Computing Atomic Nuclei

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Introduction



Weinberg's Laws of Progress in Theoretical Physics From: "Asymptotic Realms of Physics" (ed. by Guth, Huang, Jaffe, MIT Press, 1983)

First Law: "The conservation of Information" (*You will get nowhere by churning equations*)

Second Law: "Do not trust arguments based on the lowest order of perturbation theory"

Third Law: "You may use any degrees of freedom you like to describe a physical system, but if you use the wrong ones, you'll be sorry!"



Patient: Doctor, doctor, it hurts when I do this! Doctor: Then don't do that.

Nuclear Structure Theory Progress Report



Low-lying Hadron Spectrum Dürr, Fodor, Lippert et al., BMW Collaboration Science 322, 1224 November 2008

More than 99% of the mass of the visible universe is made up of protons and neutrons. Both particles are much heavier than their quark and gluon constituents, and the Standard Model of particle physics should explain this difference. We present a full ab initio calculation of the masses of protons, neutrons, and other light hadrons, using lattice quantum chromodynamics. Pion masses down to 190 mega–electron volts are used to extrapolate to the physical point, with lattice sizes of approximately four times the inverse pion mass. Three lattice spacings are used for a continuum extrapolation. Our results completely agree with experimental observations and represent a quantitative confirmation of this aspect of the Standard Model with fully controlled uncertainties



Lattice QCD calculation of nuclear force



N. Ishii, S. Aoki, T. Hatsuda, Phys. Rev. Lett. **99,** 022001 (2007) Tensor force from LQCD: http://arxiv.org/pdf/0903.5497



number of nuclei < number of processors!

Ab initio theory for light nuclei and nuclear matter

Ab initio: GFMC, NCSM, CCM (nuclei, neutron droplets, nuclear matter)



NN and NNN interactions



N³LO: Entem et al., PRC68, 041001 (2003) Epelbaum, Meissner, et al. V_{low-k} unifies NN interactions at low energy



Bogner, Kuo, Schwenk, Phys. Rep. 386, 1 (2003)

- Quality two- and three-nucleon interactions exist
 - Not uniquely defined (local, nonlocal)
 - Soft and hard-core



1-2% calculations of A = 6 - 12 nuclear energies are possible excited states with the same quantum numbers computed

Strongly paired fermions: Cold atoms and neutron matter



Gezerlis and Carlson, Phys. Rev. C 77, 032801(R) (2008)

Nuclear Coupled Cluster Theory Size Extensive!

Medium-mass nuclei from chiral nucleon-nucleon interactions

Hagen, Papenbrock, Dean, Hjorth-Jensen, Phys. Rev. Lett. 101, 092502 (2008)







Mean-Field Theory \Rightarrow Density Functional Theory



Nuclear DFT

- two fermi liquids
- self-bound
- superfluid
- mean-field \Rightarrow one-body densities
- zero-range \Rightarrow local densities
- finite-range \Rightarrow gradient terms
- particle-hole and pairing channels
- Has been extremely successful. A broken-symmetry generalized product state does surprisingly good job for nuclei.

Nuclear Energy Density Functional

isoscalar (T=0) density 0 n p
isovector (T=1) density 1 n p

+isoscalar and isovector densities: spin, current, spin-current tensor, kinetic, and kinetic-spin

^p + pairing densities

$$E = \int \mathcal{H}(r) d^3r$$

 $\mathcal{H}(r) = \frac{\hbar^2}{2m} \tau_0(r) + \sum_{t=0,1}^{\text{p-h density } p-p \text{ density (pairing functional)}} (\chi_t(r) + \breve{\chi}_t(r))$ Expansion in densities and their derivatives

- Constrained by microscopic theory: ab-initio functionals provide quasi-data!
- Not all terms are equally important. Usually ~12 terms considered
- Some terms probe specific experimental data
- Pairing functional poorly determined. Usually 1-2 terms active.
- Becomes very simple in limiting cases (e.g., unitary limit)

Nuclear DFT: works well for BE differences



Neutron-rich matter and neutron skins



Microscopic mass table





 New collective modes; polarization effects

- Nuclei are open quantum systems
- Coupling to the continuum important Virtual scattering
 - •Unbound states
 - Impact on in-medium Interactions

Wikipedia:

An open quantum system is a quantum system which is found to be in interaction with an external quantum system, the environment. The open quantum system can be viewed as a distinguished part of a larger closed quantum system, the other part being the environment.



