



Nuclear Structure I

experimental

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Sunday

Preliminaries

Nuclear binding and masses

→ Indicator of shell structure

How to measure a mass

Monday

Thursday

Friday



Preliminaries (1)

Goal: Establish physical properties of rare isotopes and their interactions to gain predictive power

Experiments: Measure observables

Observables: May or may not need interpretation to relate to physical properties

- e.g., half-life and mass connect directly to physical properties
- e.g., cross sections for reaction processes usually need interpretation to connect to physical properties (model dependencies are introduced)



Preliminaries (2)

Theories and models can relate observables to physical properties – often, experiments are motivated by theoretical predictions that need validation

But: Theories and models have their own realm of applicability that everybody involved in the experiment/data analysis/interpretation should be aware of!

Predictions or systematics come with **a warning:** Might lead to expectations that can influence the implementation of an experiment and ultimately limit the scope of discovery

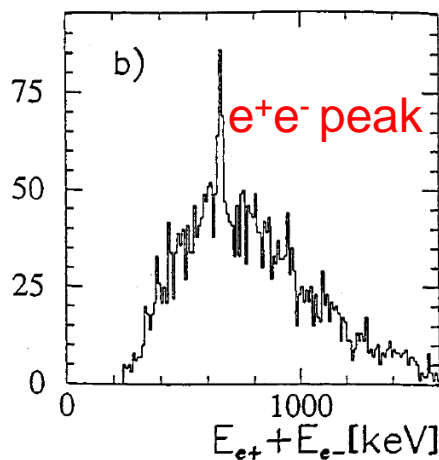
Implementation of experiments can limit the scope of discovery

Examples:

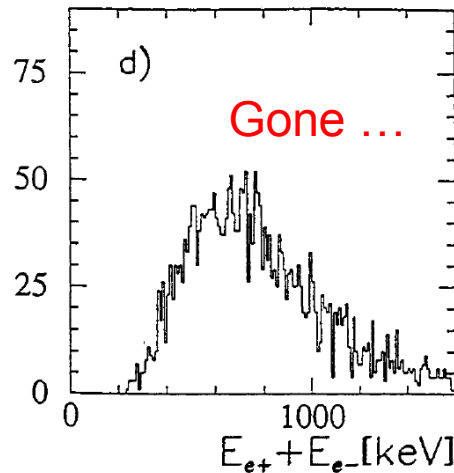
- Lifetime of a nucleus
- Excitation energy

Analysis and “selection” of data (cuts, gates, using subsets, ...) can influence the result

Example: e^+e^- resonances in heavy-ion collisions at GSI



First half of the data analyzed



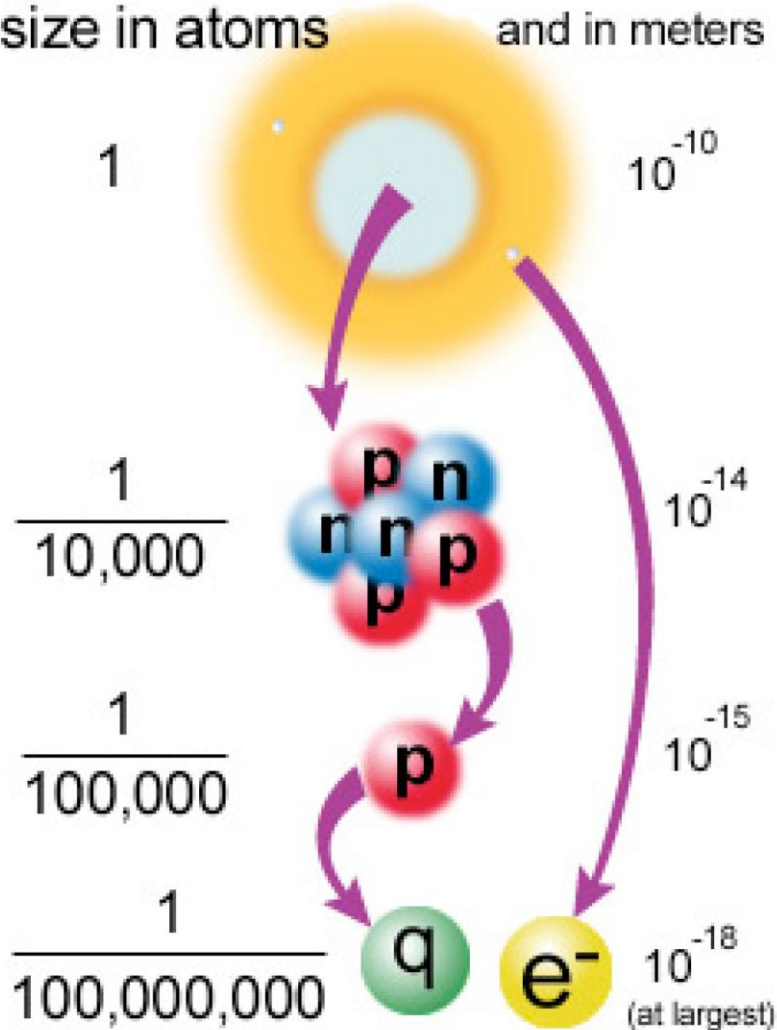
Identical analysis applied to other half of data set

