

Lecture 2 - Neutrino oscillations

- 1) ν oscill. in vacuum (2 ν case)
- 2) ν oscill. in matter - MSW effect (2 ν case)
- 3) Oscillations of 3 ν system

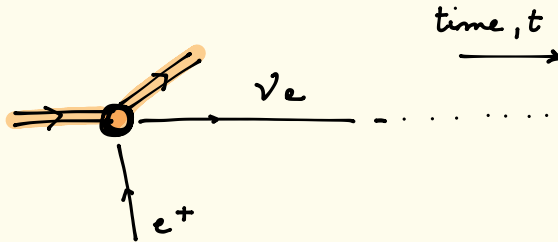
Caveats:

- idealized case of 2 ν states \rightarrow can generalize to 3 ν (realistic)
- standard, traditional formalism \rightarrow more advanced methods address shortcomings
(wavepacket formalism)

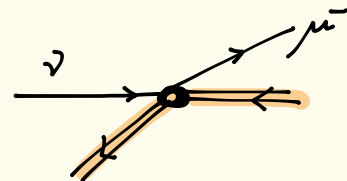
1) ν oscill. in vacuum

Periodic change of (measured) flavor, due to quantum propagation

Produced ν_e :



detected ν_μ :



$$P = P(\nu_e \rightarrow \nu_\mu, t)$$

flavor conversion probability

Toy model: 2 ν system

$\{| \nu_1 \rangle, | \nu_2 \rangle\}$ mass eigenstates

$$\langle \nu_i | \nu_j \rangle = \delta_{ij} \quad (i, j = 1, 2)$$

$\{m_1, m_2\}$ masses

$\{| \nu_e \rangle, | \nu_\mu \rangle\}$ flavor eigenstates

$$\langle \nu_\alpha | \nu_\beta \rangle = \delta_{\alpha\beta} \quad (\alpha, \beta = e, \mu)$$

Notation: $| \nu \rangle$ generic state,

$$\begin{pmatrix} \nu_e \\ \nu_\mu \end{pmatrix} \equiv \begin{pmatrix} \langle \nu_e | \nu \rangle \\ \langle \nu_\mu | \nu \rangle \end{pmatrix}$$

Mixing matrix:

$$\begin{cases} | \nu_e \rangle = c | \nu_1 \rangle + s | \nu_2 \rangle \\ | \nu_\mu \rangle = -s | \nu_1 \rangle + c | \nu_2 \rangle \end{cases}$$

$$\rightarrow U = \begin{pmatrix} c & s \\ -s & c \end{pmatrix} \quad \underline{\text{real}} \quad (\text{no physical phases for 2 } \nu)$$

$$c \equiv \cos\theta \quad s \equiv \sin\theta$$

Hamiltonian in vacuum, in $\{|V_i\rangle\}$ basis ($i=1,2$)

$$H = \begin{pmatrix} E_1 & 0 \\ 0 & E_2 \end{pmatrix}, \text{ time-independent}$$

$$E_i = \sqrt{p^2 + m_i^2} \quad (\text{same } p \text{ assumption})$$

Quantum evolution: $t=0 \quad |V(0)\rangle = |V_e\rangle$

Calculate $P(V_e \rightarrow V_\mu, t)$:

$$|V(t)\rangle = c e^{-iE_1 t} |V_1\rangle + s e^{-iE_2 t} |V_2\rangle$$

$$P(V_e \rightarrow V_\mu, t) = \left| \underbrace{(-s \langle V_1 | + c \langle V_2 |)}_{\langle V_\mu |} |V(t)\rangle \right|^2 = \left| -s c e^{-iE_1 t} + s c e^{-iE_2 t} \right|^2 =$$

► ∴ (steps)

$$P(\nu_e \rightarrow \nu_\mu, t) = \sin^2 2\theta \sin^2\left(\frac{E_2 - E_1}{2} t\right)$$

← depends on
difference of
eigenvalues.

Approximate, relativistic limit:

$$E_i = \sqrt{p^2 + m_i^2} \simeq p + \frac{m_i^2}{2p} \simeq p + \frac{m_i^2}{2E} \quad (p^2 \gg m_i^2)$$

$t \simeq L$ (propagation distance)

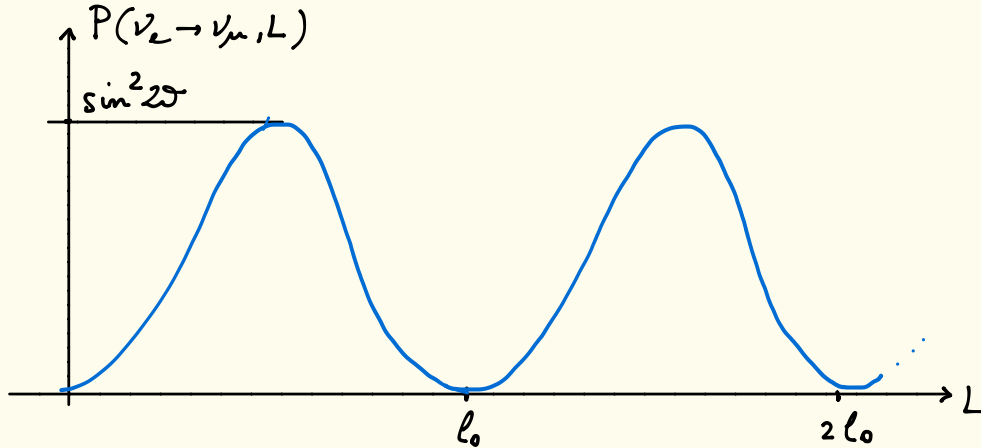
↓

$$P(\nu_e \rightarrow \nu_\mu, L) = \sin^2 2\theta \sin^2\left(\frac{\pi L}{l_0}\right)$$

$$\Delta m^2 \equiv m_2^2 - m_1^2$$

$$l_0 = \frac{4\pi E}{\Delta m^2}$$

oscill. length



$$l_0 = \frac{4\pi E}{\Delta m^2} \approx 2480 \text{ km} \left(\frac{E}{\text{GeV}} \right) \left(\frac{10^{-3} \text{ eV}^2}{\Delta m^2} \right)$$

Note:

- $P(\nu_e \rightarrow \nu_\mu, L=0) = 0$ as should.

