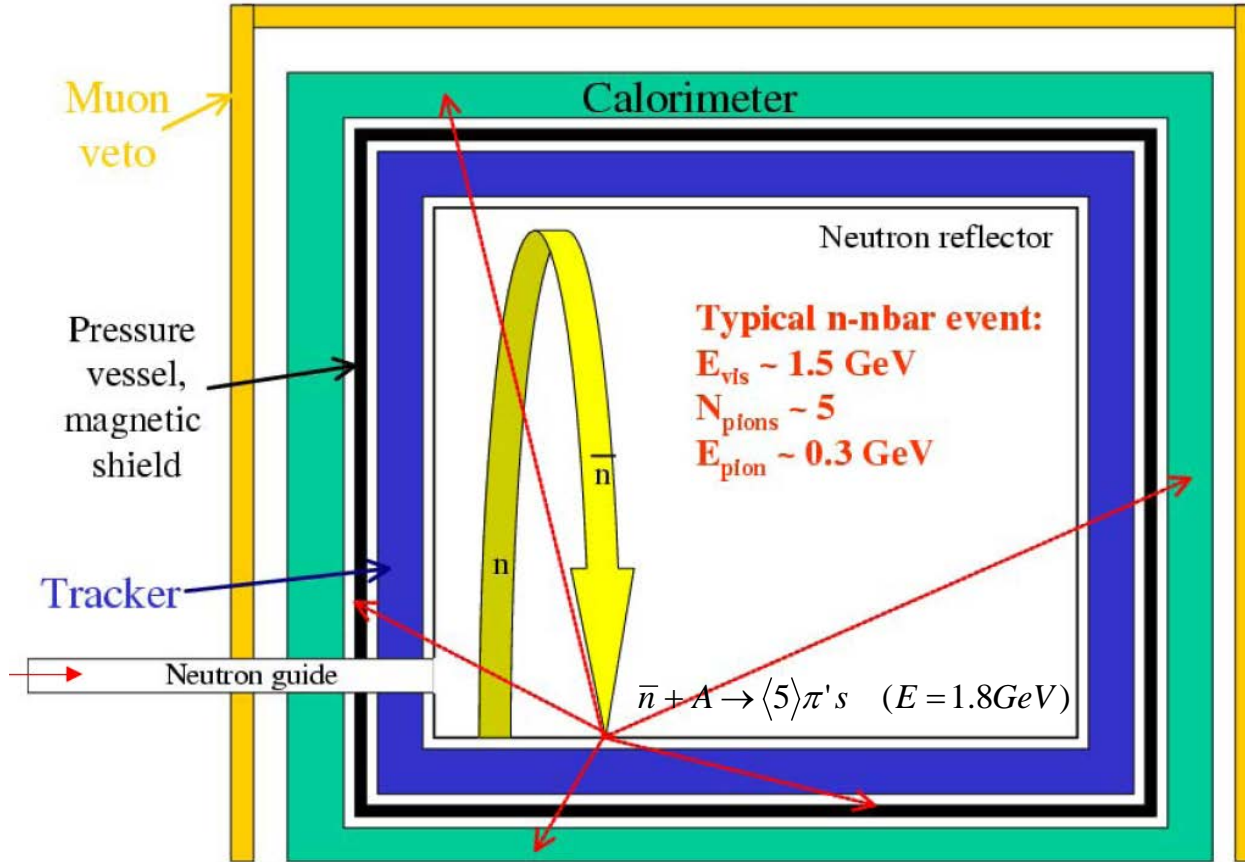


Baldo-Ceolin et al, [Z.Phys. C63 \(1994\) 409-416](#)

- $Nt^2 = 1.5 \cdot 10^9 \text{ s}^2$ ,  $P < 1.6 \cdot 10^{-18}$  (run lasted  $\sim 1$  year) and  $\tau > 0.86 \cdot 10^8 \text{ s}$
- Many subtle optimizations to minimize losses and backgrounds
- CN integrated beam flux was  $1.25 \times 10^{11} \text{ n/s}$
- Experiment was background-free
- Bound neutron limits  $\sim 3$  times better
- Many considerations make these measurements complementary to free neutron oscillations

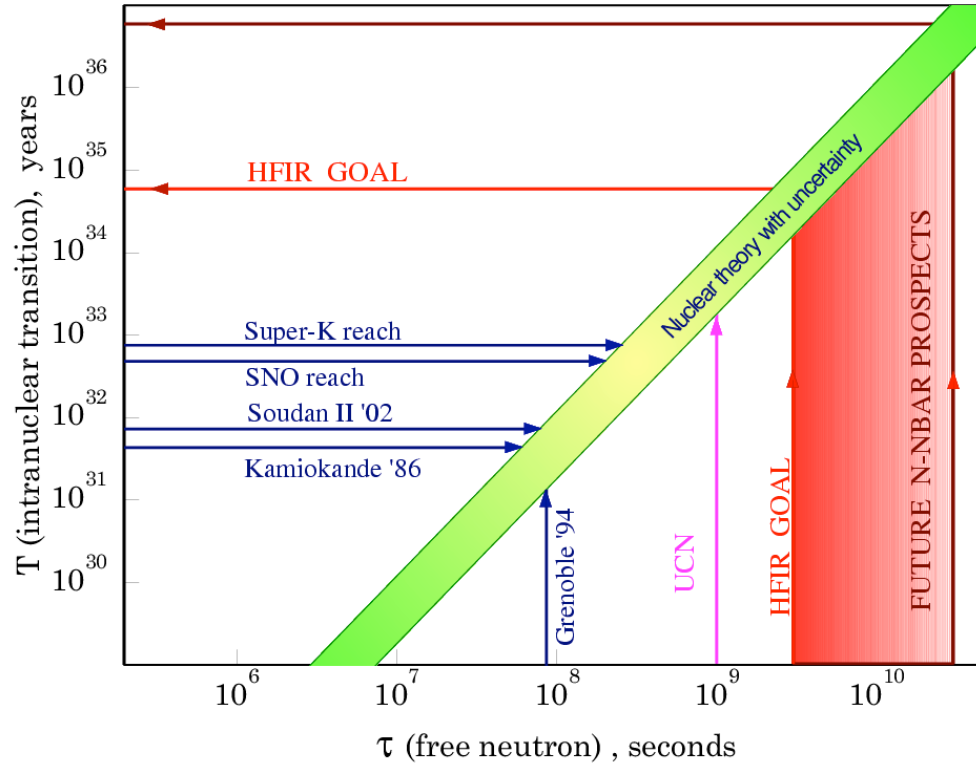
# Typical Detector for the UCN $N \rightarrow N\bar{n}$ Experiment

Typical detector size:  
height 2.5 m  
diameter 5 m



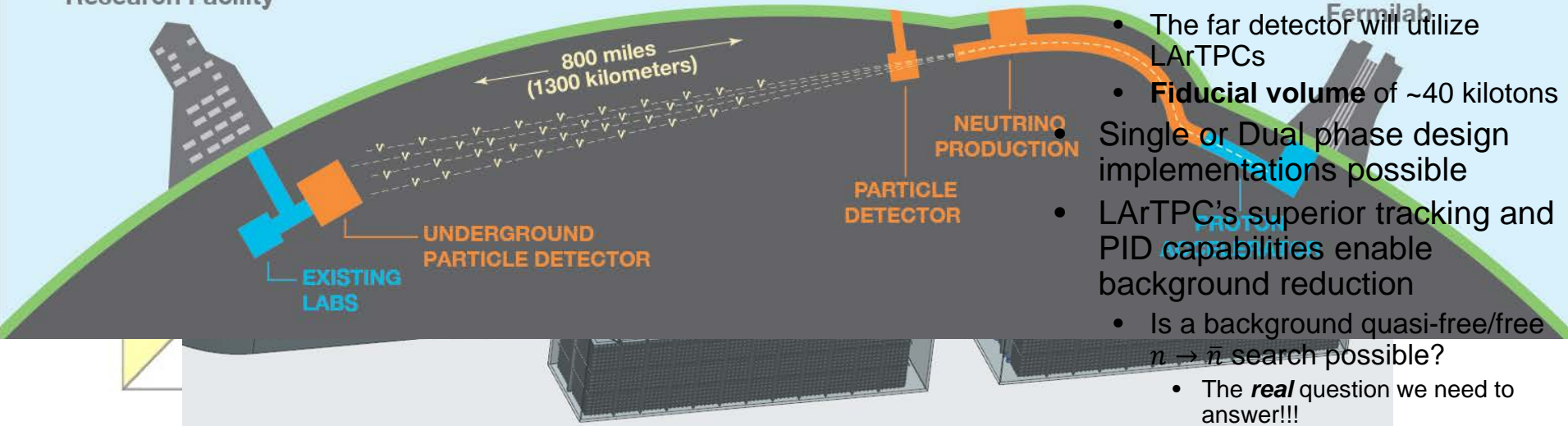
## Stability of matter from Neutron-Antineutron transition search

$T_A = R * (\tau_{\text{free}})^2$ , where R is "nuclear suppression factor" in intranuclear transition



# Bound Neutron Search at the Deep Underground Neutrino Experiment

Sanford Underground Research Facility



- **DUNE** international collaboration of 1000+
  - Partnership of Fermilab and LBNF
  - Will construct world's most intense  $\nu$  beam
  - The far detector will utilize LArTPCs
  - **Fiducial volume** of ~40 kilotons
- Single or Dual phase design implementations possible
- LArTPC's superior tracking and PID capabilities enable background reduction
  - Is a background quasi-free/free  $n \rightarrow \bar{n}$  search possible?
    - The **real** question we need to answer!!!

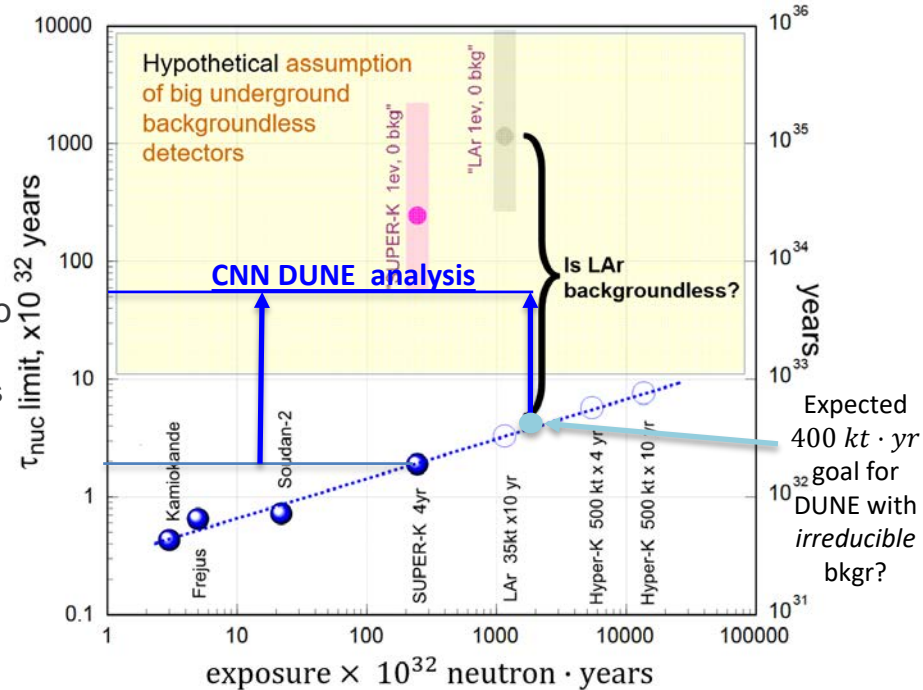
# Atmospheric $\nu$ Backgrounds Impede Intranuclear $n \rightarrow \bar{n}$ Event Identification In Large Underground Experiments

Previous searches for  $n \rightarrow \bar{n}$  suffered greatly from this!

