Hydrodynamics in Heavy Ion Collisions

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Hydrodynamics of QGP

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- How does one describe a system of $N \gg 1$ bodies?
- Depends on how much "information" one wants.
- Worst case: $\psi(x_1, \cdots, x_N)$
- Best case: Minimum information content. I.e. maximum entropy. Need only a handful of numbers such as temperature and chemical potential.
- $n(p) = 1/(e^{E_p/T \mu/T} \mp 1)$
- These quantities are the Langrange multipliers that constraints conserved quantities such as energy and charge.

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- How does one describe a system of $N \gg 1$ bodies?
- Depends on how much "information" one wants.
- Worst case: $\psi(x_1, \cdots, x_N)$
- Best case: Minimum information content. I.e. maximum entropy. Need only a handful of numbers such as temperature, and chemical potential. But locally.
- $n(p, x) = 1/(e^{p_{\mu}u^{\mu}/T(x)-\mu(x)/T(x)} \mp 1)$
- You only need to know few functions: *T*(*t*, **x**), μ(*t*, **x**) as well as the collective velocity **u**(*t*, **x**)
- These quantities are the Langrange multipliers that constraints conserved quantities such as energy, momentum and charge.

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Schematic idea of hydro evolution





- System is made up of "fluid cells".
- Each fluid cell feels a force according to the pressure difference (gradient) w.r.t. its neighbors
- System evolves by "flowing" towards lower pressure
- NR: **F** = *m***a**

 $-\nabla P = n_m \partial_t \mathbf{u}$

where *n_m*: mass density,

P: pressure,

u: flow velocity

- What is the nature of the initial condition?
- Do we reach local equilibrium in heavy ion collisions?
- How hot is it?
- How viscous is QGP?
- (Is there a phase transition? If so what kind?)

• Information content of single particle spectra

$$\frac{dN_i}{dy \, d^2 p_T} = \frac{dN_i}{2\pi p_T dp_T dy} \left(1 + \sum_{n=1}^{\infty} 2v_{i,n}(p_T, y) \cos(n\phi)\right)$$

- "Flow": *v*_{*i*,*n*}(*p*_{*T*})
- Came from

$$\varepsilon(\mathbf{x}_T,\eta) = \varepsilon(r_T,\eta) \left(1 + \sum_{n=1}^{\infty} 2\epsilon_n(r_T,\eta)\cos(n\phi)\right)$$

- **Pressure** converts it into $v_{i,n}(p_T)$
- History matters
- $\epsilon_n \rightarrow v_{i,n}$ conversion contains information on the medium and its evolution





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• Triangular Flow $-\cos(3\phi)$ component



[Alver and Roland, Phys.Rev.C81:054905, 2010]

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- Flows are about: Pressure converting Spatial morphology $\epsilon_n(r_T, \eta) \implies$ Momentum space morphology $v_n(p_T, y)$
- This is sensitive to
 - Initial Conditions
 - Flow dynamics (η/s)
 - Equation of State (to a less extent)

- Why is initial condition important?
 - Initial temperature (distribution) $T_0 > T_c$
 - Beginning time of hydro (\sim thermalization time) au_0
 - The size of the hot spots σ_0
 - What happens before the hydro stage?
- v_2 alone cannot determine all these $\implies v_3, v_4, \cdots$

- Why is η/s important?
 - One of the central properties of QGP
 - Calculable in perturbative QCD $\eta/s \sim 1/g^4 \ln(1/g)$
 - Calculable in AdS/CFT $\eta/s = 1/4\pi$
 - If $\eta/s \sim 1/4\pi$, QGP must be sQGP

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• Equations for *T*, **u** – Conservation laws

 $\partial_\mu \left< T^{\mu
u} \right> = 0$

- Stress-energy tensor *T^{μν}* has only 10 d.o.f. Cons. laws provide 4 constraints ⇒ 6 d.o.f. left.
- No dynamical content yet.
- Energy density and flow vector

$$T^{\mu\nu}u_{\nu}=-\varepsilon u^{\mu}$$

- u^{μ} : Time-like eigenvector of $T^{\mu\nu}$. Normalized to $u^{\mu}u_{\mu} = -1$.
- ε : Local energy density
- This is always possible since $T^{\mu\nu}$ is real and symmetric.