- Non trivial solution requires $\theta \neq 0$ AND $\Delta m^2 \neq 0$

- depends on $\Delta m^2 = m_2^2 - m_1^2$, NOT on absolute mass scale

for
2x2
case
only

$$P(\nu_\mu \rightarrow \nu_e, L) = P(\nu_e \rightarrow \nu_\mu, L)$$

$$P(\nu_e \rightarrow \nu_e) = P(\nu_\mu \rightarrow \nu_\mu, L) = 1 - P(\nu_e \rightarrow \nu_\mu, L)$$

$$P(\bar{\nu}_e \rightarrow \bar{\nu}_\mu, L) = P(\nu_e \rightarrow \nu_\mu, L) \quad \underline{\text{CP-conserving}}$$

Averaged oscillations:

consider L fixed, and $P(\nu_e \rightarrow \nu_\mu, E)$ (function of energy):

separation between maxima:

$$\pi = \frac{L \Delta m^2}{4E} - \frac{L \Delta m^2}{4(E + \delta E)} \longrightarrow \delta E \approx \frac{4\pi E^2}{\Delta m^2 L}$$

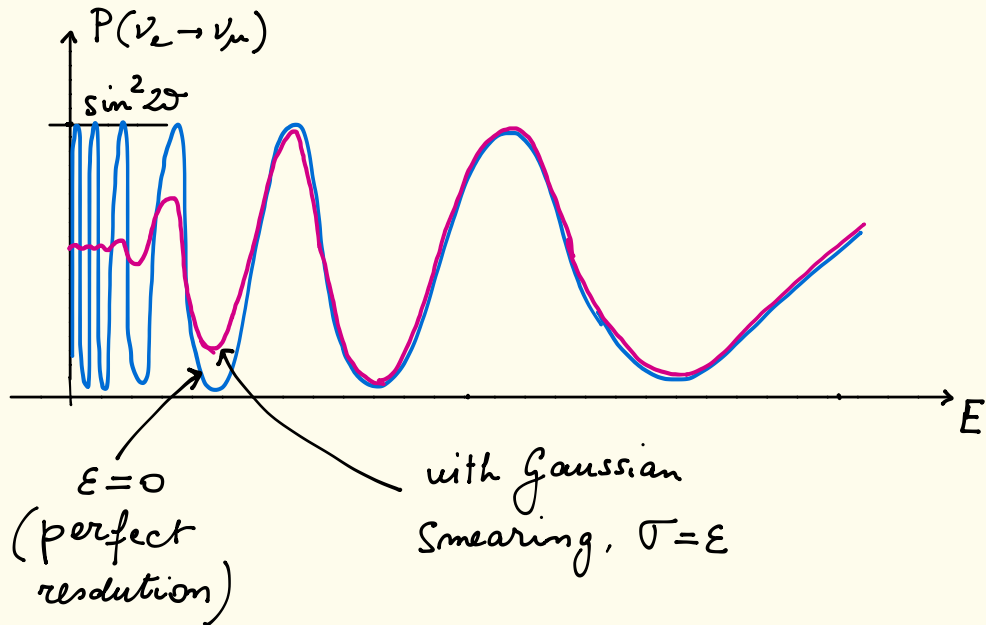
if $\delta E \ll E$ (detector resolution) \rightarrow oscillations can't be resolved!

$$\Rightarrow \boxed{P(\nu_e \rightarrow \nu_\mu) \approx \frac{1}{2} \sin^2 2\theta} \leq \frac{1}{2} \quad \leftarrow \text{avg. vacuum oscill.}$$

can't exceed 50%!

$$\delta E \approx \frac{4\pi E^2}{\Delta m^2 L} \xrightarrow[L \rightarrow \infty]{E \rightarrow 0} 0$$

For L fixed:



2) ν oscill. in matter, 2x2 case

$\left\{ \begin{array}{l} \text{Coherent scattering (refraction): } \vec{p}_\nu \text{ unchanged} \\ \text{Incoherent scattering (absorption)} \end{array} \right. \quad \begin{array}{l} V \propto G_F \\ \sigma \propto G_F^2 E \\ \text{(negligible in} \\ \text{most applications)} \end{array}$

Focus on refraction:

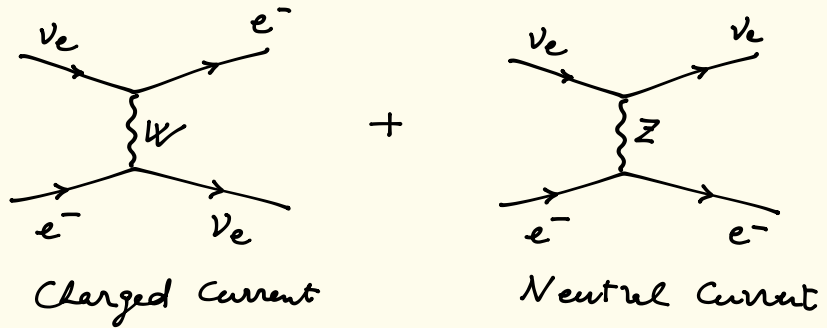
effective treatment: ν in external potential

$$H = H_{\text{vac}} + V$$

MSW: Mikheev, Smirnov (1985), Wolfenstein (1978)

2.1 The refraction potential

Example: $\nu_e - e^-$ scattering, $n_e =$ number density of e^- .



C.C. Only: Hamiltonian (after Fierz transformation):

$$\mathcal{H}_{CC}(x) = \frac{G_F}{\sqrt{2}} [\bar{\nu}_e(x) \gamma^\mu (1 - \gamma^5) \nu_e(x)] [\bar{e}(x) \gamma_\mu (1 - \gamma^5) e(x)]$$

Average over electron background (unpolarized):

$$\overline{\mathcal{H}^{cc}} = \frac{G_F}{\sqrt{2}} [\bar{\nu}_e(x) \gamma^\beta (1 - \gamma^5) \nu_e(x)] \int d^3 p_e f(p_e, T)$$

$$\times \frac{1}{2} \sum_{h_e = \pm 1} \langle e^-(p_e, h_e) | \bar{e}(x) \gamma_\beta (1 - \gamma^5) e(x) | e^-(p_e, h_e) \rangle$$

same!