- Super-Kamiokande: 24 candidate events
  - Expected *sig. eff: bkgrcount* = 12%: 24.1
- DUNE?
  - **Convolutional Neural Network**: *sig. eff: bkgrcount* = 14%: 3
    - Fully oscillated sample!
    - Important  $\nu_\tau$  interactions
  - **Truth**: *sig. eff: bkgrcount* = 27%: 0
    - Unoscillated sample! (a large grain of salt)

Theorists and computational physicists are working tirelessly to improve the accuracy of  $\nu$  generators

- This is a [requirement for understanding](#)  $\nu$  oscillation parameters precisely and atmospheric background properly
- [GENIE](#) is one of many used today
  - New cross section model from [Pastore](#) will be implemented soon
  - Implement explicit transformation radius dependence, spectral functions
- [GIBUU](#) and [NuWRO](#) getting great comparative results from [MicroBooNE](#) with their novel and unique techniques
- Plan to **run/re-run proper signal and background events** on all these platforms **for separability comparisons in the future**



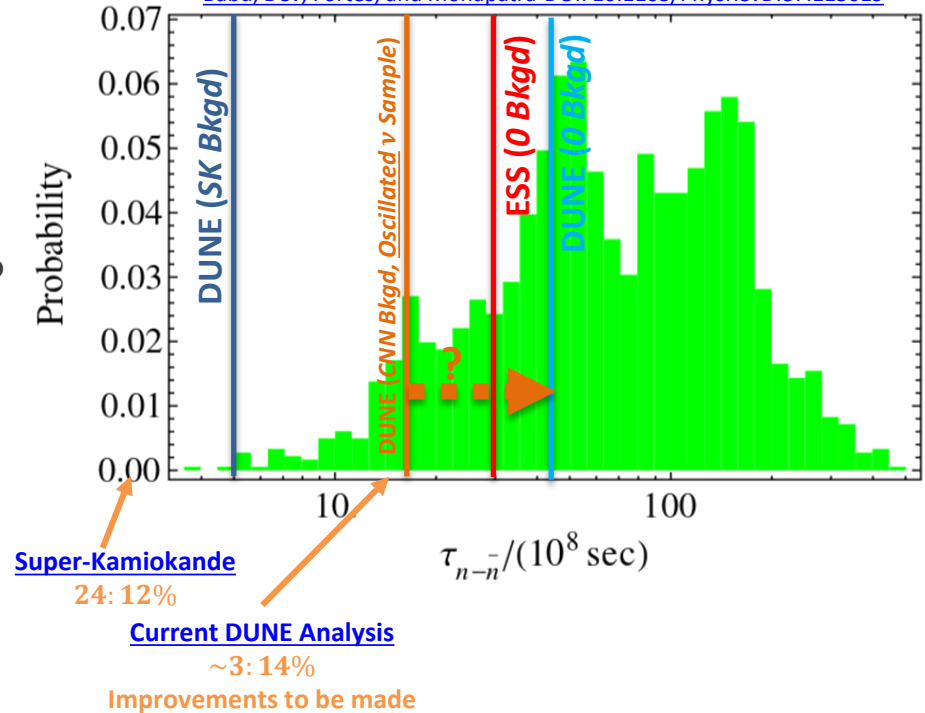
# Theoretically Important Probability Parameter Space of $\tau_{n \rightarrow \bar{n}}$

Post-sphaleron baryogenesis can predict the free  $n \rightarrow \bar{n}$  transformation time

- **Blue** shows converted limit from intranuclear transformation time
  - DUNE, 10 years,  $\sim 13,500\times$  ILL sensitivity
  - **Assumes 25% efficiency—more possible?**
  - **Assumes no background!**
- **Red** line shows free neutron transformation time
  - ESS, 3 yr, *goal* of  $\sim 1000\times$  ILL sensitivity
  - Assuming ILL-like *zero background*
    - Future work to show this definitively

$$\tau_M = R \cdot \tau_{n \rightarrow \bar{n}}^2$$

Babu, Dev, Fortes, and Mohapatra-DOI: [10.1103/PhysRevD.87.115019](https://doi.org/10.1103/PhysRevD.87.115019)



# Summary on $n\bar{n}$ experiments

- Baryon number violation is a requirement for the existence of our universe
- Arguably the best way to look for this is BSM processes such as  $n \rightarrow \bar{n}$  with pure  $\Delta B \neq 0$ 
  - Ability to say something experimentally about this depends on further integration of efforts between the neutron and HEP community
- DUNE has significant reach potential to constrain popular baryogenesis theories
  - Need to take into account zeroth—first-order corrections in nuclear physics models to understand signals and backgrounds properly
    - Transformation/annihilation radius distribution, spectral functions, new  $\nu$  cross section
- Free experiments (NNbar collaboration at ESS, see next talk from Albert) are also possible, and promise a similar reach
- Bound and free searches are ***incredibly complementary***, and, if  $n \rightarrow \bar{n}$  is definitively observed, rate differences could hint at further important BSM physics



# A New Limit on Time-Reversal-Invariance Violation in Beta Decay: Results of the emiT-II Experiment

T.E. Chupp, K.P. Coulter & R.L. Cooper  
*University of Michigan*



S.J. Freedman & B.K. Fujikawa  
*University of California - Berkeley/  
Lawrence Berkeley National Laboratory*



G.L. Jones  
*Hamilton College*

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A. Garcia  
*University of Washington*



H.P. Mumm, J.S. Nico, & A.K. Thompson  
*National Institute of Standards and Technology*

NIST



C. Trull & F.E. Wietfeldt  
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Tulane University

J.F. Wilkerson  
*University of North Carolina*



THE UNIVERSITY  
of NORTH CAROLINA  
at CHAPEL HILL

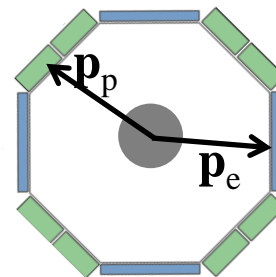
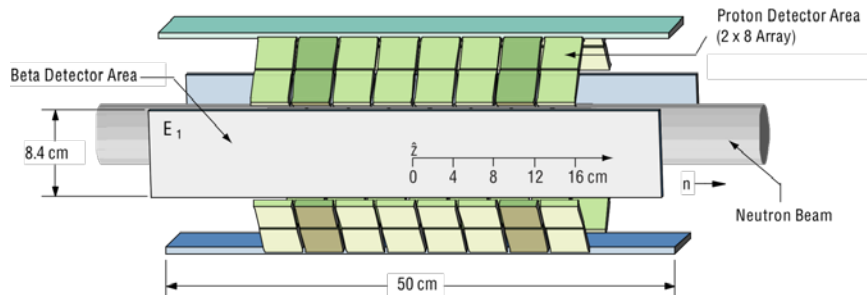
NEW RESULT:  $D = [-0.94 \pm 1.89(\text{STAT}) \pm 0.97(\text{SYS})] \times 10^{-4}$

Phys. Rev. C **86**, 035505

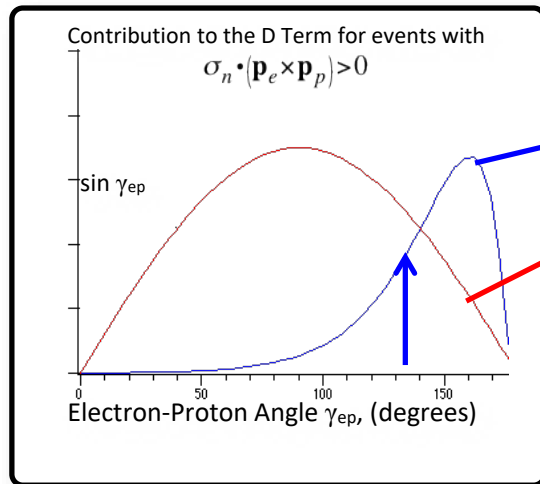
$$\varphi_{AV} = 180.012 \pm 0.028 \quad \frac{g_A}{g_V} = |\lambda| e^{i\varphi_{AV}}$$

# emiT: 8-fold symmetry

## 64 proton SBDs/4 $\beta$ scintillators



$$D\mathbf{J}_n \cdot (\hat{p}_p \times \hat{p}_e)$$

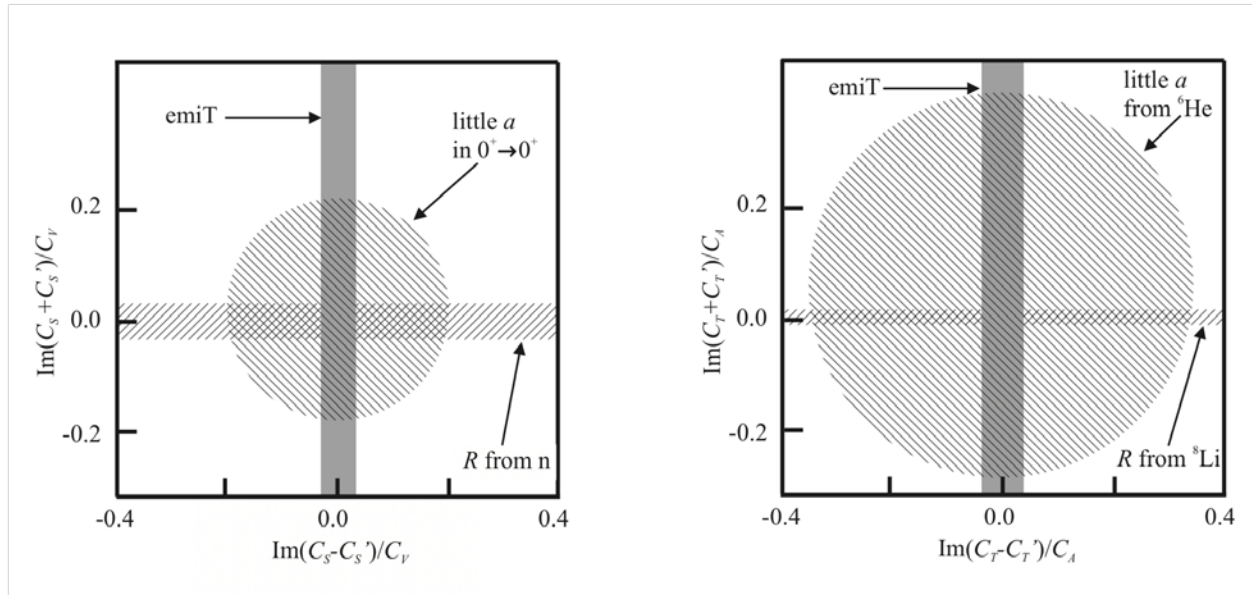


- Proton-electron momenta anticorrelated
- Coincidence rate favors  $180^\circ$ 
  - $\sin\theta_{ep}$  favors  $90^\circ$
  - FOM ( $1/\sigma^2$ ) 9x improved at  $135^\circ$
- Symmetrical, segmented detector:
  - Minimize sensitivity to  $A$  and  $B$
  - Investigate nonuniformities
  - Study systematic effects

# Final emiT Result

$$D = (-0.94 \pm 1.89 \text{ (stat)} \pm 0.97 \text{ (sys)}) \times 10^{-4}$$

$$\phi_{AV} = 180.012^\circ \pm 0.028^\circ$$



This is the most sensitive measurement of  $D$  in nuclear  $\beta$  decay. The result can be interpreted as a measurement of the phase of the ratio of the axial-vector and vector coupling constants ( $C_A/C_V = |\lambda|e^{i\phi_{AV}}$ ) with  $\phi_{AV} = 180.012^\circ \pm 0.028^\circ$  (68% confidence level). This result can also be used to constrain time-reversal-violating scalar and tensor interactions that arise in certain extensions to the Standard Model such as leptoquarks.

# QCD in a nutshell

Quantum Chromo Dynamics = gauge theory of the strong interactions

Fritzsch, Gell-Mann & Leutwyler

$f = u, d, s, c, b, t$

$SU_c(3)$  gauge group

$$\mathcal{L}_{QCD} = \sum_f \bar{q}_f (iD_\mu \gamma^\mu - m_f) q_f - \frac{1}{4} F_{\mu\nu}^a F^{a,\mu\nu}$$

$$q_f = \begin{pmatrix} q_{f,r} \\ q_{f,g} \\ q_{f,b} \end{pmatrix} \quad q'_{f,\alpha} = U_{\alpha\beta} q_{f,\beta}$$

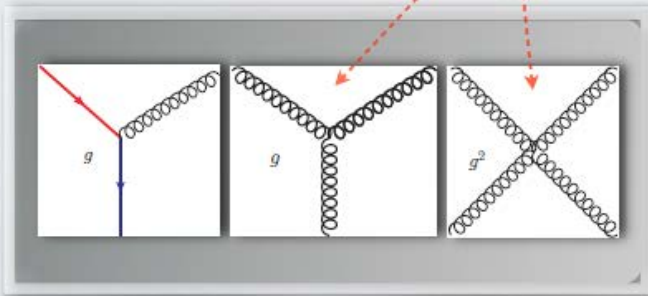
$$U \equiv \exp\left(-i \theta_a \frac{\lambda_a}{2}\right)$$

Eight non-commuting generators

$$\left[\frac{\lambda_a}{2}, \frac{\lambda_b}{2}\right] = i f^{abc} \frac{\lambda_c}{2}$$

$$D_\mu q_f \equiv (\partial_\mu + ig\mathcal{A}_\mu) q_f \quad \mathcal{A}_\mu = \mathcal{A}_\mu^a \frac{\lambda^a}{2}$$

$$F_{\mu\nu}^a = \partial_\mu \mathcal{A}_\nu^a - \partial_\nu \mathcal{A}_\mu^a - gf^{abc} \mathcal{A}_\mu^b \mathcal{A}_\nu^c$$