Hydrodynamics

So far:

$$T^{\mu
u} = arepsilon u^{\mu} u^{
u} + W^{\mu
u}$$

with

$$W^{\mu
u}u_{
u}=0$$

- This is just math. No physics input except that ε is the energy density and u^μ is the velocity of the energy flow.
- Physics Small scale physics is thermal ⇒ Local equilibrium
 ⇒ Equation of state (i.e. P = P(ε))
 - $W^{\mu\nu} = (g^{\mu\nu} + u^{\mu}u^{\nu})P(\varepsilon) + \pi^{\mu\nu}[\varepsilon, u]$ with $\pi^{\mu\nu}u_{\nu} = 0$
 - Ideal Hydro: $\pi^{\mu\nu} = 0$ gives $\partial_t((\varepsilon + P)\mathbf{u}) = -\nabla P$ for small \mathbf{u}
 - Viscous Hydro:

$$\pi^{ij} = -\frac{\eta}{2} \left(\partial^{i} u^{j} + \partial^{j} u^{i} - g^{ij} (2/3) \nabla \cdot \mathbf{u} \right) - \zeta g^{ij} \nabla \cdot \mathbf{u}$$

- $\partial_{\mu} T^{\mu\nu} = 0$: This is an operator statement. This is valid no matter what.
- ∂_μ (*T^{μν}*) = 0: This is a statement about average. This is valid no matter what.

Ideal Hydro

- $T_{id.}^{\mu\nu} = \epsilon u^{\mu} u^{\nu} + P(g^{\mu\nu} + u^{\mu} u^{\nu})$ This assumes that the the system has isotropized \implies Ideal Hydrodynamics is valid only after the system has isotropize. But this is not enough.
- P(x) = P(ε(x)): Equation of state. Valid only if local equilibrium is reached. Recent most complete characterization of QCD thermalization process: 1107.5050 by Moore and Kurkela. t_{eq} ~ α⁻²Q⁻¹.

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Viscous Hydro

- $\pi_{ij} = -\eta \partial_{\langle i} u_{j \rangle}$ (tranceless, symmetric and transverse to u^{μ})
- Gradient expansion must be valid must
- This means local equilbrium is established in the length scale much longer than the microscopic mean free path.
- In fact, $\pi_{ij} = -\eta \partial_{\langle i} u_{j \rangle}$ induces unphysical faster-than-light propagations.

 \implies Second order Israel-Stewart formalism: π_{ij} relaxes towards $-\eta \partial_{\langle i} u_{j \rangle}$

$$\frac{d}{d\tau}\pi_{ij} = -\frac{1}{\tau_r}\left(\pi_{ij} - (-\eta\partial_{\langle i}u_{j\rangle})\right)$$

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Ideal Hydro

Stress-energy tensor

$$T^{\mu\nu} = \varepsilon u^{\mu} u^{\nu} + P(u^{\mu} u^{\nu} - g^{\mu\nu})$$

Energy momentum conservation

$$\mathbf{0} = \mathbf{u}^{\mu}\partial_{\mu}\varepsilon + (\varepsilon + \mathbf{P})(\partial_{\mu}\mathbf{u}^{\mu})$$

and

$$(\varepsilon + P)u^{\mu}\partial_{\mu}u_{\alpha} = \partial_{\alpha}P - u_{\alpha}u_{\nu}\partial^{\nu}P$$

One can easily show that entropy is conserved

$$\partial_{\mu}({\it su}^{\mu})=0$$

using $sT = \varepsilon + P$ and $TdS = dU + PdV \rightarrow Tds = d\varepsilon$

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Solving Hydro – Need for τ, η

• Idealized physical picture: Two infinitely energetic (v = c) pancakes pulling away from each other



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• Dynamic rapidity *y* is defined by:

$$E = \sqrt{m^2 + p_T^2} \cosh(y)$$

 $p_z = \sqrt{m^2 + p_T^2} \cosh(y)$

- Ends with $\pm c$ are at $y = \pm \infty$ \implies The system occupies the whole rapidity axis.
- With γ = cosh Δy and γν = sinh Δy, Lorentz boost is just a translation in the rapidity space

$$E' = \gamma E + \gamma v p_z = m_T \cosh(y + \Delta y)$$
$$p'_z = \gamma p_z + \gamma v E = m_T \sinh(y + \Delta y)$$

The system must be homogeneous in y —> Independent of y

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• Space-time rapidity η defined by

 $t = \tau \cosh \eta$ $z = \tau \sinh \eta$

• Lorentz boost is just a translation in the rapidity space

$$t' = \gamma t + \gamma vz = \tau \cosh(\eta + \Delta y)$$

$$z' = \gamma z + \gamma vt = \tau \sinh(\eta + \Delta y)$$

• A boost invariant system is independent of η as well.