$f(p_e, T) \leftarrow e^-$ statistical distribution

$$\begin{aligned} \overline{\mathcal{H}^{cc}}(x) &= \underbrace{\sqrt{2} G_F m_e}_{\equiv V_{cc}} \bar{\nu}_{eL}(x) \gamma^0 \nu_{eL}(x) \\ &\equiv V_{cc} \end{aligned}$$

\hookrightarrow potential energy: $V_e^{cc} = V_{cc} = \sqrt{2} G_F m_e$

Similarly, for $\bar{\nu}_e - e^-$ scattering:

$$V_{\bar{e}}^{cc} = -V_{cc}$$

Generalize to realistic matter : e^- , p , n

$$m_p = m_e \quad , \quad m_n$$

$$\left\{ \begin{array}{l} V_e = V_e^{CC} + V_e^{NC} = \sqrt{2} G_F m_e - \frac{1}{2} \sqrt{2} G_F m_n \\ V_{\mu} = \quad \quad V_{\mu}^{NC} = \quad \quad - \frac{1}{2} \sqrt{2} G_F m_n \end{array} \right.$$

$$\left\{ \begin{array}{l} V_{\bar{e}} = V_{\bar{e}}^{CC} + V_{\bar{e}}^{NC} = -\sqrt{2} G_F m_e + \frac{1}{2} \sqrt{2} G_F m_n \\ V_{\bar{\mu}} = \quad \quad V_{\bar{\mu}}^{NC} = \quad \quad + \frac{1}{2} \sqrt{2} G_F m_n \end{array} \right.$$

Note: contributions of NC on e^- and NC on p cancel.

2.2 Constant density: $V_\alpha (\cancel{X}) = \text{const.}$ ($\alpha = e, \mu$)

H in flavor basis: time-independent

$$H_{fl} = U H_{vac} U^\dagger + V = \begin{pmatrix} c & s \\ -s & c \end{pmatrix} \begin{pmatrix} E_1 & 0 \\ 0 & E_2 \end{pmatrix} \begin{pmatrix} c & -s \\ s & c \end{pmatrix} + \begin{pmatrix} V_e & 0 \\ 0 & V_\mu \end{pmatrix}$$

Solution only depends on difference of eigenvalues \Rightarrow can subtract identity terms.

Take $H_{fl} \rightarrow H_{fl} - \mathbb{I} \left(p + \frac{m_1^2 + m_2^2}{4E} + \frac{V_e + V_\mu}{2} \right)$

$\blacktriangleright H_{fl} = \frac{\Delta m^2}{4E} \begin{pmatrix} -\cos 2\theta & \sin 2\theta \\ \sin 2\theta & \cos 2\theta \end{pmatrix} + \begin{pmatrix} V_{cc}/2 & 0 \\ 0 & -V_{cc}/2 \end{pmatrix}$

$$V_{cc} = V_e - V_\mu = \sqrt{2} G_F m_e$$

$$H_{fe} = \frac{\Delta m^2}{4E} \begin{pmatrix} -\cos 2\theta + \kappa & \sin 2\theta \\ \sin 2\theta & \cos 2\theta - \kappa \end{pmatrix}$$

$$\kappa \equiv \frac{2\sqrt{2}G_F m_e E}{\Delta m^2}$$

Calculate $P(\nu_\alpha \rightarrow \nu_\beta, L)$: use t -independent solution (H_{fl} is t -indep)

Diagonalize H : $\{\nu_{1m}, \nu_{2m}\}$ eigenstates, $\{E_{1m}, E_{2m}\}$ eigenvalues

$$U_m = \begin{pmatrix} c_m & s_m \\ -s_m & c_m \end{pmatrix} \text{ mixing matrix}$$
$$c_m \equiv \cos \vartheta_m, \quad s_m = \sin \vartheta_m$$
$$\begin{cases} |\nu_e\rangle = c_m |\nu_{1m}\rangle + s_m |\nu_{2m}\rangle \\ |\nu_\mu\rangle = -s_m |\nu_{1m}\rangle + c_m |\nu_{2m}\rangle \end{cases}$$

$$\begin{cases} \sin 2\vartheta_m = \frac{\sin 2\vartheta}{[\sin^2 2\vartheta + (\cos 2\vartheta - \kappa)^2]^{1/2}} \\ \cos 2\vartheta_m = \frac{\cos 2\vartheta - \kappa}{[\sin^2 2\vartheta + (\cos 2\vartheta - \kappa)^2]^{1/2}} \end{cases}$$

$$E_{2m} - E_{1m} = \frac{\Delta m^2}{2E} [\sin^2 2\vartheta + (\cos 2\vartheta - \kappa)^2]^{1/2}$$

$$\kappa \equiv \frac{2\sqrt{2} G_F m_e E}{\Delta m^2}$$

Calculate $P(\nu_e \rightarrow \nu_\mu, t)$: Same steps as vacuum

$$|\nu(t)\rangle = c_m e^{-iE_{1m}t} |\nu_{1m}\rangle + s_m e^{-iE_{2m}t} |\nu_{2m}\rangle$$

$$P(\nu_e \rightarrow \nu_\mu, t) = \left| \left(-s_m \langle \nu_{1m} | + c_m \langle \nu_{2m} | \right) |\nu(t)\rangle \right|^2$$