A. FRANKLIN

“Selectivity and the Production of Experimental Results”

Arch. Hist. Exact Sci. 53 (1998) 399

size in atoms and in meters



Excitation energies in molecules and nuclei

- molecular excitations:

$$E_{\text{rot}} \ll E_{\text{vib}} \ll E_{\text{el}}$$

$$(\mu\text{eV} \ll \text{meV} \ll \text{eV})$$

As a consequence, these different motions can be treated separately and the wavefunction ends up as a product of terms

- In nuclei, the energy scales are much closer

$$E_{\text{rot}} \sim E_{\text{vib}} \sim E_{\text{sp}} \text{ (MeV)}$$

Collective and single-particle excitation can be separated but interact strongly

The nucleus is a bound collection of N neutrons and Z protons mass number $A=Z+N$

AZ

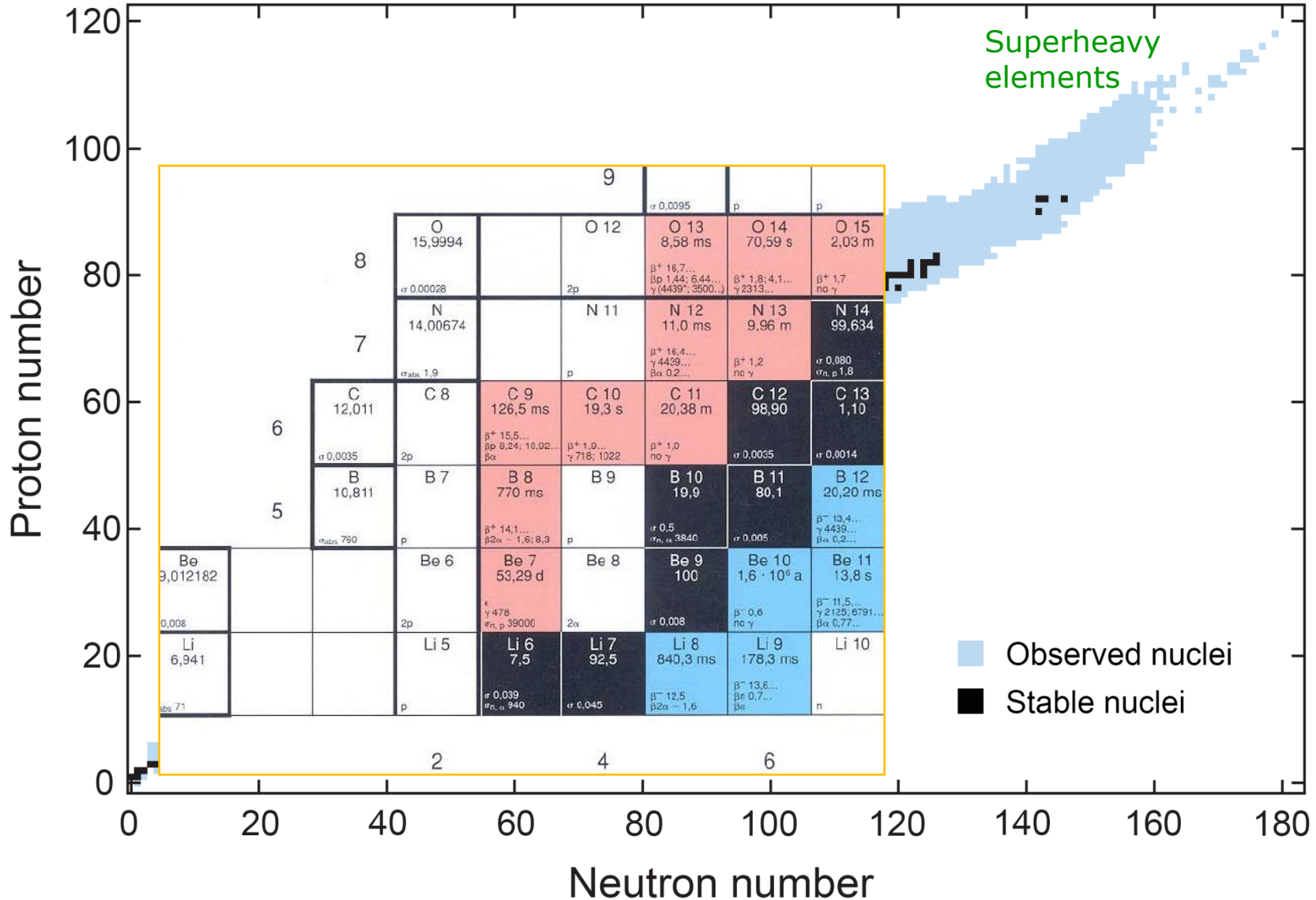
- **Isotopes:** Nuclei with the same Z but different N – e.g. ${}^9\text{C}$, ${}^{10}\text{C}$, ${}^{11}\text{C}$, ${}^{12}\text{C}$
- **Isotones:** Nuclei with the same N but different Z – e.g. ${}^9\text{C}$, ${}^8\text{B}$, ${}^7\text{Be}$, ${}^6\text{Li}$
- **Isobars:** Nuclei with the same mass number – e.g. ${}^9\text{C}$, ${}^9\text{B}$, ${}^9\text{Be}$, ${}^9\text{Li}$