Solving Hydro – 1+1 D (Bjorken)

- Simplify some more No dependence on *x*, *y*. No dissipation.
- The only thing a boost can do: Lorentz transform the fluid velocity u^{μ} .
- Boost invariance: Fluid velocity can only be $u^{\mu} = (t/\tau, 0, 0, z/\tau) = (\cosh \eta, 0, 0, \sinh \eta).$
- Let

$$arepsilon = arepsilon(au) \ {m P} = {m P}(au)$$

The energy-momentum conservation becomes

$$\frac{d\varepsilon}{d\tau} = -\frac{\varepsilon + P}{\tau}$$

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Bjorken 1+1 D - cont

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$$\frac{d\varepsilon}{d\tau} = -\frac{\varepsilon + P}{\tau}$$
 can be rewritten as
 $\frac{ds}{d\tau} = -\frac{s}{\tau}$
using $Ts = \varepsilon + P$ and $TdS = dU + PdV$.
Also,

$$\frac{d\varepsilon}{d\tau} = -\frac{(1+v_s^2)\varepsilon}{\tau}$$

using
$$v_s^2 = \frac{\partial P}{\partial \varepsilon}$$

Solutions

$$\boldsymbol{s}(\tau) = \boldsymbol{s}_0\left(\frac{\tau_0}{\tau}\right)$$

and

$$\varepsilon(\tau) = \varepsilon_0 \left(\frac{\tau_0}{\tau}\right)^{1+v_s^2}$$

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- At RHIC: $y_{\rm max} \approx \pm 5.4$
- At LHC: $y_{max} \approx \pm 8.0$
- Not $\pm\infty$, but big enough
- More technical reason: Hard to contain this system in t z as the boundary of the system linearly increases with time
- In τ , η , $\eta_{max} > y_{max}$ is enough.
- Price to pay: $\partial_{\mu}T^{\mu\nu} = 0$ becomes complicated.

- Generalized Israel-Stewart
- For example, shear viscosity: Baier, Romatschke, Son, Starinets, Stephanov (0712.2451)

$$\Delta^{\mulpha}\Delta^{
ueta} D\pi_{lphaeta} = -rac{1}{ au_{\pi}}\left(\pi^{\mu
u}-2\eta
abla^{\langle\mu}\,u^{\mu
angle}+rac{4}{3} au_{\pi}\pi^{\mu
u}(\partial_{lpha}u^{lpha})
ight)$$

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Physics Issue 1: Initial state

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Reminder

 What we want to do: Study how initial state spatial anisotropy

$$\varepsilon(\mathbf{x}_T,\eta) = \varepsilon(r_T,\eta) \left(1 + \sum_{n=1}^{\infty} 2\epsilon_n(r_T,\eta) \cos(n\phi)\right)$$

turns into the final state momentum anisotropy

$$\frac{dN_i}{dy \, d^2 p_T} = \frac{dN_i}{2\pi p_T dp_T dy} \left(1 + \sum_{n=1}^{\infty} 2v_{i,n}(p_T, y) \cos(n\phi)\right)$$

Wants to get history of physical quantities *P*, **u**, η/s, · · · from the flow coefficients v_{i,n}(p_T) – Need many different measurements



Thickness function

$$T_A(\mathbf{s}) = \int dz \, \rho_A(\mathbf{s}, z)$$

• Overlap function:

 $T_{AB}(\mathbf{s},\mathbf{b}) = T_A(\mathbf{s})T_B(\mathbf{b}+\mathbf{s})$

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- Participants: $N_{\text{part}}(\mathbf{s}, \mathbf{b}) \propto T_A(\mathbf{s}) + T_B(\mathbf{b} + \mathbf{s})$
- Binary scatterings: $N_{\rm bin}({\bf s}, {\bf b}) \propto T_{AB}({\bf s}, {\bf b})$ ٠
- Initial energy density ٠

 $\varepsilon(\mathbf{s}, \mathbf{b}) = c_1 \left[T_A(\mathbf{s}) + T_B(\mathbf{b} + \mathbf{s}) \right] + c_2 T_{AB}(\mathbf{s}, \mathbf{b})$

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- Ultimately, initial geometry determines the initial conditions and the final flow pattern.
- Initial geometry also determines number of jets at s and the path conditions for those jets.



 Smooth initial states have up-down, left-right symmetry: Initial states only has cos(2nφ) components such as v₂, v₄, v₆, · · ·

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What determines the initial shape?



• Averaged smooth initial condition \implies Only v_{even} 's survive.

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What determines the initial shape?



• Fluctuating initial condition \implies All v_n are non-zero.

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- $v_{odd} \neq 0$ due to fluctuations is obvious *once* you see it [Alver and Roland, Phys.Rev.C81:054905, 2010]
- v₂ and v₃ are sensitive to the different features of the initial condition
- Elliptic flow: Sensitive to the overall almond shape
- Triangular flow: Less so. More local in the sense that average initial condition gives zero *v*₃.
- Viscosity effect on different features is different
 - Viscosity smears out lumps.
 - Viscosity reduces differential flow Triangle is "rounder" than ellipse

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Effect of viscosity



The relative velocity of the two layers does not change.