∴ (steps)

$$\Rightarrow P(\nu_e \rightarrow \nu_\mu, L) = \sin^2 2\theta_m \sin^2 \left(\pi \frac{L}{l_m} \right)$$

$$\triangleright l_m = \frac{2\pi}{E_{2m} - E_{1m}} = \frac{l_0}{[\sin^2 2\theta + (\cos 2\theta - \alpha)^2]^{1/2}}$$

$$(l_0 = \frac{4\pi E}{\Delta m^2})$$

$$H_{fe} = \frac{\Delta m^2}{4E} \begin{pmatrix} -\cos 2\theta + \kappa & \sin 2\theta \\ \sin 2\theta & \cos 2\theta - \kappa \end{pmatrix}$$

$$\kappa \equiv \frac{2\sqrt{2} G_F m_e E}{\Delta m^2}$$

$$\begin{cases} \sin 2\theta_m = \frac{\sin 2\theta}{[\sin^2 2\theta + (\cos 2\theta - \kappa)^2]^{\frac{1}{2}}} \\ \cos 2\theta_m = \frac{\cos 2\theta - \kappa}{[\sin^2 2\theta + (\cos 2\theta - \kappa)^2]^{\frac{1}{2}}} \end{cases}$$

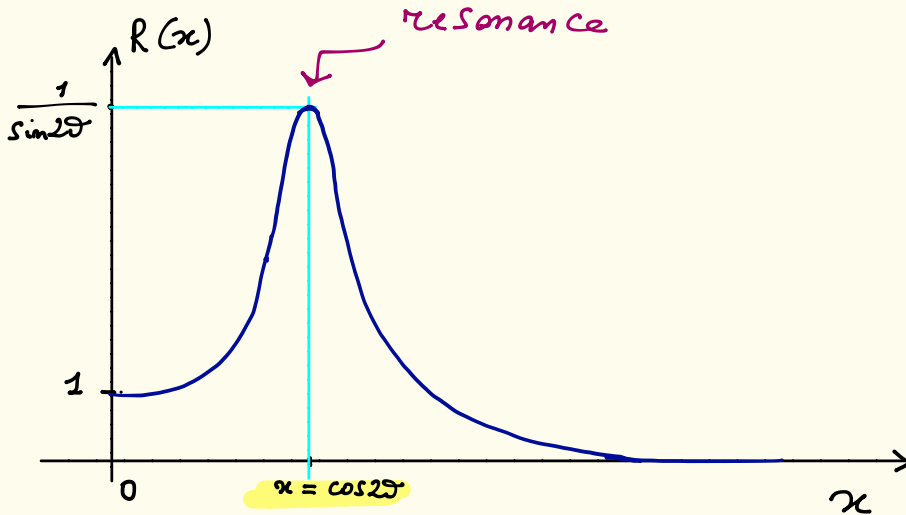
$$E_{2m} - E_{1m} = \frac{\Delta m^2}{2E} [\sin^2 2\theta + (\cos 2\theta - \kappa)^2]^{\frac{1}{2}}$$

$$l_m = \frac{2\pi}{E_{2m} - E_{1m}} = \frac{l_0}{[\sin^2 2\theta + (\cos 2\theta - \kappa)^2]^{\frac{1}{2}}}$$

$$P(\nu_e \rightarrow \nu_\mu, L) = \sin^2 2\theta_m \sin^2 \left(\pi \frac{L}{l_m} \right)$$

Note:

- Vacuum solution recovered for $\kappa=0$ ($m_e=0$), as should be.
- Resonant character: factor $\frac{1}{[\sin^2 2\vartheta + (\cos 2\vartheta - \kappa)^2]^{\frac{1}{2}}} \equiv R(\kappa)$

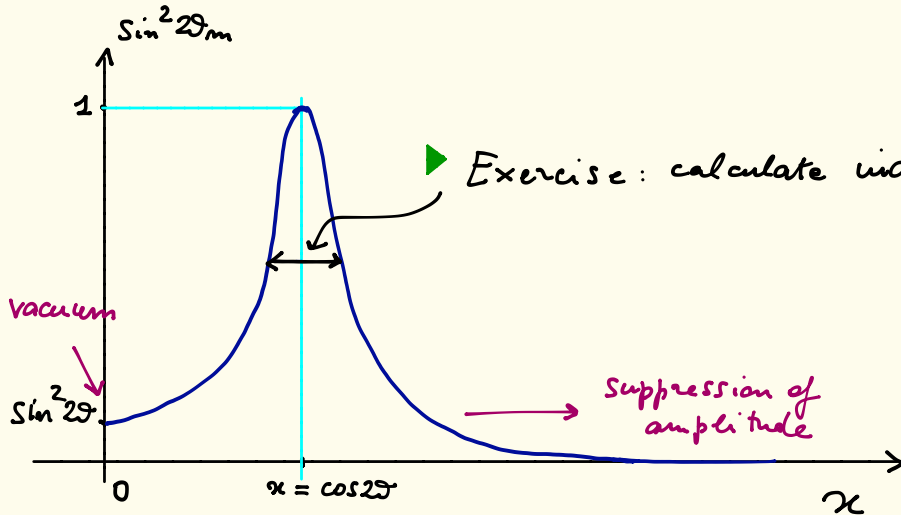


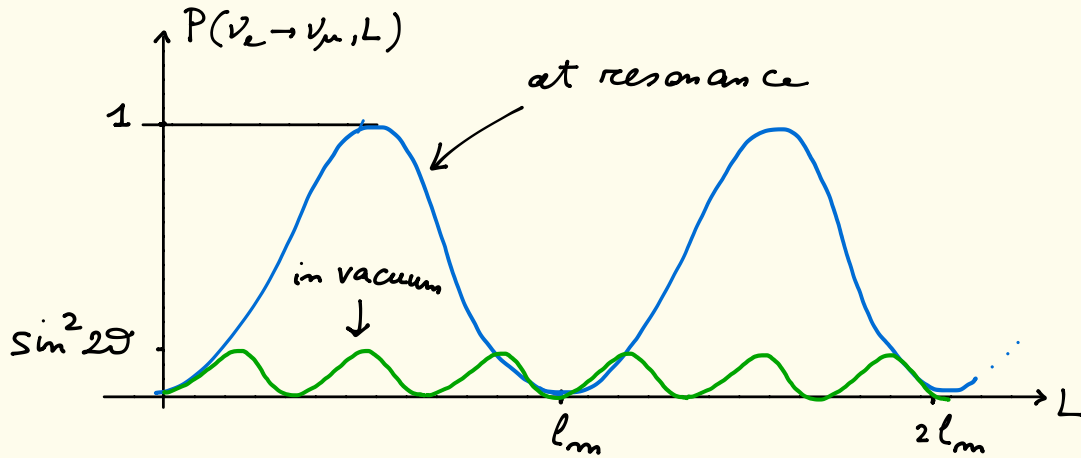
At resonance:

$$\sqrt{2} G_F m_e = \frac{\Delta m^2 \cos 2\theta}{2E}$$

(MSW resonance condition)

- $E_{1m} - E_{2m} \propto R(x)^{-1}$ has a minimum
- $l_{m} \propto R(x)$ has a maximum
- $\sin^2 2\theta_m \propto R(x)^2$ has a maximum \rightarrow max amplitude of oscill.