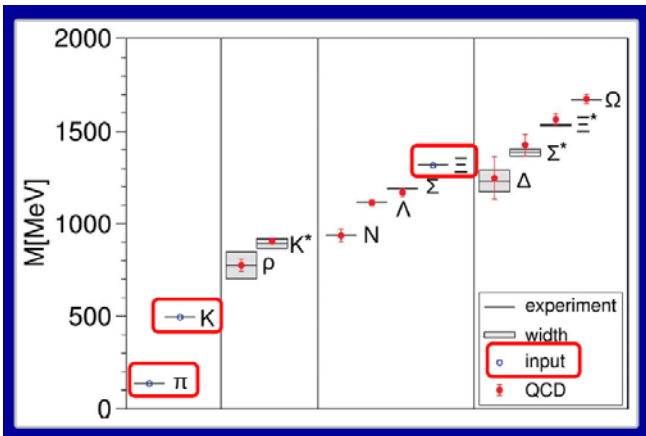
$$\mathcal{A}_\mu(x) \rightarrow \mathcal{A}_\mu(x) - \frac{1}{e} \partial_\mu \theta(x) \quad (\text{QED})$$

$$\mathcal{A}_\mu^a(x) \rightarrow \mathcal{A}_\mu^a(x) + \frac{1}{g} \partial_\mu \theta^a(x) + f^{abc} \mathcal{A}_\mu^c(x) \theta^b(x) \quad (\text{QCD})$$

Generators of the adjoint representation

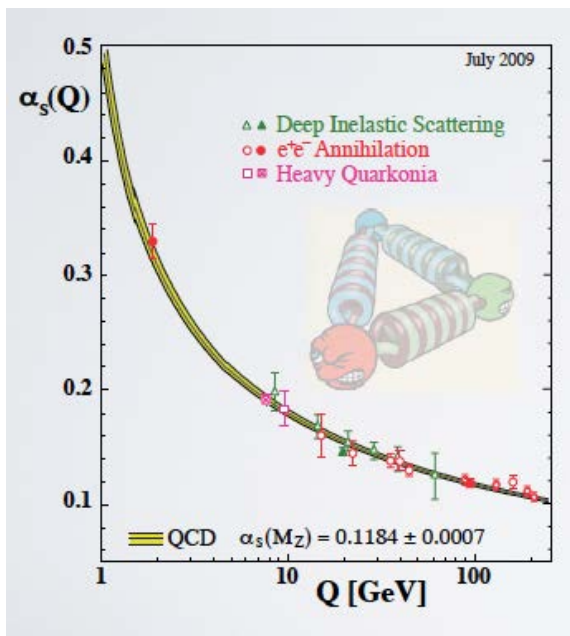
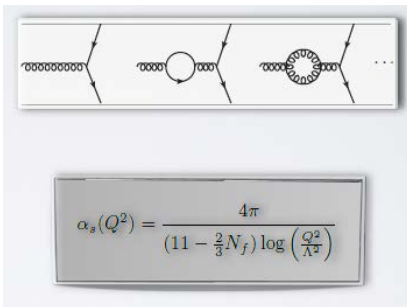
Gluons are in the adjoint representation

# Non-perturbative QCD Lattice QCD calculations



Supercomputers evaluate QCD correlation functions on a spacetime lattice.

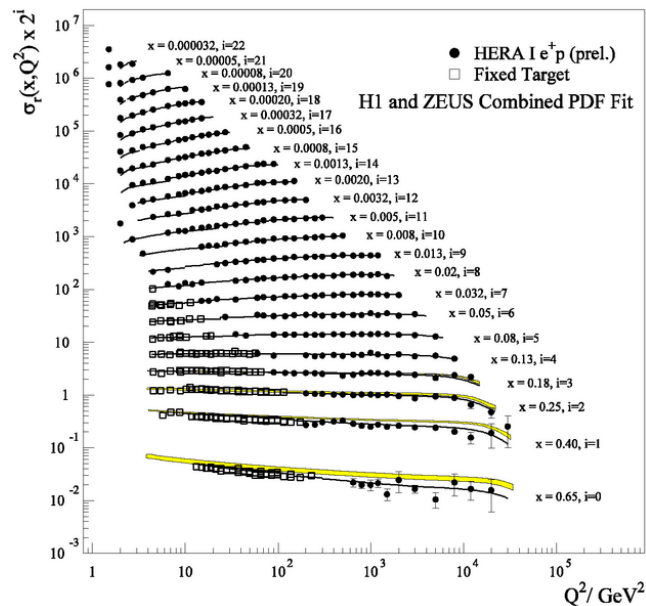
Gets the right hadron spectrum once a few masses are supplied



# Perturbative QCD

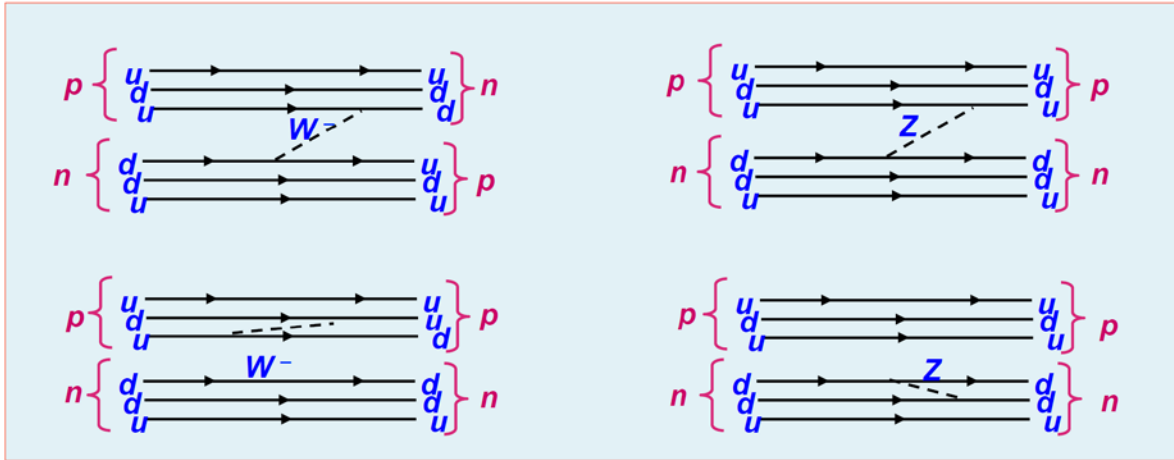
## Asymptotic Freedom

David J. Gross, H. David Politzer & Frank Wilczek, 2004



## NN Weak Interaction: use EW parity violation to probe QCD

In the Standard Model, the structure of the quark-quark weak interaction is known from the electroweak sector. However, strong QCD **confines color** and **breaks chiral symmetry**, thereby strongly correlating the quarks in both the *initial* and *final* nucleon ground states.



Two aspects of qq weak interaction make it useful as **an interesting probe of QCD**:

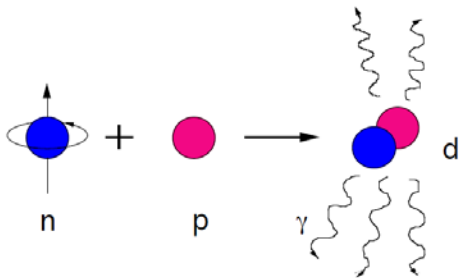
- (1) Since it is weak, it probes the nucleons in their ground states without exciting them.
- (2) Since it is short-ranged compared with the size of the nucleon, NN weak amplitudes should be first-order sensitive to **quark-quark correlation effects in the nucleon**.

QCD contains only vector quark-gluon couplings  $\rightarrow$  P is conserved.

Relative strength

$$\frac{e^2}{M_W^2} / \frac{g^2}{m_\pi^2} \approx 10^{-7}$$

# Few-Body P-odd NN in progress: n-p, n-<sup>3</sup>He, n-<sup>4</sup>He



NPDγ  
 $A_\gamma \sim 10^{-8}$

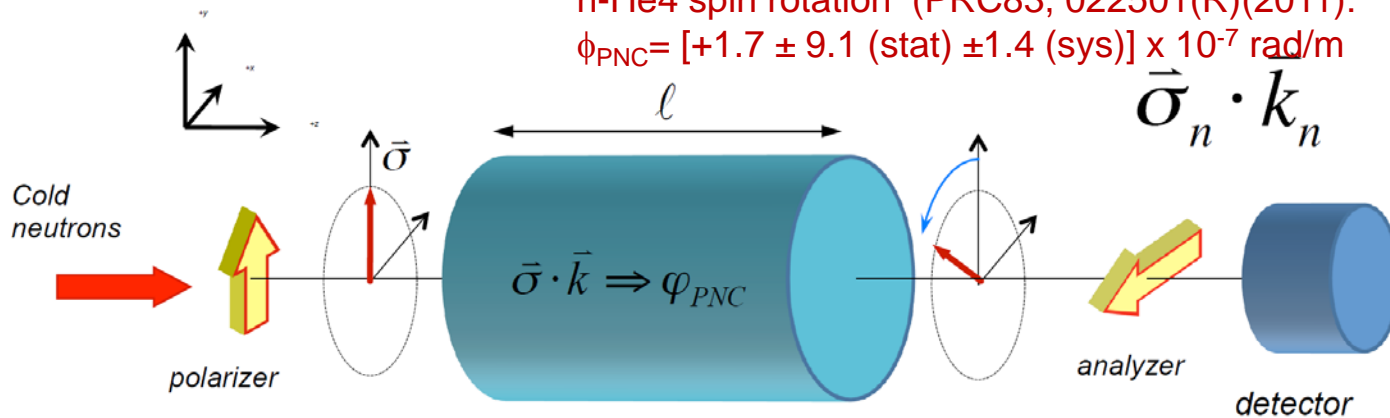
$$\vec{\sigma}_n \cdot \vec{k}_\gamma$$



N-He3  
 $A_p \sim 2 \times 10^{-8}$

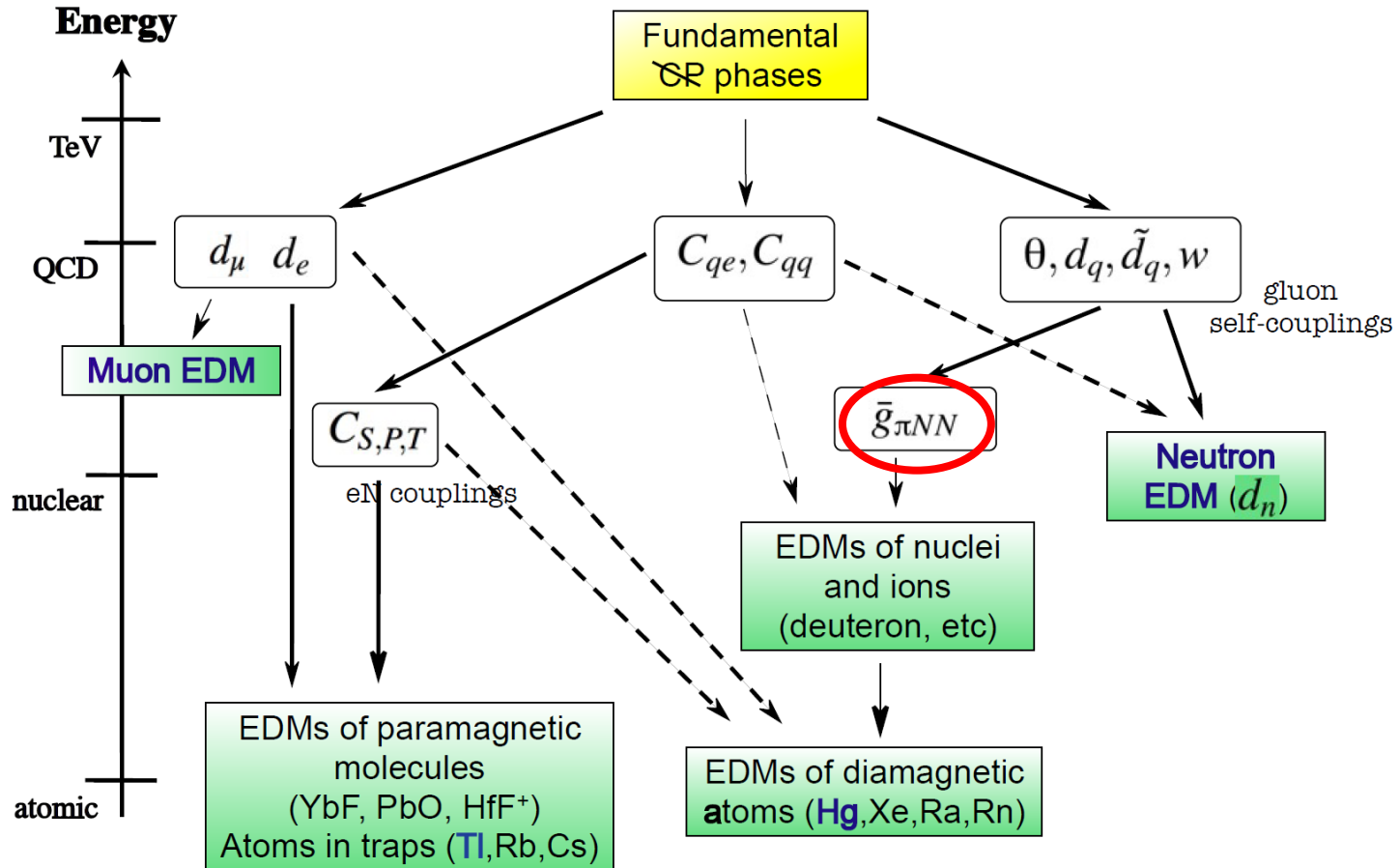
$$\vec{\sigma}_n \cdot \vec{k}_p$$

n-He4 spin rotation (PRC83, 022501(R)(2011):  
 $\phi_{\text{PNC}} = [+1.7 \pm 9.1 \text{ (stat)} \pm 1.4 \text{ (sys)}] \times 10^{-7} \text{ rad/m}$