The velocities eventually become the same.

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Effect of viscosity



- η = 0 means u₁ < u₂ < u₃ is maintained for a long time
- η ≠ 0 means that u₁ ≃ u₂ ≃ u₃ is achieved more quickly
- Shear viscosity smears out flow differences (it's a diffusion)
- Shear Viscosity reduces non-sphericity



This causes elliptic flow. It is harder to destroy this than


this (*v*₃) ...

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or this (*v*₄) ...

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or this (v_{10}) ...

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Differences in models:

- Position of the energy deposite (collision sites)
- Energy deposit at each collision sites $(xN_{part} + (1 x)N_{coll})$
- Size and spread of the initial lump

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Collision Geometry



- b: Impact parameter. Vector between two centers in transverse space
- r₁: Position vector from the center of the target nucleus
- $\mathbf{b} + \mathbf{r}_{\perp}$: Position vector from the center of the projectile nucleus



MC-Glauber

Sample Wood-Saxon thickness function

$$T_A(\mathbf{r}_{\perp}) = \int dz \, \frac{\rho_0}{1 + e^{(R-r)/a}}$$
$$T_A(\mathbf{b} + \mathbf{r}_{\perp}) = \int dz' \, \frac{\rho_0}{1 + e^{(R-r')/a}}$$

for nucleon positions

- *NN* collision occurs if two nucleons are within $D = \sqrt{\sigma_{NN}/\pi}$
- For each wounded nucleon, deposite $\epsilon_0 e^{-(\mathbf{x}-\mathbf{x}_C)^2/2\sigma_0^2}$

[BFKL, JIMWLK, BK]



- Gluon distributions for protons for $Q^2 = 10 \text{ GeV}^2$ and $Q^2 = 100 \text{ GeV}^2$.
- Looks like growing indefinitely: Unphysical

[BFKL, JIMWLK, BK]

QCD – Interaction of quarks and gluons



- Leading order BFKL equation (evolution in *x*) takes into account splitting, but not recombination.
- When density is high, recombination must be taken into account
 JIMWLK & BK
- Density is high: Classical field limit
- Recombination: Non-linear effect

[BFKL, JIMWLK, BK]

- Saturation (or Recombination) scale
 - Transverse gluon density

$$ho \sim rac{xg_{\mathcal{A}}(x,Q^2)}{\mathcal{S}_{\perp}} \sim rac{\mathcal{A}xg(x,Q^2)}{\mathcal{A}^{2/3}} \sim \mathcal{A}^{1/3}xg(x,Q^2)$$

Recombination cross-section

$$\frac{\sigma_{gg \to g} \sim \alpha_s^2}{Q^2}$$

Saturation when

$$ho\sigma_{\gg
ightarrow g}\sim 1$$

Saturation scale

$$Q_s^2 = \alpha_s(Q_s) A^{1/3} x g(x, Q_s^2)$$

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Classical field equation of QCD

 $D_\mu G^{\mu
u} = J^
u$

where

$$D_{\mu} = \partial_{\mu} - igA^a_{\mu}T_a$$

and

$$G^{a}_{\mu
u} = \partial_{\mu}A^{a}_{
u} - \partial_{
u}A^{a}_{\mu} + gf_{abc}A^{b}_{\mu}A^{c}_{
u}$$

• $J_a^{\mu} = \rho_A \delta^{\mu +} + \rho_B \delta^{\mu -}$: Color source

Gluon field

$$m{A}^{\mu}=m{A}^{\mu}_{m{A}}+m{A}^{\mu}_{m{B}}+m{A}^{\mu}_{m{P}} heta(au)$$

The produced field A_P after the collision is what we are after

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MC-KLN Drescher and Nara, PRC 75:034905

Sample Wood-Saxon thickness function

$$T_{\mathcal{A}}(\mathbf{r}_{\perp}) = \int dz \, \frac{\rho_0}{1 + e^{(R-r)/a}}$$
$$T_{\mathcal{A}}(\mathbf{b} + \mathbf{r}_{\perp}) = \int dz' \, \frac{\rho_0}{1 + e^{(R-r')/a}}$$

for nucleon positions



MC-KLN Drescher and Nara, PRC 75:034905

Calculate thickness function again:



where $S = \sigma_{NN}$ is the cross-section of the tube



MC-KLN Drescher and Nara, PRC 75:034905

• Calculate the saturation scale

$$Q_{s,\mathcal{A}}^2(x,\mathbf{r}_{\perp}) = 2\,\mathrm{GeV}^2\left(\frac{t_{\mathcal{A}}(\mathbf{r}_{\perp})}{1.53}\right)\left(\frac{0.01}{x}\right)^{\lambda}$$