- Useful : resonance energy, resonance density ($\rho = 2m_N m_e$)

$$E_R = \frac{\Delta m^2 \cos 2\theta}{2\sqrt{2} G_F m_e} = 14 \text{ GeV} \left(\frac{\Delta m^2}{10^{-3} \text{ eV}^2} \right) \left(\frac{1 \text{ g cm}^{-3}}{\rho} \right) \cos 2\theta$$

$$m_{eR} = \frac{\Delta m^2 \cos 2\theta}{2\sqrt{2} G_F E}$$

$$\hookrightarrow \rho_R = 2m_N m_{eR} = 1.4 \text{ g cm}^{-3} \cos 2\theta \left(\frac{\Delta m^2}{10^{-3} \text{ eV}^2} \right) \left(\frac{10 \text{ GeV}}{E} \right)$$

- Resonance exists only if $\Delta m^2 > 0$!

For $\Delta m^2 < 0$ $\sin^2 2\theta_m \leq \sin^2 2\theta \rightarrow$ amplitude always suppressed.

- for antineutrinos: $V_\alpha \rightarrow -V_\alpha < 0$

Resonance for $\Delta m^2 < 0$, suppression for $\Delta m^2 > 0$.

Resonance affects only ν or $\bar{\nu}$, not both.

2.2) Varying density: $\rho = \rho(t)$, 2×2 case

$H = H(t) \rightarrow$ need full Schrodinger equation.

Idea: use basis of instantaneous eigenstates:

$$|v_{1m}(t)\rangle, |v_{2m}(t)\rangle, \quad \underline{\theta_m = \theta_m(t)}$$

$$\begin{cases} |v_e\rangle = c_m(t) |v_{1m}(t)\rangle + s_m(t) |v_{2m}(t)\rangle \\ |v_\mu\rangle = -s_m(t) |v_{1m}(t)\rangle + c_m(t) |v_{2m}(t)\rangle \end{cases} \longrightarrow U_m = \begin{pmatrix} c_m & s_m \\ -s_m & c_m \end{pmatrix}$$

$$c_m(t) \equiv \cos \theta_m(t), \quad s_m(t) = \sin \theta_m(t)$$

$$U(\theta_m(t)) = U_m(t)$$

Schrodinger equation, flavor basis:

$$i \frac{d\nu_{fl}}{dt} = H_{fl} \nu_{fl}$$

$$\nu_{fl} = U_m(t) \nu_m$$

Rewrite in instantaneous matter basis:

$$i \frac{d}{dt} (U_m(t) \nu_m(t)) = H_{fl} (U_m(t) \nu_m(t))$$

$$i \left(\frac{d}{dt} U_m \right) \nu_m + i U_m \frac{d}{dt} \nu_m = H_{fl} U_m \nu_m$$

↓ × U_m^+ :

$$i \frac{d}{dt} \nu_m = \left[U_m^+ H_{fl} U_m - \underbrace{i U_m^+ \left(\frac{d}{dt} U_m \right)} \right] \nu_m$$

depends on $\dot{\theta}_m = \frac{d}{dt} \theta_m$

Result, explicit form:

$$i \frac{d}{dt} \begin{pmatrix} v_{1m} \\ v_{2m} \end{pmatrix} = \begin{pmatrix} E_{1m}(t) & -i \dot{\theta}_m(t) \\ i \dot{\theta}_m(t) & E_{2m}(t) \end{pmatrix} \begin{pmatrix} v_{1m} \\ v_{2m} \end{pmatrix}$$

Note: off-diagonal terms $\propto \dot{\theta}_m \rightarrow$ cause $v_{1m} \leftrightarrow v_{2m}$ transitions

A diabatic approximation: neglect $\dot{\theta}_m$

Adiabaticity condition :

$$|\dot{\theta}_m| \ll |E_{1m} - E_{2m}|$$

$$\gamma^{-1} \equiv \frac{2|\dot{\theta}_m|}{|E_{1m} - E_{2m}|} \quad \blacktriangleright \quad \frac{\sin 2\theta \Delta m^2 / 2E}{|E_{1m} - E_{2m}|^3} |\dot{V}_{cc}| \ll 1$$

Meaning: potential varies slowly

→ $\nu_{1m} \leftrightarrow \nu_{2m}$ transitions suppressed

→ $\langle \nu_{1m} | \nu \rangle$, $\langle \nu_{2m} | \nu \rangle$ vary only by a phase

Apply adiabatic approximation:

t_i : initial time

t_f : final time

$$|v(t_i)\rangle = |v_e\rangle = \cos\theta_{mi} |v_{1m}\rangle + \sin\theta_{mi} |v_{2m}\rangle$$

$$|v(t_f)\rangle = \cos\theta_{mi} e^{-i \int_{t_i}^{t_f} E_{1m}(t') dt'} |v_{1m}\rangle + \sin\theta_{mi} e^{-i \int_{t_i}^{t_f} E_{2m}(t') dt'} |v_{2m}\rangle$$

at $t = t_f$: $|v_\mu\rangle = -\sin\theta_{mf} |v_{1m}\rangle + \cos\theta_{mf} |v_{2m}\rangle$

$$P(v_e \rightarrow v_\mu, t) = \left| \langle v_\mu | v(t_f) \rangle \right|^2$$