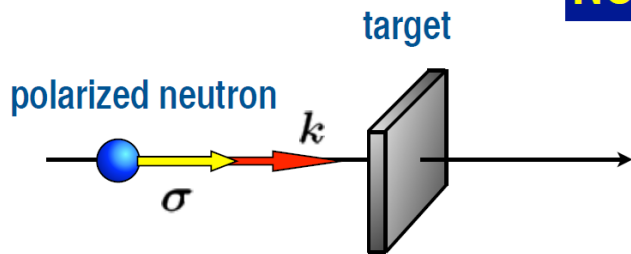




# CP-violation in Low Energy Phenomena



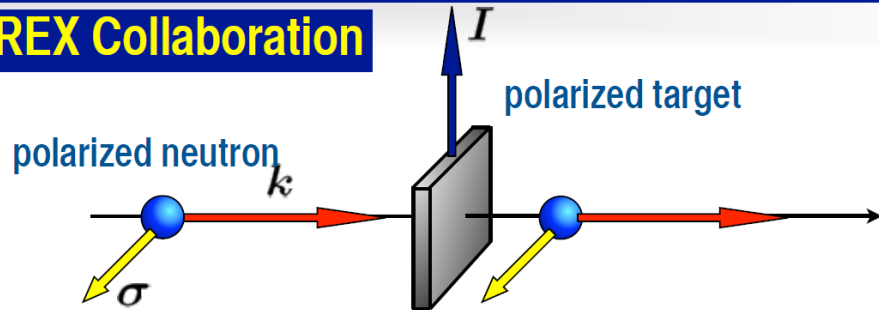
NOPTREX Collaboration



$$\sigma \cdot \hat{k}$$

$$P : \sigma \cdot \hat{k} \rightarrow \sigma \cdot (-\hat{k})$$

P-odd



$$\sigma \cdot (\hat{k} \times \hat{I})$$

$$P : \sigma \cdot (\hat{k} \times \hat{I}) \rightarrow \sigma \cdot ((-\hat{k}) \times \hat{I})$$

$$T : \sigma \cdot (\hat{k} \times \hat{I}) \rightarrow (-\sigma) \cdot ((-\hat{k}) \times (-\hat{I}))$$

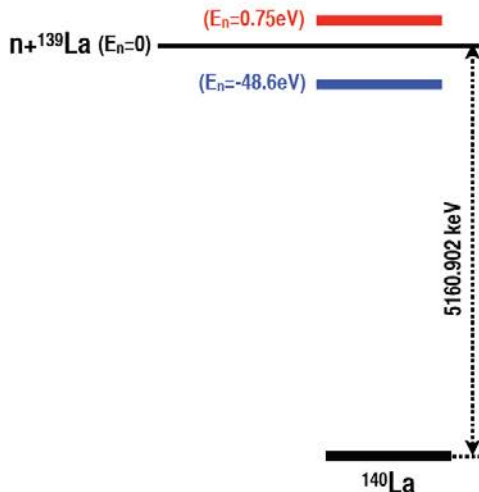
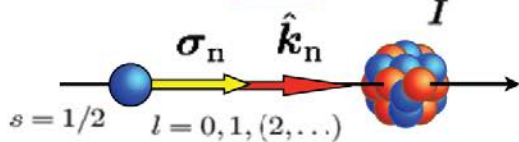
P-odd T-odd



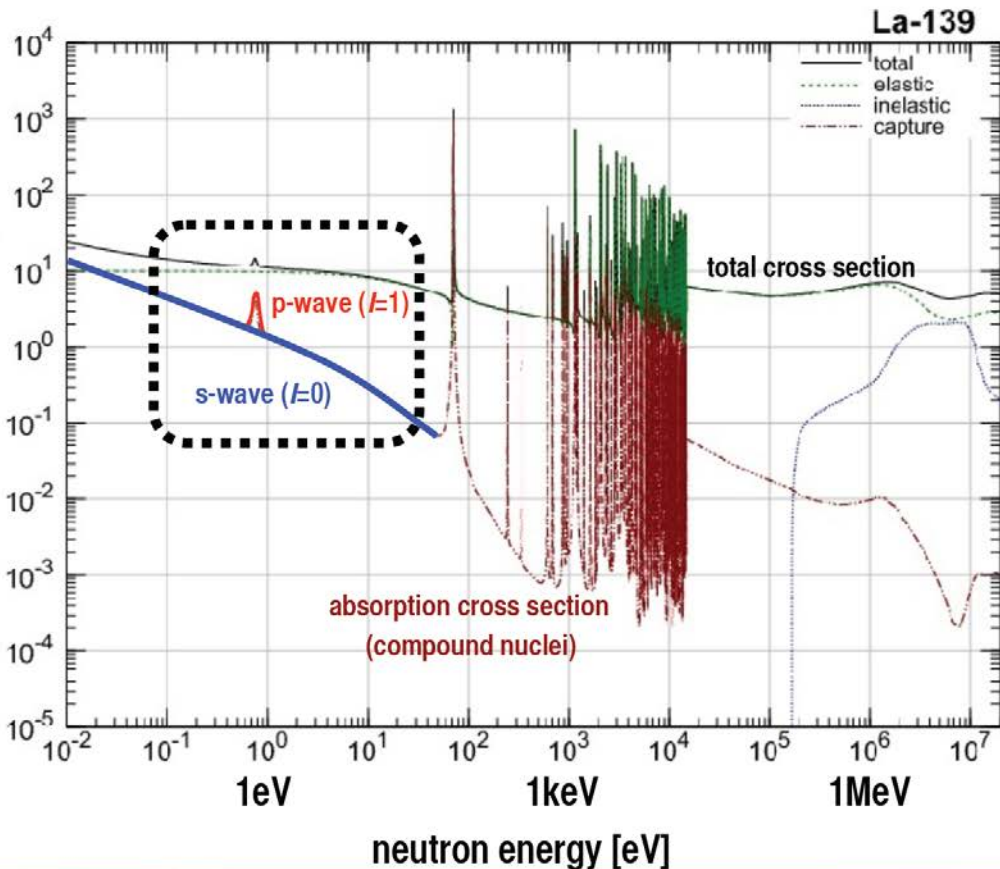
# Compound States

**P-violation**

$$\sigma = \sigma_0 + \Delta\sigma(\sigma_n \cdot \hat{k}_n)$$



cross section [b]



thermal

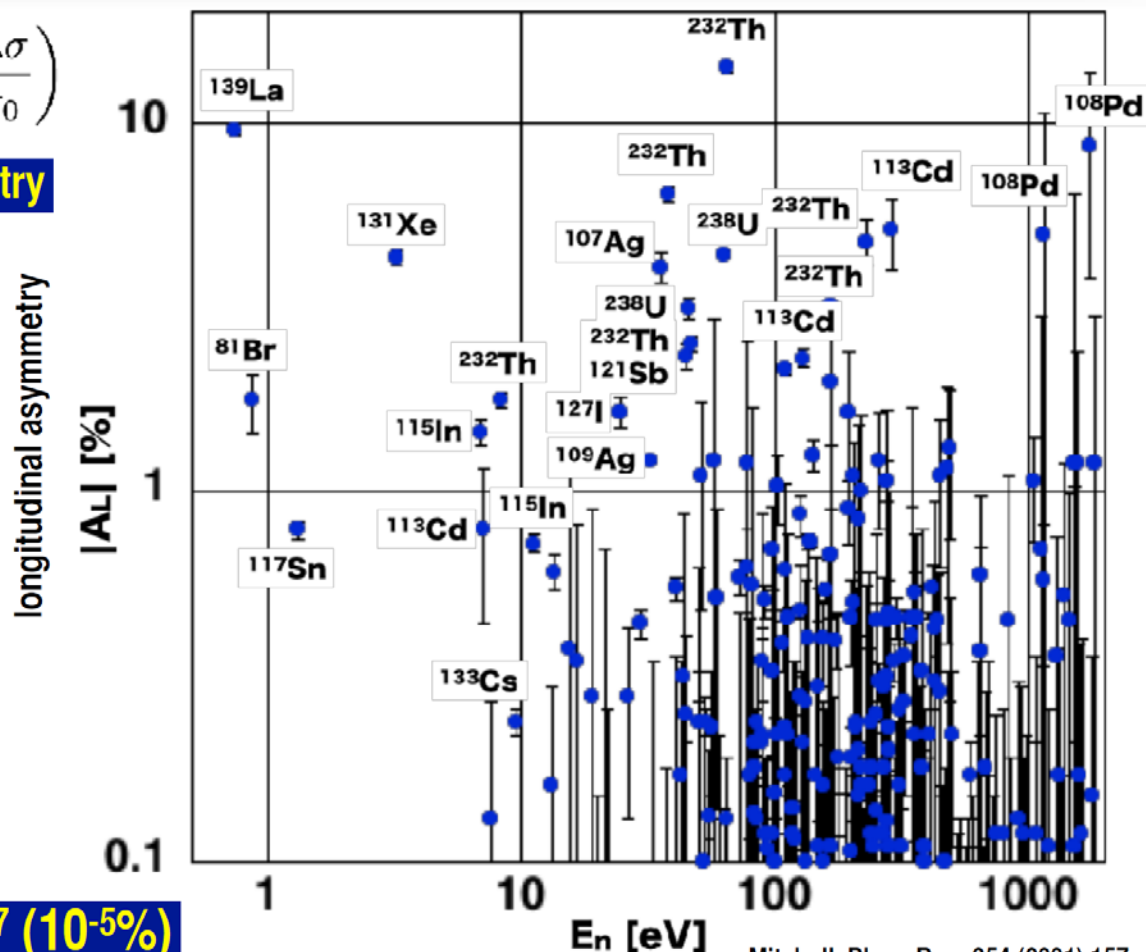
epithermal

fast

# Enhanced P-violation in Compound States

$$A_L = \frac{\sigma_+ - \sigma_-}{\sigma_+ + \sigma_-} \left( = \frac{\Delta\sigma}{\sigma_0} \right)$$

**Longitudinal Asymmetry**



**NN-interaction  $10^{-7}$  ( $10^{-5}\%$ )**

# T-violation in Neutron Optics

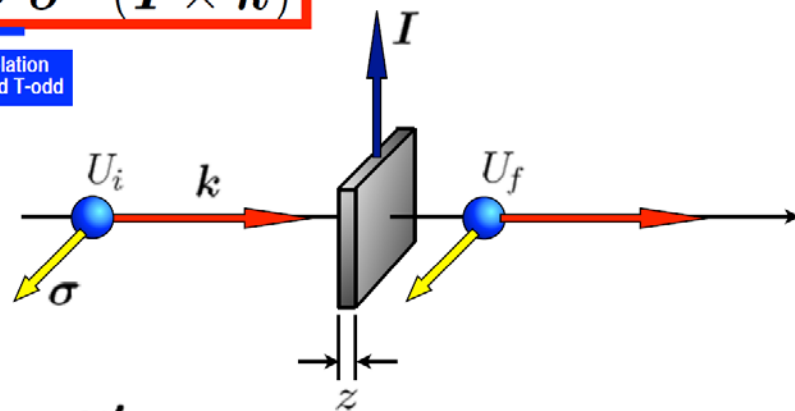
$$f = \underbrace{A'}_{\text{Spin Independent}} + \underbrace{B' \sigma \cdot \hat{I}}_{\text{Spin Dependent}} + \underbrace{C' \sigma \cdot \hat{k}}_{\text{P-violation}} + \underbrace{D' \sigma \cdot (\hat{I} \times \hat{k})}_{\text{T-violation}}$$

Spin Independent  
P-even T-even

Spin Dependent  
P-even T-even

P-violation  
P-odd T-even

T-violation  
P-odd T-odd



T-violating matrix element

$$D' \rightarrow \Delta\sigma_{CP} = \kappa(J) \frac{W_T}{W} \Delta\sigma_P \leftarrow C'$$

Gudkov, Phys. Rep. 212 (1992) 77

T-violation

angular  
momentum  
factor

P-violation

P-violating matrix element

$$\kappa(J) = 0.99_{-0.07}^{+0.88}, 4.84_{-1.69}^{+5.58}$$

$$\left| \frac{W_T}{W} \right| < 3.9 \times 10^{-4}$$

# T-violation in Neutron Optics

$$f = \underbrace{A'}_{\text{Spin Independent}} + \underbrace{B'\sigma \cdot \hat{I}}_{\text{Spin Dependent}} + \underbrace{C'\sigma \cdot \hat{k}}_{\text{P-violation}} + \underbrace{D'\sigma \cdot (\hat{I} \times \hat{k})}_{\text{T-violation}}$$