Neutron number \rightarrow

				9				
					α 0,0095	p	p	
8		O 15,9994		O 12	O 13 8,58 ms β^+ 16,7... β p 1,44; 6,44... γ (4439; 3500...)	O 14 70,59 s β^+ 1,8; 4,1... γ 2313...	O 15 2,03 m β^+ 1,7 no γ	
					α 0,00028	2p		
7		N 14,00674		N 11	N 12 11,0 ms β^+ 16,4... γ 4439... $\beta\alpha$ 0,2...	N 13 9,96 m β^+ 1,2 no γ	N 14 99,634 α 0,080 α n, p 1,8	
					α 1,9	p		
6		C 12,011	C 8	C 9 126,5 ms β^+ 15,5... β p 0,24; 10,02... $\beta\alpha$	C 10 19,3 s β^+ 1,9... γ 718; 1022	C 11 20,38 m β^+ 1,0 no γ	C 12 98,90 α 0,0035	C 13 1,10 α 0,0014
					α 0,0035	2p		
5		B 10,811	B 7	B 8 770 ms β^+ 14,1... β 2 α - 1,6; 8,3	B 9	B 10 19,9 α 0,5 α n, α 3840	B 11 80,1 α 0,005	B 12 20,20 ms β^- 13,4... γ 4439... $\beta\alpha$ 0,77...
					α 760	p		
4			Be 6	Be 7 53,29 d ϵ 478 α n, p 39000	Be 8	Be 9 100 α 0,008	Be 10 $1,6 \cdot 10^9$ a β^- 0,6 no γ	Be 11 13,8 s β^- 11,5... γ 2125; 6791... $\beta\alpha$ 0,77...
					2p	2 α		
3			Li 5	Li 6 7,5 α 0,039 α n, α 340	Li 7 92,5 α 0,045	Li 8 840,3 ms β^- 12,5 β 2 α - 1,6	Li 9 178,3 ms β^- 13,6... β n 0,7... $\beta\alpha$	Li 10 n
					p			
					2	4	6	