Full result:

$$\blacktriangleright P(\nu_e \rightarrow \nu_\mu, t) = \frac{1}{2} - \frac{1}{2} \cos 2\theta_{mi} \cos 2\theta_{mf} - \frac{1}{2} \sin 2\theta_{mi} \sin 2\theta_{mf} \cos \varphi$$

$$\varphi = \int_{t_i}^{t_f} (E_{1m}(t') - E_{2m}(t')) dt'$$

If $\sin 2\theta_{mi} \simeq 0$ OR $\sin 2\theta_{mf} \simeq 0$ OR $\langle \cos \varphi \rangle \simeq 0$
(averaged oscill.):

$$P(\nu_e \rightarrow \nu_\mu, t) = \frac{1}{2} - \frac{1}{2} \cos 2\theta_{mi} \cos 2\theta_f$$

Common situation : $m_e(t_i) \gg m_{eR}$ $\rightarrow \cos 2\theta_{mi} \simeq -1$

$m_e(t_f) \ll m_{eR} \rightarrow \cos 2\theta_{mf} \simeq \cos 2\theta$

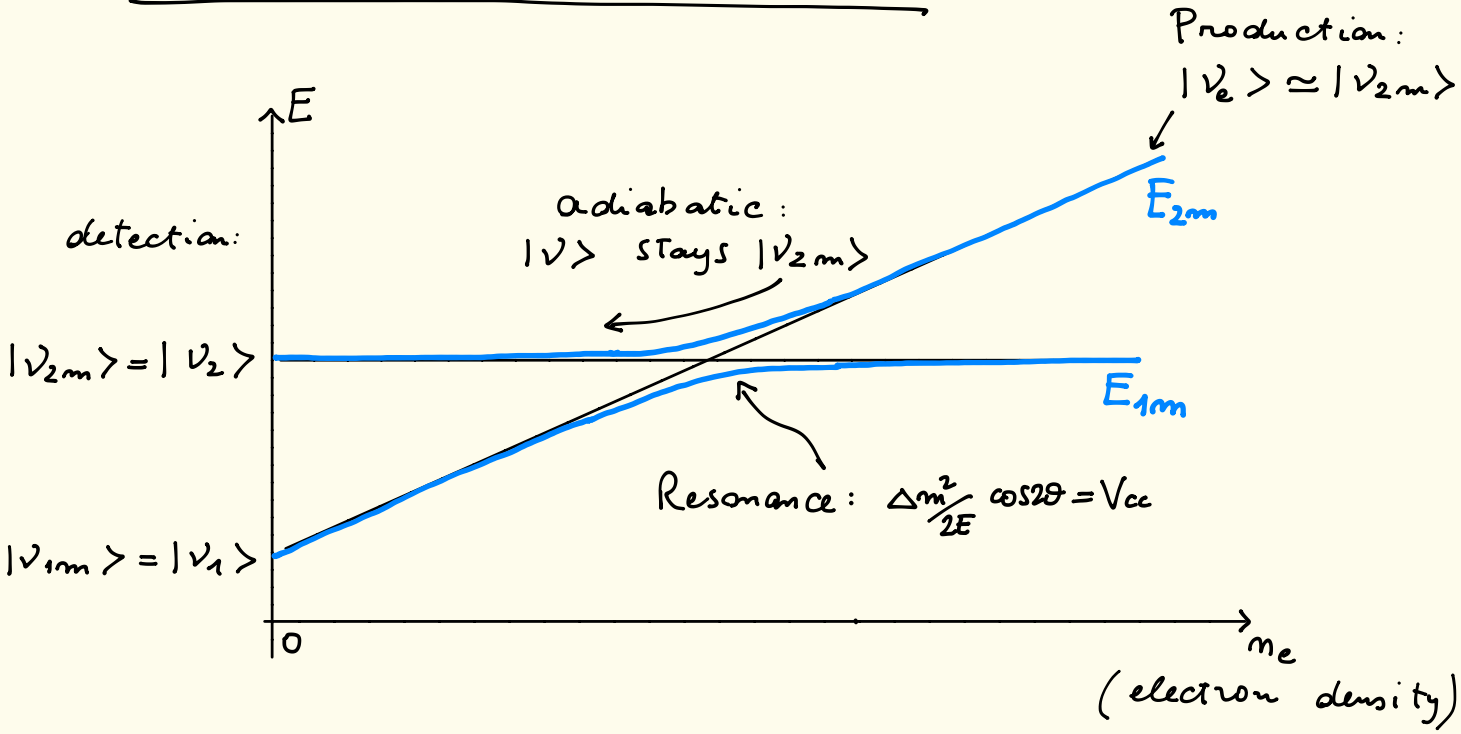
$$\begin{aligned} \hookrightarrow P(\nu_e \rightarrow \nu_\mu, t) &\simeq \frac{1}{2} (1 + \cos 2\theta) \\ &= \cos^2 \theta \geq \frac{1}{2} \end{aligned}$$

Strong flavor
conversion!
($\theta < \pi/4$)

Qualitative picture:

at $t = t_i$, $|\nu_e\rangle \sim |\nu_{2m}\rangle$
adiabatic evolution: $|\nu_{2m}\rangle \longrightarrow |\nu_2\rangle$ (up to a phase)
at t_f : $P(\nu_e \rightarrow \nu_\mu) = |\langle \nu_\mu | \nu_2 \rangle|^2 = \cos^2 \theta$

Useful: level-crossing diagram



A paradox?

$$P(\nu_e \rightarrow \nu_\mu, t) \approx \frac{1}{2}(1 + \cos 2\vartheta) \xrightarrow{\vartheta \rightarrow 0} 1$$

Mixing-less
Conversion ???

No, for $\vartheta \rightarrow 0$ adiabatic condition is violated!

Generalization to non-adiabatic propagation (for $\vartheta \ll \pi/4$)

$$\gamma^{-1} \equiv \frac{2|\dot{\vartheta}_{nm}|}{|E_{1m} - E_{2m}|} = \frac{\sin 2\vartheta \frac{\Delta m^2}{2E}}{|E_{1m} - E_{2m}|^3} |\dot{V}_{ee}|$$

At resonance: $|E_{1m} - E_{2m}|$ is min. $\rightarrow \gamma^{-1}$ is max.