• Calculate the unintegrated gluon density function

$$\phi(\mathbf{x}, \mathbf{k}_{\perp}^2; \mathbf{r}_{\perp}) = \frac{1}{\alpha_s(\mathbf{Q}_s^2)} \frac{\mathbf{Q}_s^2}{\max(\mathbf{Q}_s^2, \mathbf{k}_{\perp}^2)}$$

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MC-KLN Drescher and Nara, PRC 75:034905

• Deposite energy with an approximate the $g_A g_B \rightarrow g_P$ process

$$\frac{dE_g}{d^2 \mathbf{r}_\perp dy d^2 p_\perp} = \frac{4N_c}{N_c^2 - 1} \frac{1}{|\mathbf{p}_\perp|} \int d^2 k_\perp \alpha_s \phi_A((\mathbf{p}_\perp + \mathbf{k}_\perp)^2/4) \phi_A((\mathbf{p}_\perp - \mathbf{k}_\perp)^2/4)$$

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[Tribedy & Venugopalan, Nucl.Phys.A850 136; Tribedy & Venugopalan, PLB710 125; Schenke, Tribedy & Venugopalan, PRL108 252301; Gale, Jeon, Schenke, Tribedy & Venugopalan, PRL110 012302]



IP-Glasma

- Sample the position of the nucleons.
- Calcualte the saturation momentum for each nucleon using the IP-Sat model (Kowalski & Teaney, PRD68 114005)
- Calculate the color charge density by summing over all *Q_s* at the given global position

$$g\mu(x,\mathbf{b}) = c\sum_{i} Q_s(x,\mathbf{b}_i)$$

[Tribedy & Venugopalan, Nucl.Phys.A850 136; Tribedy & Venugopalan, PLB710 125; Schenke, Tribedy & Venugopalan, PRL108 252301; Gale, Jeon, Schenke, Tribedy & Venugopalan, PRL110 012302]



IP-Glasma

• Sample the color charge distribution of each nucleus using the Gaussian distribution

$$W_{A}[
ho] = \exp\left(-
ho_{a}
ho_{b}/(g^{2}\mu_{A}^{2})
ight)$$

- Solve the Classical Yang-Mills equation
- After evolving for au_0 calculate $T^{\mu
 u}$
- Connect it to Hydro

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Comparison

[Schenke, Tribedy & Venugopalan, PRL108 252301; Schenke, Gale & Jeon, arXiv:1301.5893v1]



- Different size and distribution of ϵ_n
- Test: Need to get the $v_n(p_T)$ for various centralities
- Test: Need to get the e-by-e distribution of integrated v_n

Physics Issue 2: Viscosity

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Effect of viscosity



The relative velocity of the two layers does not change.



The velocities eventually become the same.

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Effect of viscosity



- η = 0 means u₁ < u₂ < u₃ is maintained for a long time
- η ≠ 0 means that u₁ ≃ u₂ ≃ u₃ is achieved more quickly
- Shear viscosity smears out flow differences (it's a diffusion)
- Shear Viscosity reduces non-sphericity

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Interaction Strength and Viscosity

Weak coupling allows rapid momentum diffusion



Large η/s : $u_{\mu}(x)$ changes due to pressure gradient and diffusion

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Interaction Strength and Viscosity

Strong coupling *does not* allow momentum diffusion



Small η/s : $u_{\mu}(x)$ changes due to pressure gradient only

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Kinetic Theory estimate



• Rough estimate (fluid rest frame, or $u_z(x) = 0$)

The momentum density: T_{0z} = (ε + P)u₀u_z diffuses in the x direction with v_x = u_x/u₀. Net change:

$$\begin{array}{l} \langle \epsilon + \mathcal{P} \rangle \left| v_{x} \right| u_{0} (u_{z} (x - l_{\mathrm{mfp}}) - u_{z} (x + l_{\mathrm{mfp}})) \\ \approx -2 \left\langle \epsilon + \mathcal{P} \right\rangle \left| v_{x} \right| u_{0} l_{\mathrm{mfp}} \partial_{x} u_{z} (x) \\ \sim -\eta u_{0} \partial_{x} u_{z} \end{array}$$

Here I_{mfp} : Mean free path

• Recall thermo. id.: $\langle \epsilon + \mathcal{P} \rangle = \mathbf{s} \mathbf{T}$

$$\eta \sim \langle \epsilon + \mathcal{P} \rangle \ \textit{I}_{mfp} \ \langle |\textit{v}_{\textit{x}}| \rangle \sim$$
 s *T* $\textit{I}_{mfp} \ \langle |\textit{v}_{\textit{x}}| \rangle$