Spin Independent  
P-even T-even

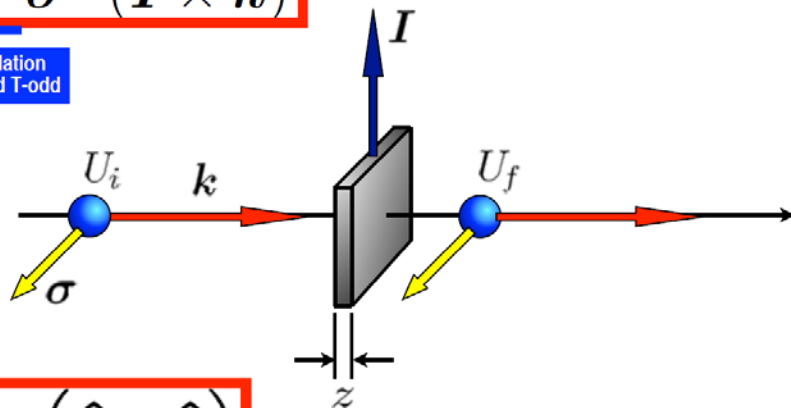
Spin Dependent  
P-even T-even

P-violation  
P-odd T-even

T-violation  
P-odd T-odd

$$U_f = \delta U_i$$

$$\delta = e^{i(n-1)kz} \quad n = 1 + \frac{2\pi\rho}{k^2} f$$



$$\delta = \underbrace{A}_{\text{Spin Independent}} + \underbrace{B\sigma \cdot \hat{I}}_{\text{Spin Dependent}} + \underbrace{C\sigma \cdot \hat{k}}_{\text{P-violation}} + \underbrace{D\sigma \cdot (\hat{I} \times \hat{k})}_{\text{T-violation}}$$

Spin Independent  
P-even T-even

Spin Dependent  
P-even T-even

P-violation  
P-odd T-even

T-violation  
P-odd T-odd

$$A = e^{iZA'} \cos b$$

$$B = ie^{iZA'} \frac{\sin b}{b} ZB'$$

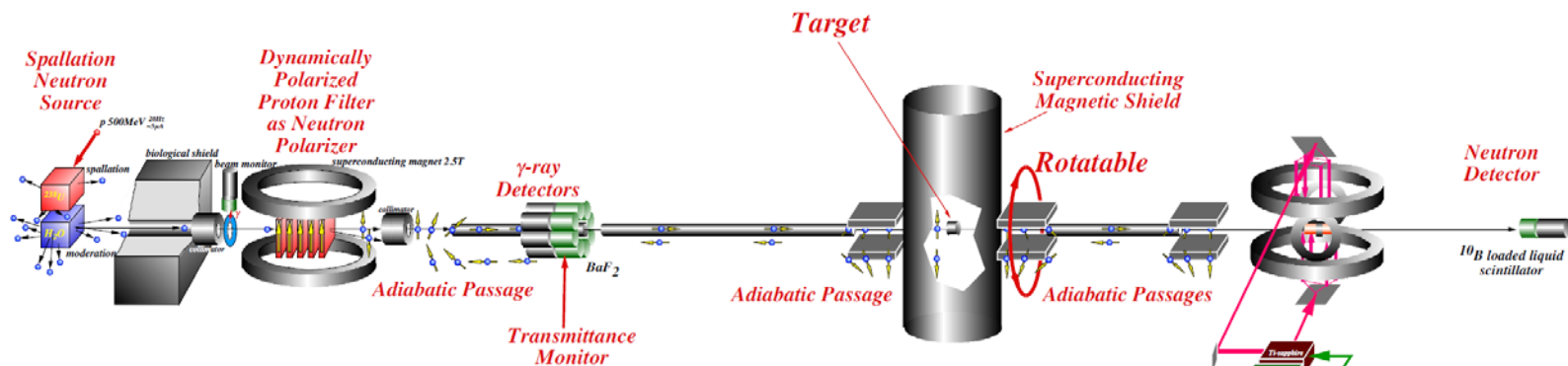
$$Z = \frac{2\pi\rho}{k} z$$

$$C = ie^{iZA'} \frac{\sin b}{b} ZC'$$

$$b = Z(B'^2 + C'^2 + D'^2)^{1/2}$$

$$D = ie^{iZA'} \frac{\sin b}{b} ZD'$$

**D ≠ 0 ⇒ D' ≠ 0**

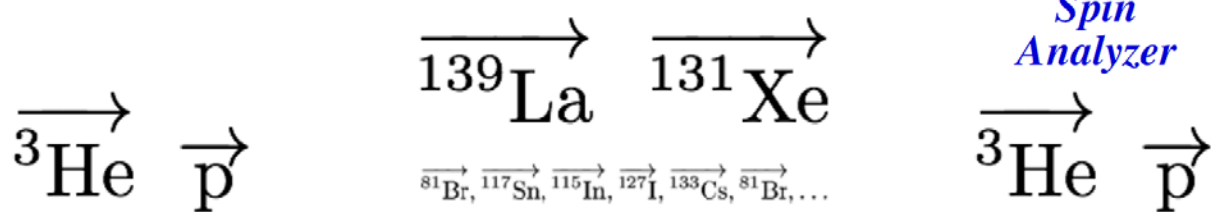


**epithermal  
neutron  
polarizer**

**polarized target  
spin control**

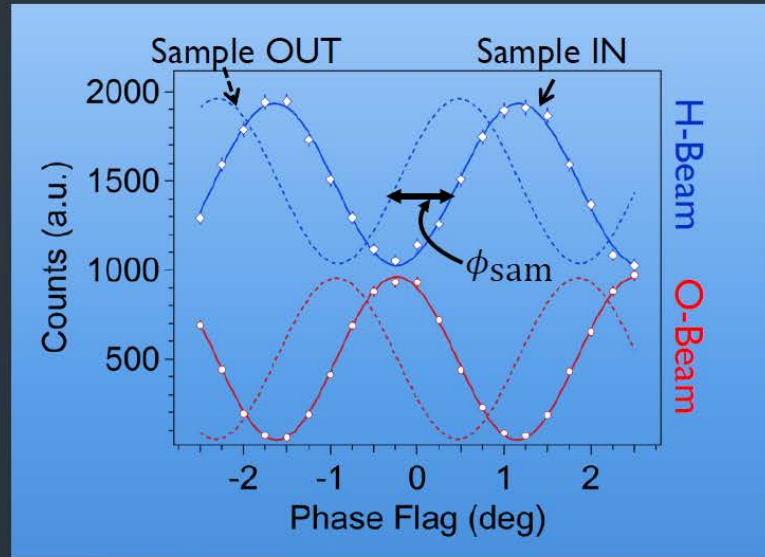
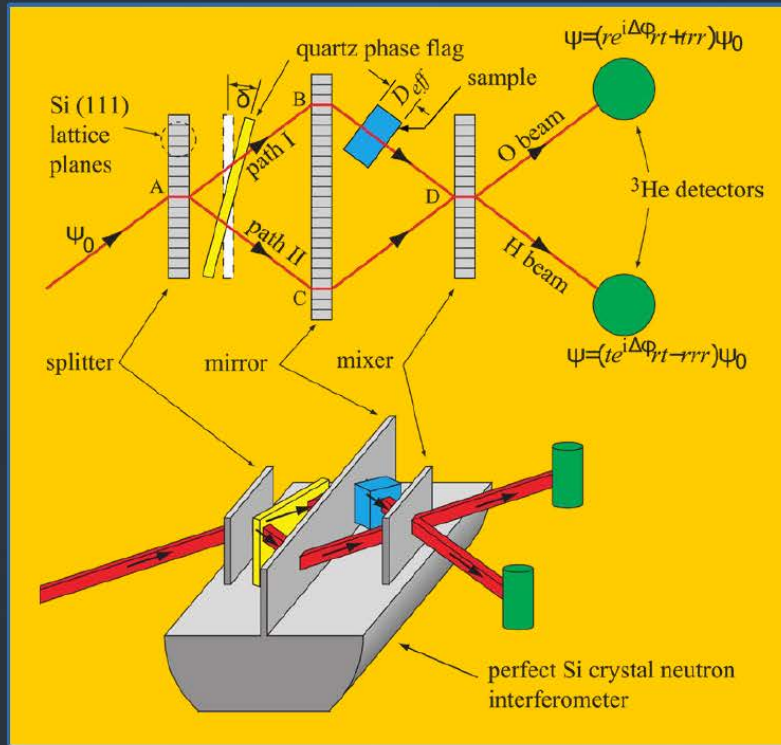
**epithermal  
neutron  
spin analyzer**

← Polarized Neutron Source → Target station → Neutron Spin Analyzer →





# Neutron Interferometer



Interferogram:

$$I_O = A \left[ 1 + C * \cos(\phi_{\text{sam},1} + \phi_{\text{sam},2}(d_{\text{eff}}) + \phi_0) \right]$$

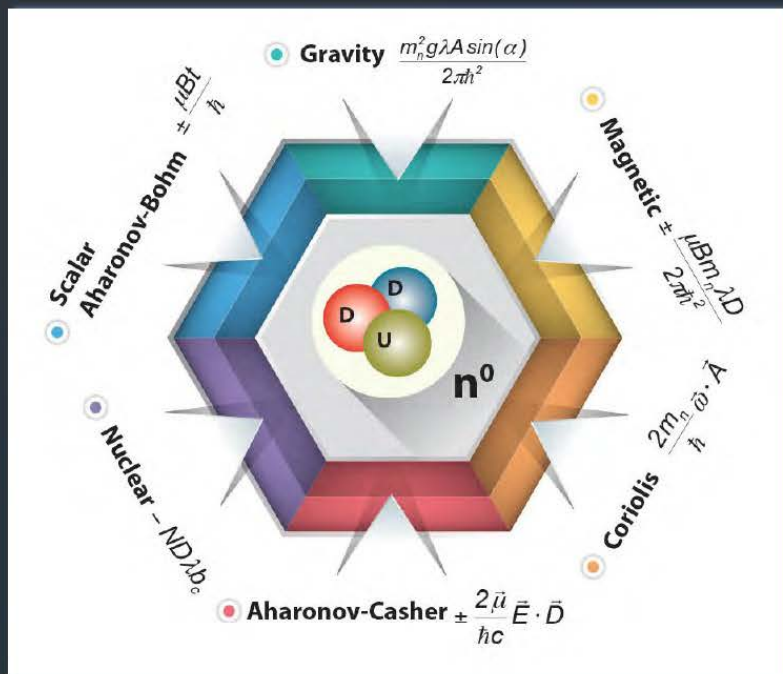
$$I_H = A \left[ \frac{B}{A} - C * \cos(\phi_{\text{sam},1} + \phi_{\text{sam},2}(d_{\text{eff}}) + \phi_0) \right]$$

The fringe visibility or contrast ( $C$ ) is an important parameter as  $\delta\phi \propto C^{-1}$  and is used to evaluate an interferometer's quality.

# Phase Shifts



neutron interferometry  
is a *diverse* instrument!