\rightarrow max probability of $\nu_{1m} \leftrightarrow \nu_{2m}$

γ at resonance:

$$\gamma_r = \frac{\sin^2 2\vartheta}{\cos 2\vartheta} \frac{\Delta m^2}{2E} \left| \frac{\dot{m}_e}{m_e} \right|_{res}^{-1} \xrightarrow{\vartheta \rightarrow 0} 0 \quad \text{max. violation of adiabaticity}$$

Generalized probability (oscill. averaged):

$$P(\nu_e \rightarrow \nu_\mu, t) = \frac{1}{2} - \frac{1}{2} \cos 2\theta_{mi} \cos 2\theta_{mf} (1 - 2 P_{LZ})$$

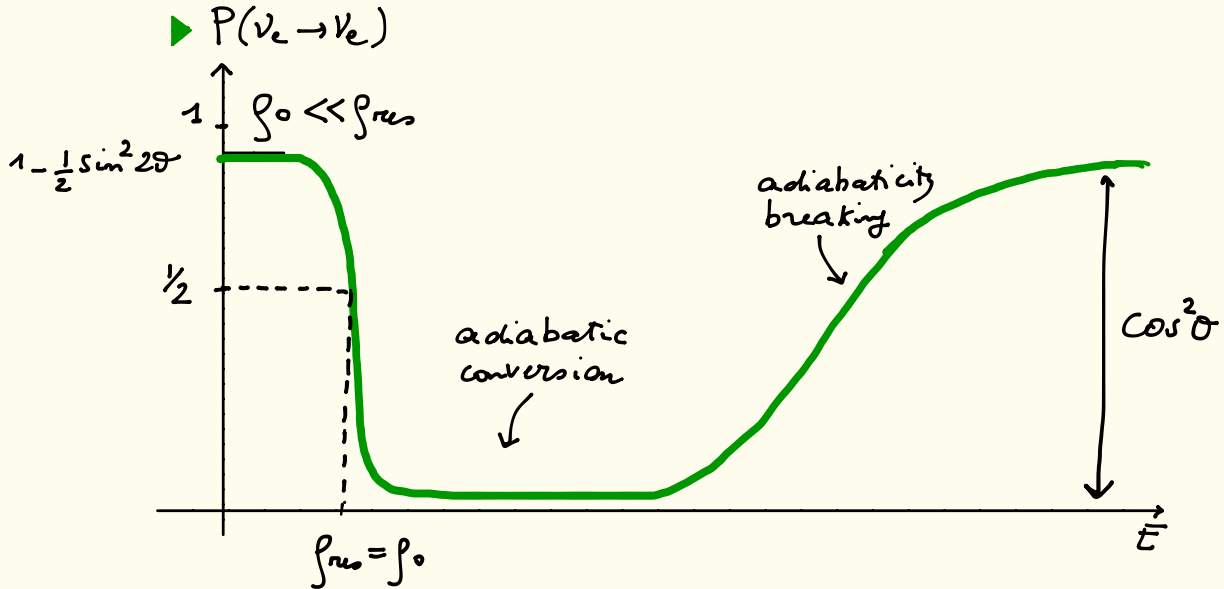
$$P_{LZ} = e^{-\frac{\pi}{2} \delta_2}$$

Landau-Zener hopping probability ($\theta \ll \pi/4$)

Useful exercise: $P(\nu_e \rightarrow \nu_e)$ as function of E (oscill. averaged)

Example: Sun: $\rho(r) \approx \rho_0 e^{-r/R_\odot}$ $\rho_0 \sim 10^2 \text{ g cm}^{-3}$

Take $\theta \ll \pi/4$ (small mixing)



Beyond basics: 3 ν in vacuum

$$\{\nu_1, \nu_2, \nu_3\}, \{m_1, m_2, m_3\} \quad \{\nu_e, \nu_\mu, \nu_\tau\}$$

Mixing matrix:

$$U = \begin{pmatrix} c_{12}c_{13} & s_{12}c_{13} & s_{13}e^{-i\delta} \\ -s_{12}c_{23} - c_{12}s_{23}s_{13}e^{i\delta} & c_{12}c_{23} - s_{12}s_{23}s_{13}e^{i\delta} & s_{23}c_{13} \\ s_{12}s_{23} - c_{12}c_{23}s_{13}e^{i\delta} & -c_{12}s_{23} - s_{12}c_{23}s_{13}e^{i\delta} & c_{23}c_{13} \end{pmatrix}$$

$$\theta_{12}, \theta_{23}, \theta_{13}$$

$$c_{ij} = \cos\theta_{ij}$$

$$s_{ij} = \sin\theta_{ij}$$

[Note: Majorana phases do not affect oscillations]

A note on notation

$\nu_\alpha \equiv \langle \nu_\alpha | \nu \rangle$ component of flavor $\alpha = e, \mu, \tau$

$\nu_\alpha = U_{\alpha j} \nu_j$ (summed over $j = 1, 2, 3$):

$$\begin{pmatrix} \nu_e \\ \nu_\mu \\ \nu_\tau \end{pmatrix} = \begin{pmatrix} U_{e1} & U_{e2} & U_{e3} \\ U_{\mu1} & U_{\mu2} & U_{\mu3} \\ U_{\tau1} & U_{\tau2} & U_{\tau3} \end{pmatrix} \begin{pmatrix} \nu_1 \\ \nu_2 \\ \nu_3 \end{pmatrix}$$

Instead:

$$|\nu_\alpha\rangle = U_{\alpha j}^* |\nu_j\rangle$$

Calculate probabilities : Vacuum case

$$P(\nu_\alpha \rightarrow \nu_\beta, t) = |A(\nu_\alpha \rightarrow \nu_\beta)|^2$$

$$A(\nu_\alpha \rightarrow \nu_\beta) = \left(\sum_{i,j} \langle \nu_i | U_{\beta i} \right) \left(\sum_i e^{-i E_j t} U_{\alpha j}^* | \nu_j \rangle \right)$$

▶ ∴ (steps)

$$P(\nu_\alpha \rightarrow \nu_\beta, t) = \sum_{i,j} U_{\alpha j}^* U_{\beta j} U_{\beta i}^* U_{\alpha i} e^{-i(E_j - E_i)t}$$