Perturbative estimate

High Temperature limit: $\langle |v_x| \rangle = O(1)$ • $\eta/s \approx T I_{mfp} \approx \frac{T}{n\sigma} \sim \frac{1}{T^2\sigma}$ • The only energy scale: T

$$\sigma \sim \frac{(\text{coupling constant})^{\#}}{T^2}$$

Hence

$$rac{\eta}{s} \sim rac{1}{(ext{coupling constant})^{\#}}$$

• Perturbative QCD partonic 2-2 cross-section

$$\frac{d\sigma_{\rm el}}{dt} = C \frac{2\pi\alpha_{\rm S}^2}{t^2} \left(1 + \frac{u^2}{s^2}\right)$$

Naively expect

$$\eta/\mathrm{S}\sim \frac{1}{lpha_{s}^{2}}$$

• Coulomb enhancement (cut-off by m_D) leads to

$$\eta/\mathrm{s} \sim rac{1}{lpha_s^2 \ln(1/lpha_s)}$$

QCD η calc

Relevant processes



Use kinetic theory

$$\frac{df}{dt} = \mathcal{C}_{2\leftrightarrow 2} + \mathcal{C}_{1\leftrightarrow 2}$$

Complication: 1 \leftrightarrow 2 process needs resummation (LPM effect, AMY)

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QCD Estimates of η/s

- Danielewicz and Gyulassy [PRD 31, 53 (1985)]:
 - η/s bound from the kinetic theory: Recall: $\eta \sim s T I_{mfp} \langle |v_x| \rangle$ Use $I_{mfp} \langle |v_x| \rangle \sim \Delta x \Delta p/m$ to get

$$rac{\eta}{s} \gtrsim rac{1}{12} pprox 0.08 pprox (1/4\pi)$$

• QCD estimate in the small α_S limit with $N_f = 2$ and $2 \rightarrow 2$ only (min. at $\alpha_S = 0.6$):

$$\eta pprox rac{T}{\sigma_\eta} pprox rac{0.57 T^3}{lpha_S^2 \ln(1/lpha_S)} \gtrsim rac{0.2 s}{lpha_S} pprox (2.5/4\pi) s$$

Baym, Monien, Pethick and Ravenhall [PRL 64, 1867 (1990)]

$$\eta pprox rac{1.16 T^3}{lpha_S^2 \ln(1/lpha_S)} \gtrsim 0.4 s pprox (5/4\pi) s$$

• M. Thoma [PLB 269, 144 (1991)]

$$\eta pprox rac{1.02 \, T^3}{lpha_S^2 \ln(1/lpha_S)} \gtrsim 0.4 s pprox (5/4\pi) s$$

Full leading order calculation of η/s

• Arnold-Moore-Yaffe (JHEP 0305, 051 (2003)) [Plots: Guy]:



Shear viscosity in $\mathcal{N}=4$ SYM

Son, Starinets, Policastro, Kovtun, Buchel, Liu, ...

- Strong coupling limit, 4 ingredients
 - Kubo formula

$$\eta = \lim_{\omega \to 0} \frac{1}{2\omega} \int dt \, d^3x \, e^{i\omega t} \, \langle [T_{xy}(x), T_{xy}(0)] \rangle$$

Gauge-Gravity duality

$$\sigma_{\rm abs}(\omega) = rac{8\pi G}{\omega} \int dt \, d^3x \, e^{i\omega t} \, \langle [T_{xy}(x), T_{xy}(0)]
angle$$

- $\lim_{\omega \to 0} \sigma_{abs}(\omega) = A_{blackhole}$
- Entropy of the BH : $s = A_{\text{blackhole}}/4G$

Therefore, (including the first order correction)

$$\frac{\eta}{\rm s} = \frac{1}{4\pi} \left(1 + \frac{7.12}{(g^2 N_c)^{3/2}} \right)$$

Correction is small if $g \gg 1$ (10% at g = 2.4).