**Nuclear:  $\phi_{nuc}$**

$$\phi_{nuc} \sim 135\pi \text{ for 1 cm of Al}$$

**Gravity:  $\phi_{gravity}$**

$$\phi_{gravity} \sim 50\pi \text{ for } 0.002 \text{ m}^2 \text{ Area}$$

**Magnetic:  $\phi_{mag}$**

$$\phi_{mag} \sim 15\pi \text{ for 1 cm of 100 G field}$$

**Aharonov-Casher:  $\phi_{AC}$**

$$\phi_{AC} \sim 1 \text{ mrad for } 30\text{kV field}$$

**Geometric:  $\phi_{geo}$**

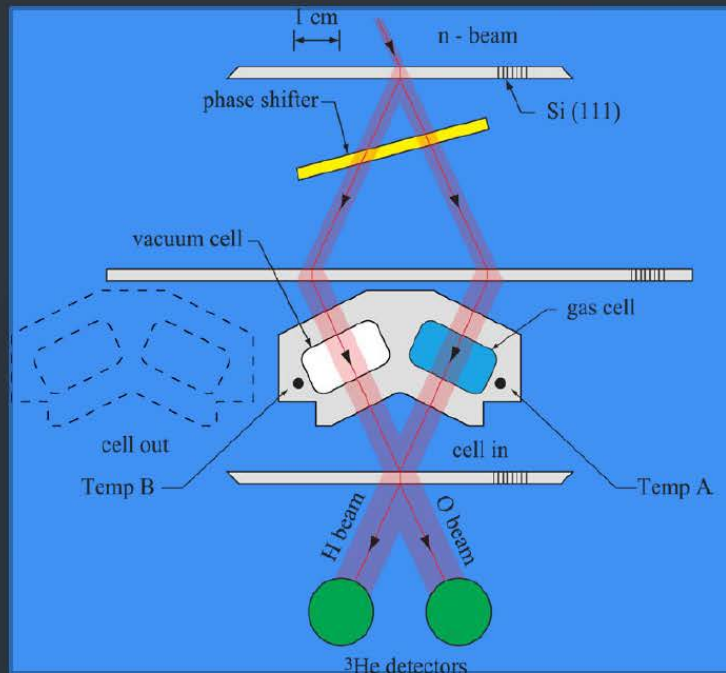
$$\phi_{geo} \sim \pi$$

**Sagnac:  $\phi_{sag}$**

$$\phi_{sag} \sim \pi \text{ for } 0.002 \text{ m}^2 \text{ Area}$$

# Precision Scattering Lengths

- Tighter constraints for NN and few-nucleon potential models
- Neutron scattering lengths of light nuclei are benchmarks for chiral effective field theories and can be used to calculate low-energy coefficients required by the theory
- $^4\text{He}$  is often utilized in fundamental neutron experiments exploring physics beyond the standard model



$$b = \frac{\phi_{gas}}{N(T, P)\lambda D(T)}$$

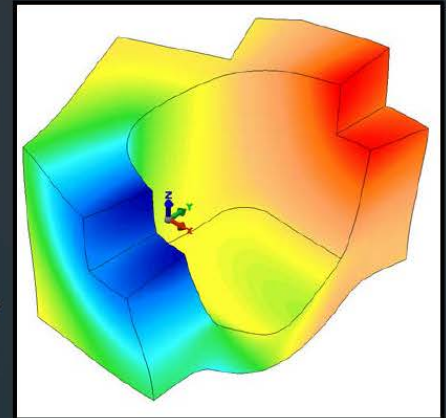
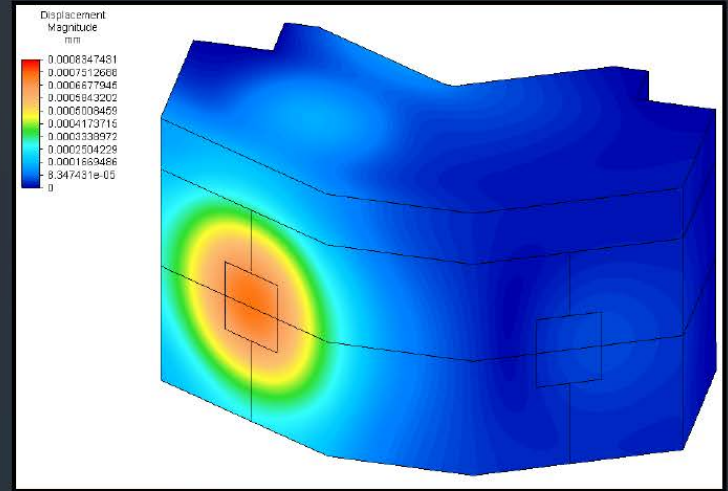
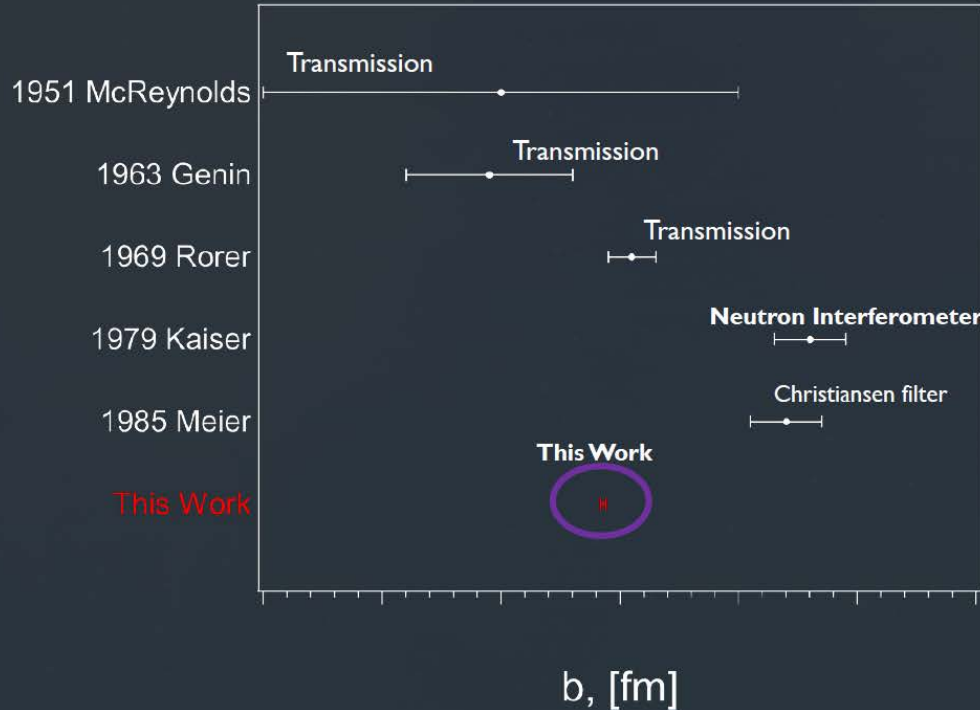
Path Length  $D(T) = D_0[1 + \alpha(T - T_0)]$

Density  $N(T, P) = \frac{P}{k_B T(1 + B_P P + C_P P^2)}$

Wavelength  $\lambda = 2.70913(15) \text{ \AA}$



# n-<sup>4</sup>He Preliminary Results



FEA calculations of the effects of cell deformation when pressurized will shift this result  $\sim 1\sigma$  are pending...

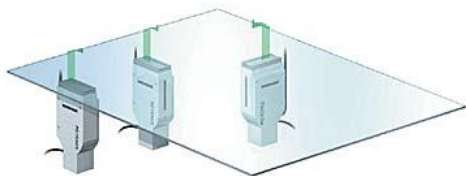
10x more precise result and shifts the world average



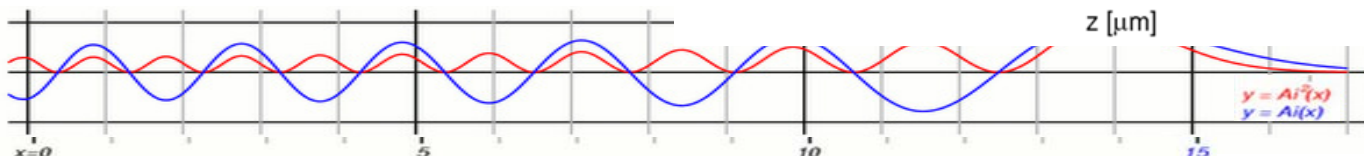
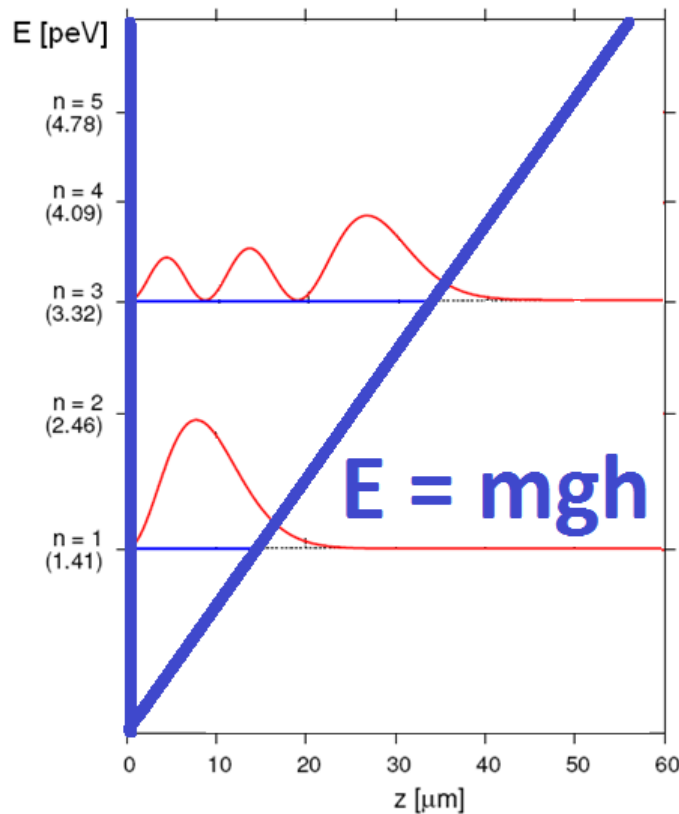
# Quantum States in the Gravity Potential

- Bound States
- Discrete energy levels
- Ground state 1.4 peV
- Airy-Functions

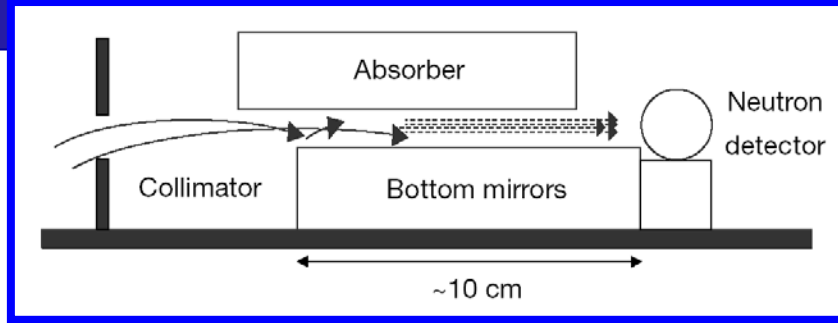
$$\left( -\frac{\hbar^2}{2m} \frac{\partial^2}{\partial z^2} + mgz \right) \varphi_n(z) = E_n \varphi_n(z)$$



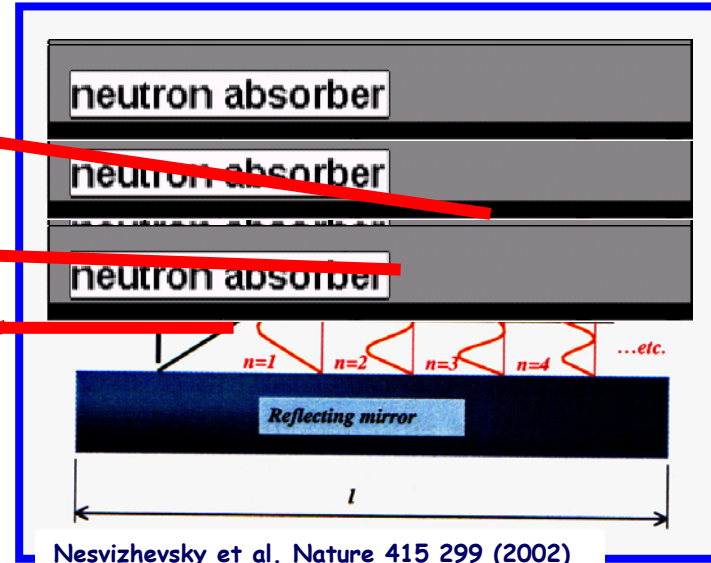
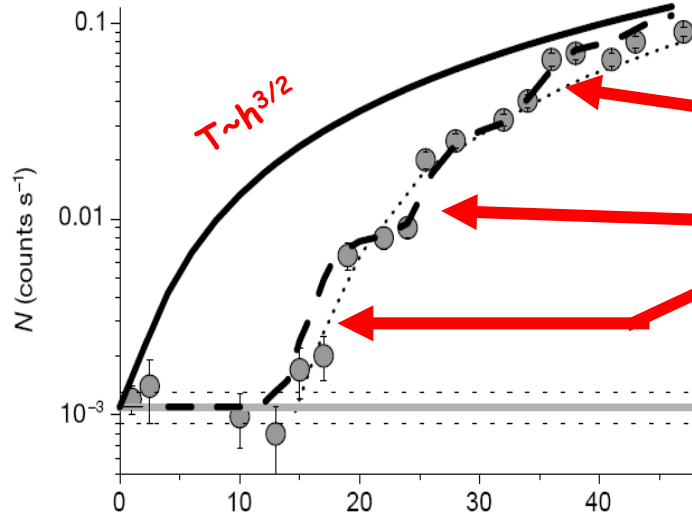
$\psi^2$



# Nesvizhevsky et al. 2002: Observation of Bound Quantum States



Neutron mirror:  
polished glass plate 10 cm long



# Neutrons test Newton

$$V(r) = G \frac{m_1 \cdot m_2}{r} (1 + \alpha \cdot e^{-r/\lambda})$$

● Strength  $\alpha$

● Range  $\lambda$

## Hypothetical Gravity Like Forces

### Extra Dimensions:

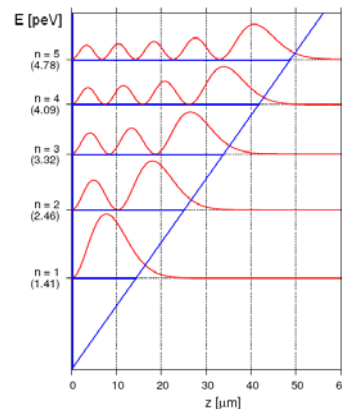
The string and  $D_p$ -brane theories predict the existence of extra space-time dimensions