Effective 2ν descriptions : use $\Delta m_{21}^2 \ll |\Delta m_{31}^2| \simeq |\Delta m_{32}^2|$

1) $\frac{\Delta m_{21}^2}{2E} L \ll 1 \quad \rightarrow \text{take } \Delta m_{21}^2 = 0$

▶
$$P(\nu_e \rightarrow \nu_e, L) \simeq 1 - \sin^2 2\theta_{13} \sin^2 \left(\frac{\Delta m_{31}^2}{4E} L \right)$$

(ν beams)

2) $\left| \frac{\Delta m_{31}^2}{2E} L \right| \sim \left| \frac{\Delta m_{32}^2}{2E} L \right| \gg 1 \rightarrow 3 \rightarrow 1, 3 \rightarrow 2$ oscill. average out

▶
$$P(\nu_e \rightarrow \nu_e, L) \simeq S_{13}^2 + C_{13}^2 P_{2\nu}(\Delta m_{21}^2, \theta_{12})$$

(reactor ν)

Exact solution of 3×3 case :

$$P(\nu_\alpha \rightarrow \nu_\beta, t) = \sum_{i,j} U_{\alpha j}^* U_{\beta j} U_{\beta i}^* U_{\alpha i} e^{-i(E_j - E_i)t}$$

► \vdots (steps, use unitarity)

$$P(\nu_\alpha \rightarrow \nu_\beta, L) = \delta_{\alpha\beta} - 4 \sum_{j>i} \text{Re} [U_{\alpha j}^* U_{\beta j} U_{\beta i}^* U_{\alpha i}] \sin^2 \left(\frac{\Delta m_{ij}^2 L}{4E} \right) \\ + 2 \sum_{j>i} \text{Im} [U_{\alpha j}^* U_{\beta j} U_{\beta i}^* U_{\alpha i}] \sin \left(\frac{\Delta m_{ij}^2 L}{2E} \right)$$

CP symmetry on oscillation probabilities

$$\nu_{\alpha L} = U_{\alpha j} \nu_{jL} \quad : \quad P(\nu_{\alpha L} \rightarrow \nu_{\beta L}, t) \quad \checkmark$$

$\downarrow C$

$$C : \quad \bar{\nu}_{\alpha L} = U_{\alpha j}^* \bar{\nu}_{jL} \quad P(\bar{\nu}_{\alpha L} \rightarrow \bar{\nu}_{\beta L}, t) \quad \times \text{not observed!}$$

(charge conjugation)

$\downarrow P$

$$CP : \quad \bar{\nu}_{\alpha R} = U_{\alpha j}^* \bar{\nu}_{jR} \quad P(\bar{\nu}_{\alpha R} \rightarrow \bar{\nu}_{\beta R}, t) \quad \checkmark$$

(charge and parity)

\Leftrightarrow CP symmetry :

$$P(\nu_{\alpha} \rightarrow \nu_{\beta}) = P(\bar{\nu}_{\alpha} \rightarrow \bar{\nu}_{\beta})$$

Look for CP violation:

calculate $P(\nu_\alpha \rightarrow \nu_\beta) - P(\bar{\nu}_\alpha \rightarrow \bar{\nu}_\beta) \equiv \Delta P_{\alpha\beta}$

$$P(\nu_\alpha \rightarrow \nu_\beta, L) = \delta_{\alpha\beta} - 4 \sum_{j>i} \text{Re}[U_{\alpha j}^* U_{\beta j} U_{\beta i}^* U_{\alpha i}] \sin^2\left(\frac{\Delta m_{ij}^2 L}{4E}\right) \\ + 2 \sum_{j>i} \text{Im}[U_{\alpha j}^* U_{\beta j} U_{\beta i}^* U_{\alpha i}] \sin\left(\frac{\Delta m_{ij}^2 L}{2E}\right)$$

$$\blacktriangleright P(\bar{\nu}_\alpha \rightarrow \bar{\nu}_\beta, L) = \delta_{\alpha\beta} - 4 \sum_{j>i} \text{Re}[U_{\alpha j}^* U_{\beta j} U_{\beta i}^* U_{\alpha i}] \sin^2\left(\frac{\Delta m_{ij}^2 L}{4E}\right) \\ - 2 \sum_{j>i} \text{Im}[U_{\alpha j}^* U_{\beta j} U_{\beta i}^* U_{\alpha i}] \sin\left(\frac{\Delta m_{ij}^2 L}{2E}\right)$$

$$\Delta P_{\alpha\beta} = 4 \sum_{j>i} \text{Im}[U_{\alpha j}^* U_{\beta j} U_{\beta i}^* U_{\alpha i}] \sin\left(\frac{\Delta m_{ij}^2 L}{2E}\right)$$

CP violation in the lepton sector?

$$\triangleright \Delta P_{e\mu} = \Delta P_{\mu\tau} = \Delta P_{e\tau} = 4 s_{12} c_{12} s_{13} c_{13}^2 s_{23} c_{23} \sin \delta \sum_{j>i} \sin \left(\frac{\Delta m_{ij}^2}{2E} L \right)$$

CP is violated if

$$\Delta P_{\alpha\beta} \neq 0 \rightarrow \sin \delta \neq 0$$

Note:

$$\Delta P_{\alpha\beta} = 0 \text{ if any } \theta_{ij} = 0 \rightarrow \text{CP is a 3}\nu \text{ effect!}$$

3 ν propagation in matter: (outline only)

- $V_{\bar{\alpha}} = -V_{\alpha} \rightarrow P(\nu_{\alpha} \rightarrow \nu_{\beta}) \neq P(\bar{\nu}_{\alpha} \rightarrow \bar{\nu}_{\beta}) \rightarrow$ mimics effect of δ
consequence of CP-asymmetric background ($m_e \neq m_{\bar{e}}$)
- generalized MSW: factorization in multiple 2 ν problems

For further study: ν in ν background

- ν - ν scattering : non-linear effects
relevant for supernovae, cosmology

[e.g. Duan & Kneller, arxiv:0904.0974]