N = 4 SYM



Perturbative calculation and the strong coupling calculation behave very differently

Experimental Evidence for $\eta/s \sim 1/4\pi$

- Theoretical situation:
 - Perturbative calculations: $\eta/s \ge 7.5/(4\pi)$
 - AdS/CFT in the infinite coupling limit: $\eta/s = 1/(4\pi)$
 - Roughly an order of magnitude difference \implies Testable!
- A relativistic heavy ion collision produces a complicated system
 Need a hydrodynamics simulation suite
- We use MUSIC (3+1D e-by-e viscous hydrodynamics)
- Viscosity measurement is through the flow coefficients

$$\frac{dN}{dyd^2p_T} = \frac{dN}{2\pi dyp_T dp_T} \left(1 + 2\sum_{n=1}^{\infty} v_n \cos(n(\phi - \psi_n))\right)$$

• v_n is a translation of the eccentricities ϵ_n via pressure gradient

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Aride The Music

MUScl for Ion Collisions

MUSCL: Monotone Upstream-centered Schemes for Conservation Laws

Current MUSIC (and MARTINI) Team

- Charles Gale (McGill)
- Sangyong Jeon (McGill)
- Björn Schenke (Formerly McGill, now BNL)
- Clint Young (Formerly McGill, now UMN)
- Gabriel Denicol (McGill)
- Matt Luzum (McGill/LBL)
- Sangwook Ryu (McGill)
- Gojko Vujanovic (McGill)
- Jean-Francois Paquet (McGill)
- Michael Richard (McGill)
- Igor Kozlov (McGill)

MUSIC

3+1D Event-by-Event Viscous Hydrodynamics

- 3+1D parallel implementation of new *Kurganov-Tadmor Scheme* in (τ, η) with an additional baryon current (No need for a Riemann Solver. Semi-discrete method.)
- Ideal and Viscous Hydro
- Event-by-Event fluctuating initial condition
- Sophisticated Freeze-out surface construction
- Full resonance decay (3+1D version of Kolb and Heinz)
- Many different equation of states including the newest from Huovinen and Petreczky
- New Development: Glasma Initial Conditions & UrQMD after-burner

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Fluctuating Initial Condition

Each event is *not* symmetric: Fluctuating initial condition \implies All v_n are non-zero.







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Ideal vs. Viscous

[Movies by B. Schenke]



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Ideal vs. Viscous

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- Magnitude of higher harmonics, v₃, v₄, ···, (almost) independent of centrality – Local fluctuations dominate
- Higher harmonics are easier to destroy that *v*₂ which is a *global* distortion Viscosity effect.
- To get a good handle on flow: Both fluctuations and viscosity are essential

E-by-E MUSIC vs LHC Data

[Schenke, Jeon and Gale, Phys. Rev. C 85, 024901 (2012)] Best value $\eta/s = 0.16 = 2/(4\pi)$.



Glasma Initial Condition

[Gale, Jeon, Schenke, Tribedy and Venugopalan, arXiv:1209.6330] Best value $\eta/s = 0.2 = 2.5/(4\pi)$.



Jeon (McGill)

Glasma Initial Condition

[Gale, Jeon, Schenke, Tribedy and Venugopalan, arXiv:1209.6330] E-by-E distributions



Jeon (McGill)

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New Development: UrQMD Afterburner

v_2 at RHIC (Midrapidity). In each centrality class: 100 UrQMD times 100 MUSIC events. [Ryu, Jeon, Gale, Schenke and Young, arXiv:1210.4558]





 Using previous MUSIC parameters that were tuned to reproduce PHENIX v_n

LHC Spectra

In each centrality class: 100 MUSIC times 10 UrQMD events. $\eta/s = 2/(4\pi)$. ALICE data from QM12.



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LHC Flows

In each centrality class: 100 MUSIC times 10 UrQMD events



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- Strong flows: Strongest evidence that η/s has to be small
- η/s much larger than 0.2 cannot be accommodated within current understanding of the system.
- Perturbative result of $\eta/s = 0.4 0.6$ is out.
- Using the LQCD EoS.
- LQCD estimate (η + 3ζ/4)/s ≈ 0.20 0.26 between 1.58T_c 2.32T_c.
 [H. Meyer, Eur.Phys.J.A47:86,2011]
- Does this mean very large coupling?

Jet Quenching

• Fact: Jets lose energy (ATLAS images).



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Jet Quenching

• Fact: Jets lose energy (ATLAS images).





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 Collisional energy loss rate [Wicks, Horowitz, Djordjevic and Gyulassy, NPA 784, 426 (2007), Qin, Gale, Moore, Jeon and Ruppert, Eur. Phys. J. C 61, 819 (2009)]:

$$\frac{dE}{dx} \approx C_1 \pi \alpha_S^2 T^2 \left[\log \left(\frac{E_p}{\alpha_S T} \right) + C_2 \right]$$

 $C_{1,2}$: Depends on the process. O(1).

• Radiational $\propto \alpha_s^2$ (Arnold, Moore, Yaffe, JHEP 0206, 030 (2002))



What we want to get at

What α_S do we need for these?



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- Event generator
 - Jet propagation through evolving QGP medium.
- Several on the market. We use MARTINI.