**Infinite-Volume Extra Dimensions:** Randall and Sundrum

**Exchange Forces** from new Bosons: a deviation from the ISL can be induced by the exchange of new (pseudo)scalar and (pseudo)vector bosons

- **Axion** - - - - -  $\rightarrow 0.2 \mu\text{m} < \lambda < 0.2 \text{cm}$
- **Scalar boson. Cosmological consideration**
- **Bosons from Hidden Supersymmetric Sectors**
- **Gauge fields in the bulk (ADD, PRD 1999)** - - -  $\rightarrow 10^6 < \alpha < 10^9$

**Supersymmetric large Extra Dimensions (B.& C.)** - - -  $\rightarrow \alpha < 10^6$





# Short range fundamental forces

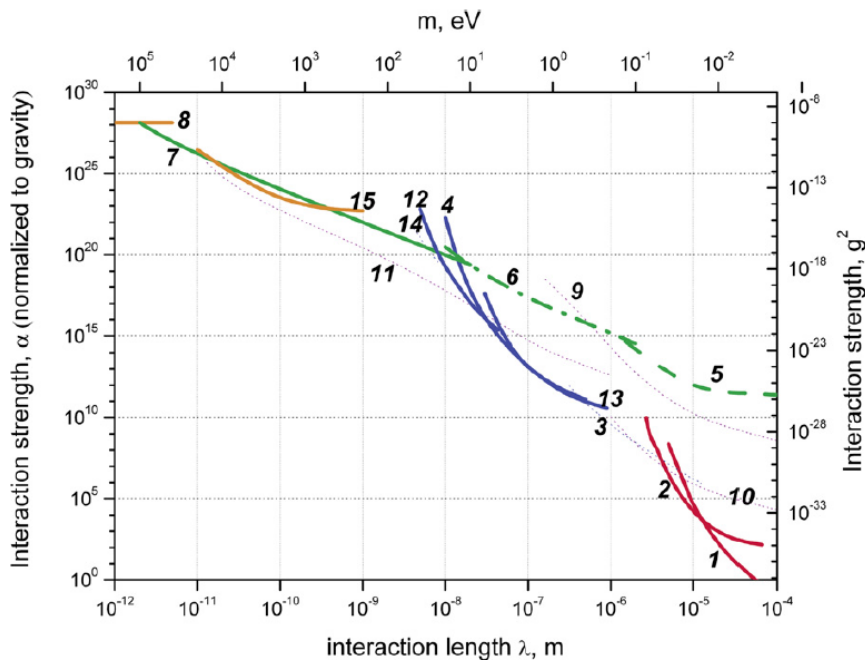
Ultra cold neutron quantum states / États quantiques des neutrons ultra froids

Short-range fundamental forces

Forces fondamentales à courte portée

I. Antoniadis<sup>a</sup>, S. Baessler<sup>b,c</sup>, M. Büchner<sup>d</sup>, V.V. Fedorov<sup>e</sup>, S. Hoedl<sup>f</sup>, A. Lambrecht<sup>g</sup>,  
V.V. Nesvizhevsky<sup>g,\*</sup>, G. Pignol<sup>h</sup>, K.V. Protasov<sup>h</sup>, S. Reynaud<sup>i</sup>, Yu. Sobolev<sup>j</sup>

*I. Antoniadis et al. / C. R. Physique 12 (2011) 755–778*



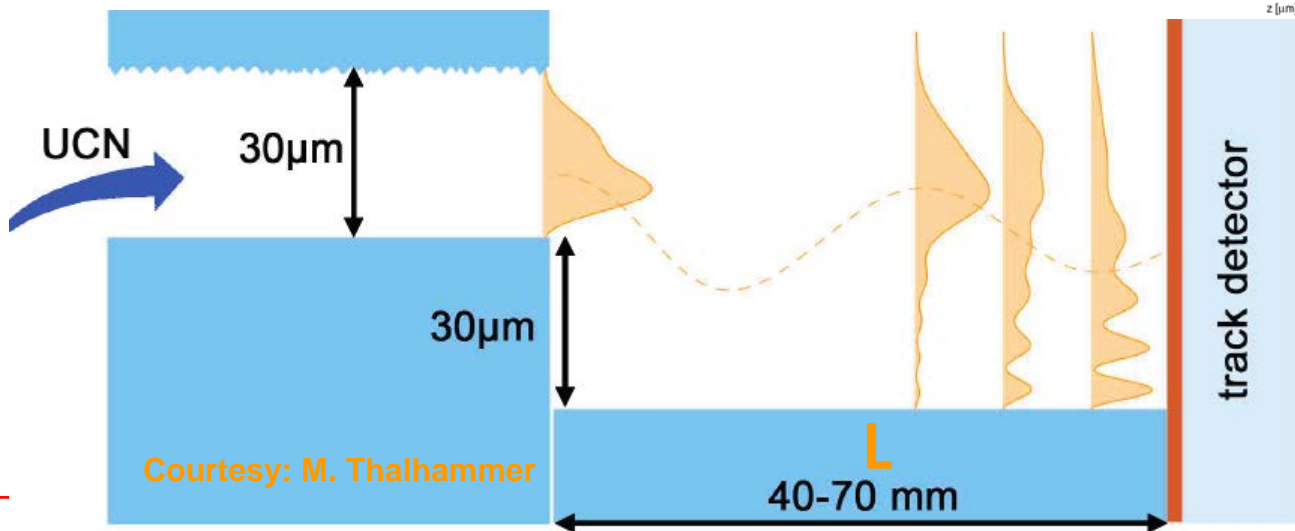
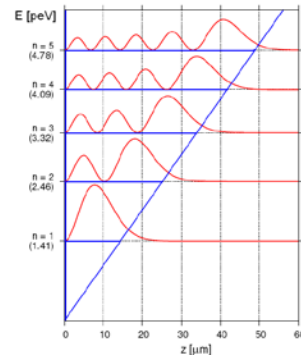
$$V(r) = G \frac{m_1 \cdot m_2}{r} (1 + \alpha \cdot e^{-r/\lambda})$$

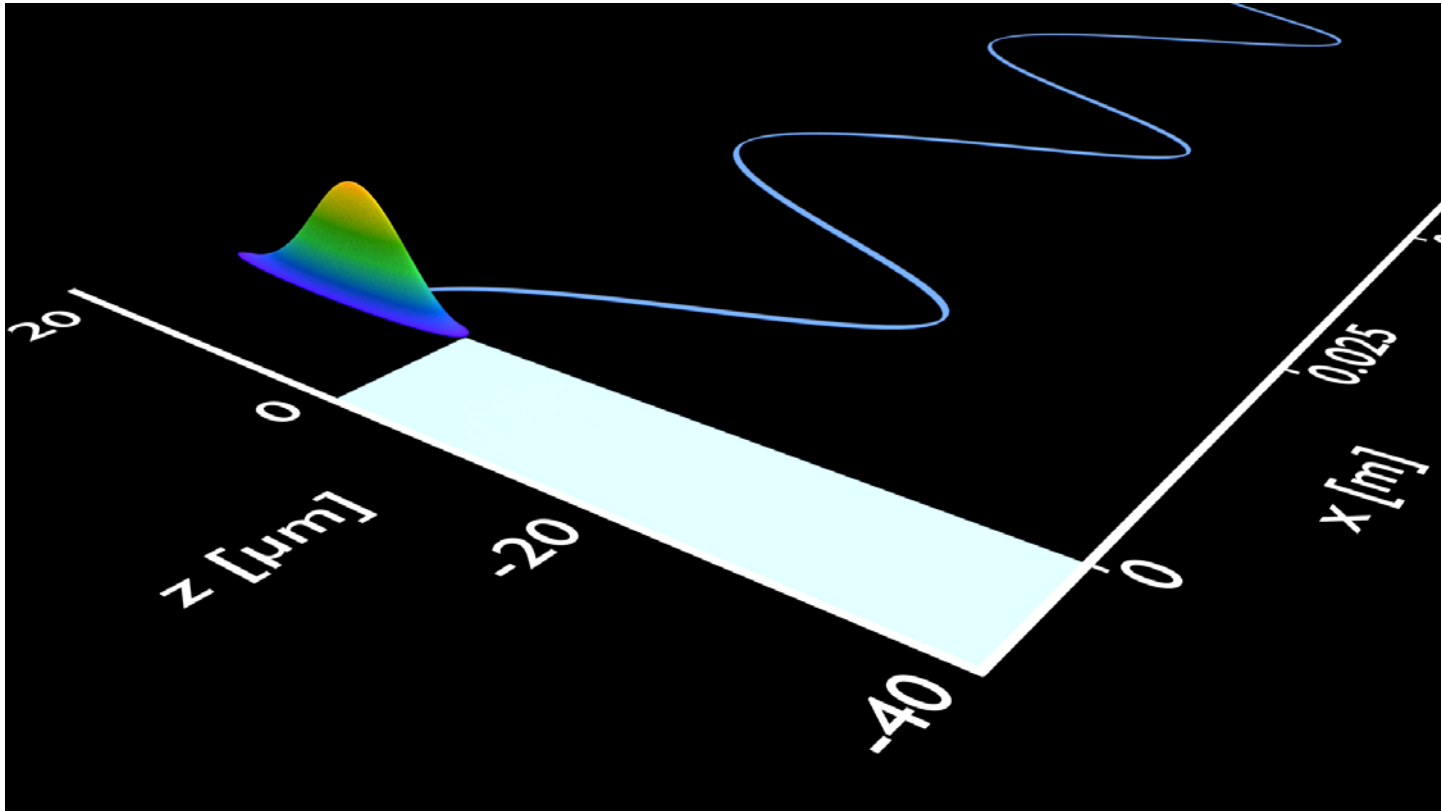
- Quantum interference:
  - sensitivity to fifth forces
  - coming from extra dimensions
  - string theories (higher dimensional field theories)
    - axion fields
- stroboscopic snapshots
  - spatial resolution  $1\mu\text{m}$
  - low background: 1 neutron every 100s

- Snapshots with spatial resolution detectors  $\sim 1.5 \mu\text{m}$

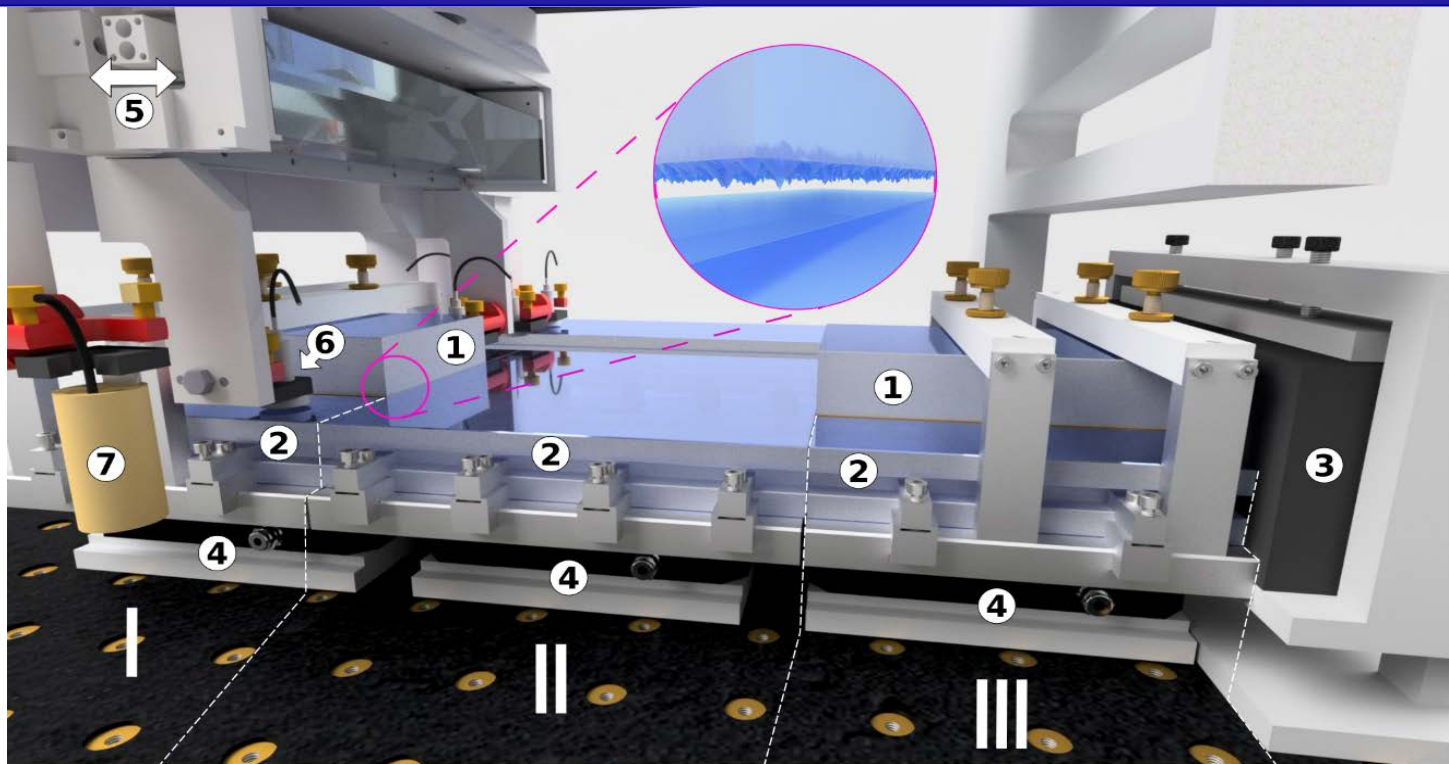
$$\Psi(z, t) = \sum_{n=0}^{\infty} c_n e^{-iE_n t/\hbar} \psi_n(z)$$

$$\psi_n(z) \sim Ai\left[\frac{z}{z_0} - \frac{E_n}{E_0}\right]; c_n = \int_0^{\infty} \Psi(z, 0) \psi(z) dz$$