Prog. Part. Nucl. Phys. 59, 432 (2007)



Neutron number —

Halos ²H (deuteron) Riisager, Fedorov, Jensen $S_n=2.2$ MeV, $r_{np}=4$ fm Europhys. Lett. 49, 547 (2000) ⁴He₂ (atomic helium dimer) S=0.13 µeV, r=100 Å ³H_L (hypertriton) S_{Λ} =0.08 MeV 10^{2} .⁴He₂ 11 = 09 7 <r3>/R² 0 (fm) 5 3 1 10 15 20 5 l = 2(fm) Cobis, Jensen, Fedorov 1 o ۴B ¹⁹C J. Phys. G23, 401 (1997) 10⁻² 10-1 1 μ B R²/h²

Rigged Hilbert Space: the natural framework to formulate quantum mechanics

In mathematics, a rigged Hilbert space (Gel'fand triple, nested Hilbert space, equipped Hilbert space) is a construction designed to link the distribution and square-integrable aspects of functional analysis. Such spaces were introduced to study spectral theory in the broad sense. They can bring together the 'bound state' (eigenvector) and 'continuous spectrum', in one place.

Mathematical foundations in the 1960s by Gel'fand et al. who combined Hilbert space with the theory of distributions. Hence, the RHS, rather than the Hilbert space alone, is the natural mathematical setting of Quantum Mechanics

I. M. Gel'fand and N. J. Vilenkin. Generalized Functions, vol. 4: Some Applications of Harmonic Analysis. Rigged Hilbert Spaces. Academic Press, New York, 1964.

$$\psi = \psi(r)e^{-iE_0t/\hbar - wt/2} = \psi(r)e^{-iEt/\hbar}$$

 $E = E_0 - i\frac{\Gamma}{2}; \quad \Gamma = \hbar w$
J.J. Thompson, 1884
G. Gamow, 1928 relation between decay width
and decay probability

J. Phys. G: Nucl. Part. Phys. 36 (2009) 013101 (40pp)

doi:10.1088/0954-3899/36/1/013101

TOPICAL REVIEW

Shell model in the complex energy plane

N Michel^{1,2}, W Nazarewicz^{3,4,5}, M Płoszajczak⁶ and T Vertse^{7,8}

Abstract

This work reviews foundations and applications of the complex-energy continuum shell model that provides a consistent many-body description of bound states, resonances and scattering states. The model can be considered a quasi-stationary open quantum system extension of the standard configuration interaction approach for well-bound (closed) systems.

Complex-energy Shell Model Gamow Shell Model

GSM: N. Michel et al., Phys.Rev.Lett. 89, 042502 (2002) 8 ⁹He ¹⁰He ⁵He ⁶He ⁷He ⁸He 7 6 finite-range SDI GSM EXP x 10 HF-basis 5 0 +s.p. basis: ⁵He unbound! energy (MeV) Δ p3/2 0.75 -i0.3 MeV 3 p_{1/2} 2.13 -i2.9 MeV 2 5/2-1/2-1 2+2+3/2-1/2.0 3/2--1 0 +<u>?</u> 1/2--2 0+ **Helium isotopes** -3 · 0+ -4

Real-energy Continuum Shell Model A. Volya and V. Zelevinsky, Phys. Rev. C 67 (2003) 54322



Connections to quantum many-body systems



- Understanding the transition from microscopic to mesoscopic to macroscopic
- Symmetry breaking and emergent phenomena
 - Pairing in finite systems
- Quantum chaos
- Open quantum systems
- Dynamical symmetries and collective dynamics



Dilute fermion matter:

- strongly correlated very large scattering length (unitary limit)
- Low-density neutron matter
- Cold fermions in traps



Trapped ⁶Li gas (Rice)



-1/2K

Computational Strategy

Connections to computational science

1Teraflop=10¹² flops 1peta=10¹⁵ flops (next 2-3 years) 1exa=10¹⁸ flops (next 10 years)



Million-fold increase in computing and data capabilities (ORNL)



Scientific Grand Challenges Workshop Series

Enabling science communities to address scientific grand challenges through extreme scale computational science

Workshop series:

- •Climate Science
- High-Energy Physics
- Nuclear Physics
- Fusion Energy Sciences
- Nuclear Energy
- Biology
- Materials Science and Chemistry



26-28 January 2009, Washington, DC 109 participants; DOE/NSF/NNSA reps

The Nuclear Physics Workshop defined Priority Research Directions in

- Nuclear Astrophysics
- Cold QCD and Nuclear Forces
- Nuclear Structure and Reactions
- Accelerator Physics
- Hot and Dense QCD



Nuclear Structure and Nuclear Reactions

Forefront Questions in Nuclear Science and the Role of High Performance Computing January 26-28, 2009 · Washington, D.C.