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- Modular Alogorithm for Relativistic Treatment of Heavy IoN Interactions
- Hybrid approach
 - Calculate Hydrodynamic evolution of the soft mode (MUSIC)
 - Propagate jets in the evolving medium according to McGill-AMY

Heavy Ion Collisions



HIC Jet production scheme:

 $\frac{d\sigma_{AB}}{dt} = \int_{\text{geometry}} \int_{abcdc'} \times f_{a/A}(x_a, Q_f) f_{b/B}(x_b, Q_f) \\ \times \frac{d\sigma_{ab \to cd}}{dt} \\ \times \mathcal{P}(x_c \to x'_c | T, u^{\mu}) \\ \times D(z'_c, Q)$

 $\mathcal{P}(x_c \rightarrow x'_c | T, u^{\mu})$: Medium modification of high energy parton property

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MARTINI - Basic Idea



$$\frac{d\sigma_{AB}}{dt} = \int_{\text{geometry}} \int_{abcdc'} \times f_{a/A}(x_a, Q_f) f_{b/B}(x_b, Q_f) \\ \times \frac{d\sigma_{ab \to cd}}{dt} \\ \times \mathcal{P}(x_c \to x'_c | T, u^{\mu}) \\ \times D(z'_c, Q)$$

 Sample collision geometry using Wood-Saxon

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MARTINI - Basic Idea



$$\frac{d\sigma_{AB}}{dt} = \int_{\text{geometry}} \int_{abcdc'} \times f_{a/A}(x_a, Q_f) f_{b/B}(x_b, Q_f) \\ \times \frac{d\sigma_{ab \to cd}}{dt} \\ \times \mathcal{P}(x_c \to x'_c | T, u^{\mu}) \\ \times D(z'_c, Q)$$

- PYTHIA 8.1 generates high *p_T* partons
- Shadowing included
- Shower (Radiation) stops at $Q = \sqrt{p_T/\tau_0}$

MARTINI - Basic Idea



$$\frac{d\sigma_{AB}}{dt} = \int_{\text{geometry}} \int_{abcdc'} \times f_{a/A}(x_a, Q_f) f_{b/B}(x_b, Q_f) \\ \times \frac{d\sigma_{ab \to cd}}{dt} \\ \times \mathcal{P}(x_c \to x'_c | T, u^{\mu}) \\ \times D(z'_c, Q)$$

- Hydrodynamic phase (MUSIC)
- AMY evolution MC simulation of the rate equ's.

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Parton propagation

Process include in MARTINI (all of them can be switched on & off):



Photon: emission & conversion

Jeon (McGill

Parton propagation

Resummation for the inelastic processes included:



- All such graphs are leading order (BDMPS)
- Full leading order SD-Eq (AMY): (Figure from G. Qin)



Parton propagation



An example path in MARTINI. (Figure from B. Schenke)

Jeon (M

• While this is happening in the background ...

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Projection on to the longitudinal plane

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Projection onto the transverse plane

Jeon (McGill)

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Pion production

[Schenke, Jeon and Gale, Phys. Rev. C 80, 054913 (2009)]



- E - N

Photon production

[Schenke, Jeon and Gale, Phys. Rev. C 80, 054913 (2009)]

• Spectra and R_{AA}^{γ}



• $R_{AA}(p_T, \Delta \phi)$



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MARTINI – LHC dN/dA

[Young, Schenke, Jeon, Gale, Phys. Rev. C 84, 024907 (2011)].

•
$$A = (E_t - E_a)/(E_t + E_a)$$

- This is with ideal hydro with a smooth initial condition
- Full jet reconstruction with FASTJET
- $\alpha_S = 0.27$ seems to work.



ATLAS, PRL 105 (2010) 252303



CMS, arXiv: 1102.1957 (2011)

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MARTINI – LHC dN/dA

[Young, Schenke, Jeon, Gale, Phys. Rev. C 84, 024907 (2011)].

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CMS, arXiv: 1102.1957 (2011)

ATLAS, QM 2011

Not the full story

[Clint Young's HP2012 Proceedings]



- R_{AA} For LHC, constant α_S suppresses jets too much.
- Need to incorporate finite length effect (Caron-Huot-Gale) and running α_s . This is with maximum $\alpha_s = 0.27$.
- Don't quite get azimuthal dependence yet. Δφ broadening may be due to the background fluctuations ⇒ Need to combine UrQMD background?

- So many nuclear experiments are being done/planned. RHIC, LHC, Raon, FRIB, FAIR, JPARC, Dubna, HIRFL-CSR ...
- There never have been a time in history when so much information is so readily available.
- This is a great time to be/become a nuclear physicist.
- Work hard. Think hard. Dream big.

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