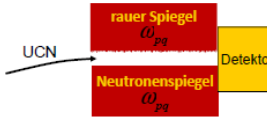
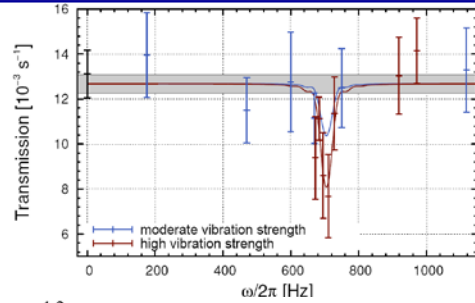


# Acoustic Rabi Transitions



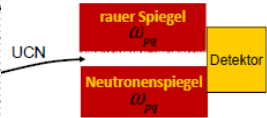
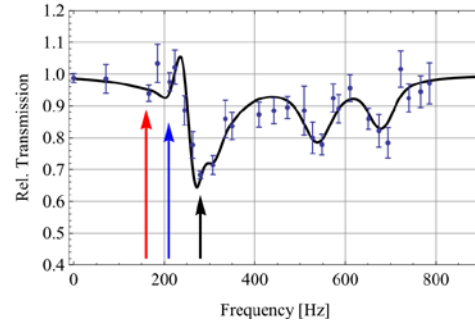
Three regions (marked I, II, III). rough neutron mirror on top (1) the neutron mirror (2) neutron detector (3) All neutron mirrors are mounted on nano-positioning tables (4). An optical system (parts in 5) controls the induced mirror oscillations. A movable system based on highly precise capacitive sensors (6) controls and levels steps between the regions. The experiment is shielded by  $\mu$ -metal against the magnetic field of the Earth. Flux-gate magnetic field sensors (7) log the residual magnetic fields.

# qBounce – Gravity Resonance Spectroscopy



$$|1\rangle \leftrightarrow |3\rangle$$

● T. Jenke et al. NP 2011



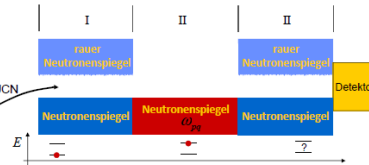
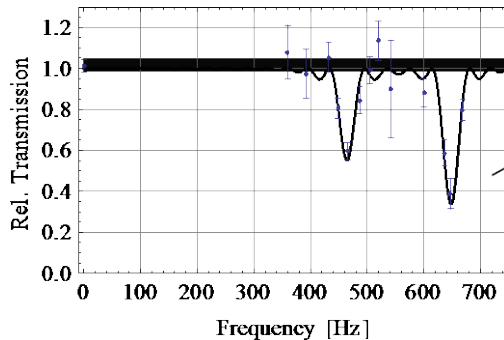
$$|1\rangle \leftrightarrow |2\rangle : 266 \text{ Hz}$$

$$|1\rangle \leftrightarrow |3\rangle : 563 \text{ Hz}$$

$$|2\rangle \leftrightarrow |3\rangle : 296 \text{ Hz}$$

$$|2\rangle \leftrightarrow |4\rangle : 701 \text{ Hz}$$

● T. Jenke et al. PRL 2014



$$|1\rangle \leftrightarrow |3\rangle : 462 \text{ Hz}$$

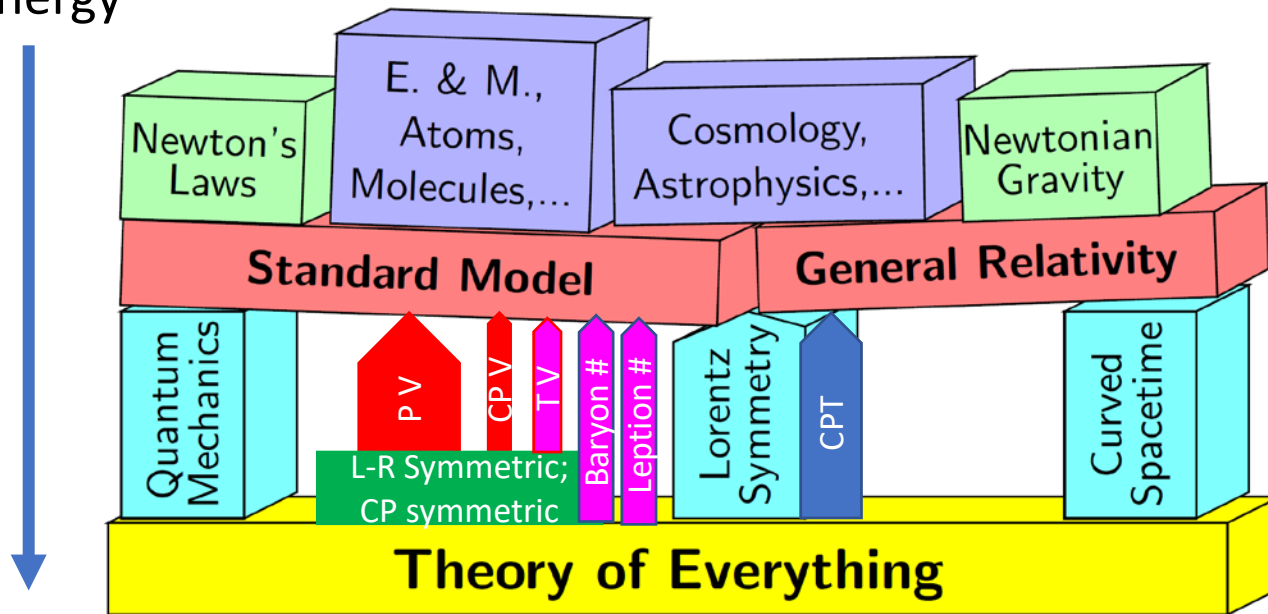
$$|1\rangle \leftrightarrow |4\rangle : 647 \text{ Hz}$$

● C. Cronenberg et al.

# Question: Why Lorentz Violation?

& Discrete Symmetries

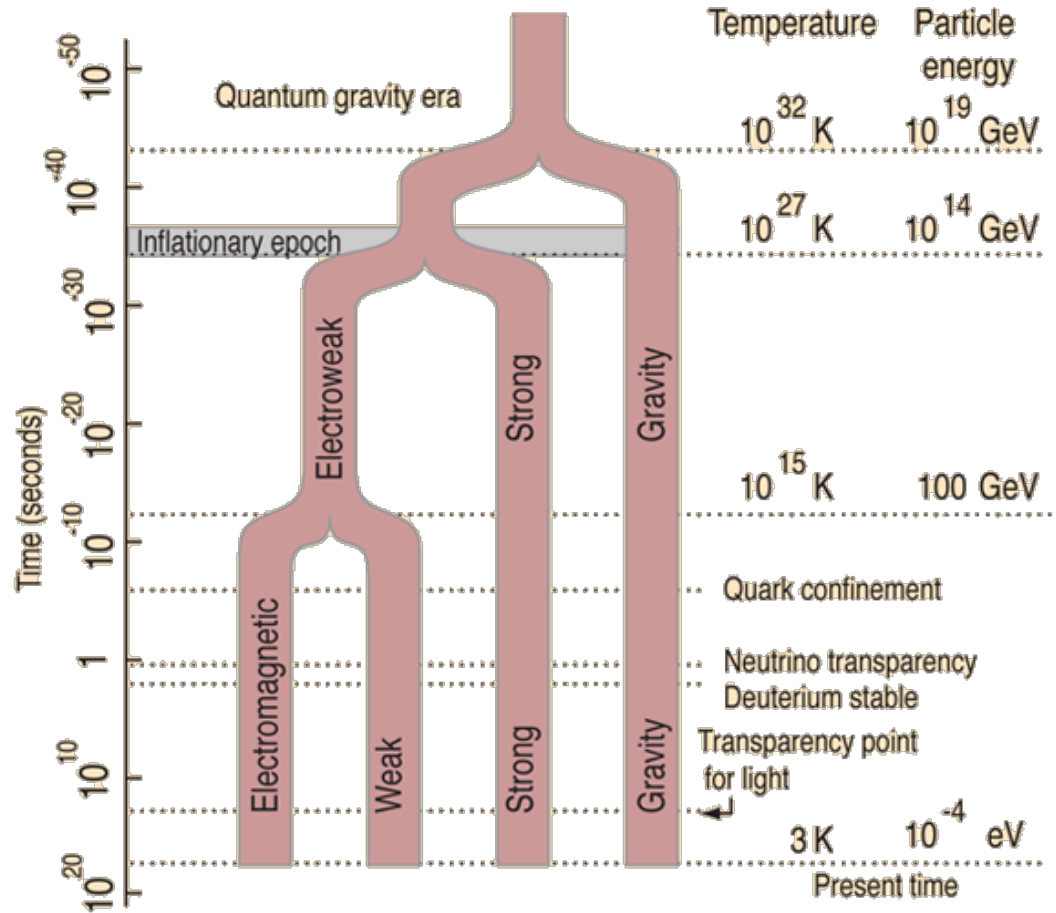
Energy



**Answer:** Symmetry violations (at low E-scales) are evidences, pointing to new physics that unifies all forces at high E-scales.

**Experimental Approach:** Precision measurements on small values of symmetry-violating observables.

# A (Possible) Unified Theory of Everything





# History: Unifications through Symmetries

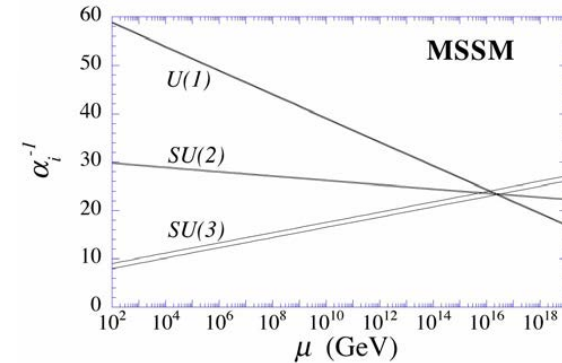
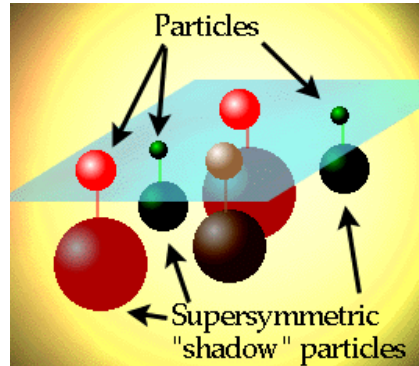
$$\begin{aligned}\nabla \times \nabla \times \mathbf{E} &= -\nabla^2 \mathbf{E} = \\ &= -\mu \frac{\partial}{\partial t} (\nabla \times \mathbf{H}) \\ &= -\mu \frac{\partial}{\partial t} \left( \frac{\partial \mathbf{D}}{\partial t} + \mathbf{J} \right) \\ &= -\mu \epsilon \frac{\partial}{\partial t} \left( \frac{\partial \mathbf{E}}{\partial t} \right)\end{aligned}$$

$$\Rightarrow \nabla^2 \mathbf{E} = \mu \epsilon \frac{\partial^2 \mathbf{E}}{\partial t^2} \quad \text{[The Vector wave equation]}$$

Light = Electromagnetic wave

- **Stern and Gerlach:** Intrinsic spin, properties with respect to the rotation operator  $J$  doubles the number of electron states
- **Dirac:** particle/antiparticle, properties with respect to the Lorentz boost generator,  $K$ , doubling the number of electron states: electron-positron
- **Supersymmetry:** introduces a new generator  $Q$  doubling the number of states once again: electron and scalar electron (selectron)

Mike Berger



$$\begin{pmatrix} \gamma \\ Z^0 \end{pmatrix} = \begin{pmatrix} \cos \theta_W & \sin \theta_W \\ -\sin \theta_W & \cos \theta_W \end{pmatrix} \begin{pmatrix} B \\ W_3 \end{pmatrix}$$

$$M_Z = \frac{M_W}{\cos \theta_W}$$

$$G_F^0 \simeq \frac{\pi \alpha}{\sqrt{2} M_W^2 (1 - M_W^2/M_Z^2)}$$

Electro-weak

Questions?