List of Priority Research Directions

- Physics of extreme neutron-rich nuclei and matter
- Microscopic description of nuclear fission
- Nuclei as neutrino physics laboratories
- Reactions that made us triple α process and ¹²C(α,γ)¹⁶O



Nuclear Physics Requires Exascale Computing

Nuclear Astrophysics	stellar : 3D turbulance 3D SN progenitors 3D SN , neut mixing,. 3D core-coll. SN whole star 3D SN la turbulant nuclear burn 3D SN la whole star
Cold QCD and Nuclear Forces	g_A to 3%Nucleon Spin, Parton Distsflavor-GPD'sNNN-ints α Low-Lying Hadron SpectrumDeuteron $\langle d j^{\mu}_{wk} d\rangle$ Excited Hadron Spectrum
Nuclear Structure	Light Ion Reactions t-dep. Fission, ab initio Ni isotopes Sn Fusion in $etaeta$ - rates $3lpha$ capture Sn Med. Nuclei ${}^{12}C(lpha,\gamma){}^{16}O$
Accelerator Physics	Isotope separator ECR ion src ERL Heating of cryo's optimization e- cooling of H.I. in ERL
Hot and Dense	transport in QCD bulk thermo (quenched) (staggered) From QCD to detector Phase structure phase-diagram mu(B)/T < 3 transport
1	0 ⁻³ 10 ⁻² 10 ⁻¹ 1 10 10 ² 10 ³ 1 Pflop-Yrs on Task (sustained)



SciDAC 2 Project: Building a Universal Nuclear Energy Density Functional

- Understand nuclear properties "for element formation, for properties of stars, and for present and future energy and defense applications"
- Scope is all nuclei, with particular interest in reliable calculations of unstable nuclei and in reactions
- Order of magnitude improvement over present capabilities
 Precision calculations
- Connected to the best microscopic physics
- Maximum predictive power with well-quantified uncertainties

Universal Nuclear Energy Density Functional



•Funded (on a competitive basis) by •Office of Science •ASCR

- •NNSA
- 15 institutions
- ~50 researchers
 •physics
 •computer science
 •applied mathematics
- foreign
 collaborators
- 5 years http://unedf.org/

...unprecedented theoretical effort !

[See http://www.scidacreview.org/0704/html/unedf.html by Bertsch, Dean, and Nazarewicz]

Ab-initio nuclear structure: towards ${}^{12}C(\alpha,\gamma)$



In January 2009: calculations of ¹²C with a complete Hamiltonian (two- and three-nucleon potentials -- AV18+IL7) on 32,000 processors of the Argonne BGP. These are believed to be the best converged ab initio calculations of ¹²C ever made. The result is quite good; the computed binding energy is 93.5(6) MeV compared to the experimental value of 92.16 MeV and the point rms radius is 2.35 fm vs 2.33 from experiment. The figure compares the computed ¹²C density with that extracted from electron-scattering experiments. Note the good reproduction of the dip at small radius.

Example: Large Scale Mass Table Calculations

Science scales with processors

M. Stoitsov

HFB+LN mass table, HFBTHO

Even-Even Nuclei

- The SkM* mass table contains 2525 even-even nuclei
- A single processor calculates each nucleus 3 times (prolate, oblate, spherical) and records all nuclear characteristics and candidates for blocked calculations in the neighbors
- Using 2,525 processors about 4 CPU hours (1 CPU hour/configuration)

All Nuclei

- 9.210 nuclei
- 599,265 configurations
- Using 3,000 processors about 25 CPU hours

see MassExplorer.org







Multimodal fission in nuclear DFT

A. Staszczak, A.Baran,

J. Dobaczewski, W.N.



Neutron number N









Recent years: very successful period for theory of nuclei

- many new ideas leading to new understanding
- new theoretical frameworks
- exciting developments
- high-quality calculations



- Effective Field Theory/Renormalization Group provides missing links Short-range repulsion: a red herring!
- Accurate ab-initio methods allow for interaction tests
- Worldwide attack on the nuclear energy density functional
- Quantitative microscopic nuclear structure
- Integrating nuclear structure and reactions
- High-performance computing continues to revolutionize microscopic nuclear many-body problem: impossible becomes possible
- Some of the most interesting physics outcomes will be at the interfaces:
 - QCD to forces to structure
 - structure and reactions with nuclear astrophysics



- Exciting science; old paradigms revisited
- Interdisciplinary (quantum many-body problem, cosmos,...)
- Relevant to society (energy, medicine, national security, ...)
- Theory gives the mathematical formulation of our understanding and predictive ability
- New-generation computers provide unprecedented opportunities
- Large coherent international theory effort is needed to make a progress

Guided by data on short-lived nuclei, we are embarking on a comprehensive study of all nuclei based on the most accurate knowledge of the strong internucleon interaction, the most reliable theoretical approaches, and the massive use of the computer power available at this moment in time. The prospects look good.





(Nuclear) Many-Body Physics: "Old" vs. "New"

One Hamiltonian for all problems and energy/length scales	Infinite # of low-energy potentials; different resolutions => different dof's and Hamiltonians
Find the "best" potential	There is no best potential
Two-body data may be sufficient; many-body forces as last resort	Many-body data needed and many-body forces inevitable
Avoid (hide) divergences	Exploit divergences (cutoff dependence as tool)
Choose diagrams by "art"	Power counting determines diagrams and truncation error

Short-range correlations: a red herring

Dick Furnstahl	The UNEDF Project