Proton number \rightarrow

About 3000 isotopes have been made in laboratories

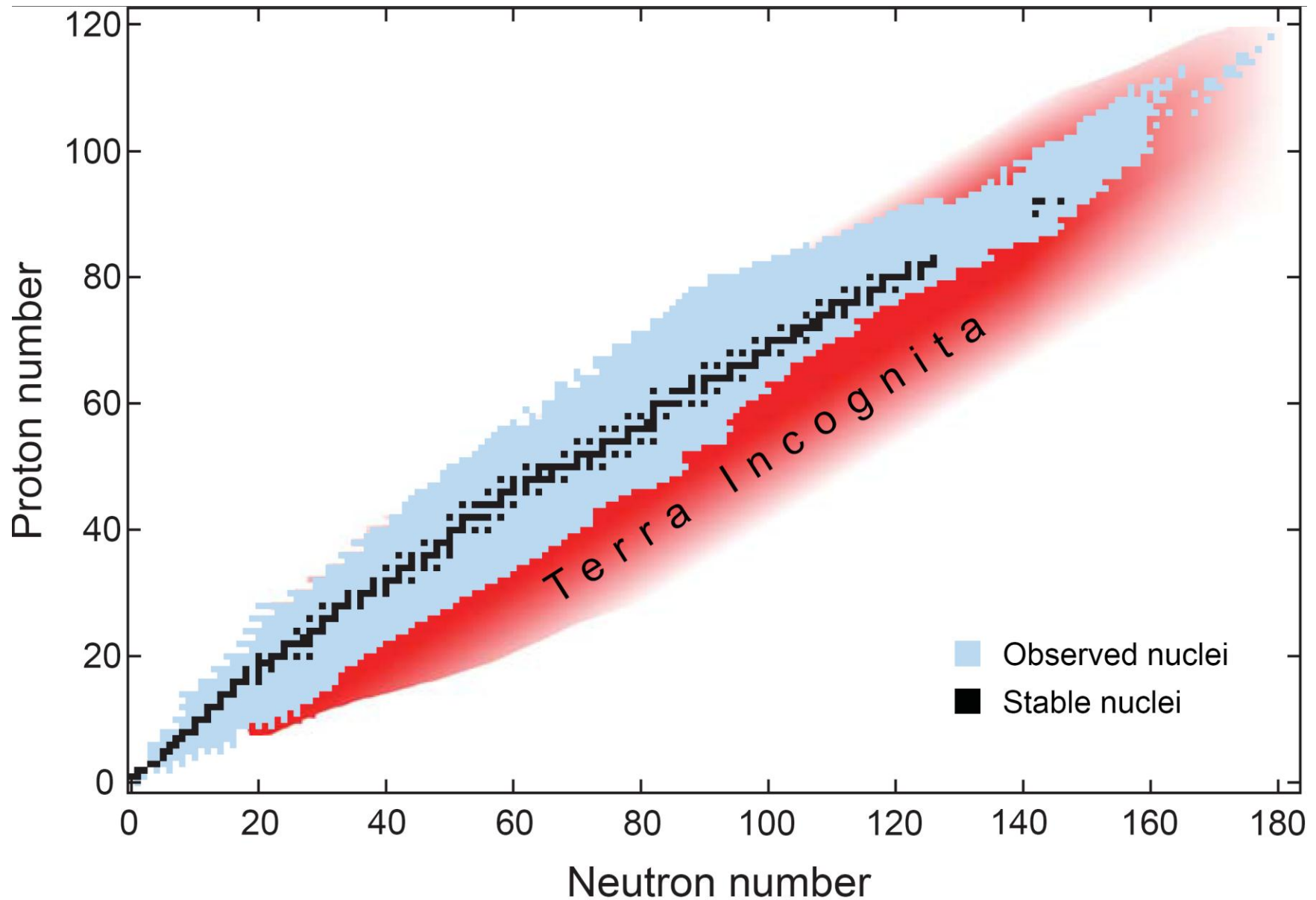




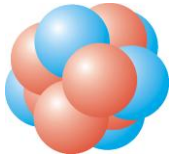
6000-8000 isotopes might be out there

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Normal Nucleus:

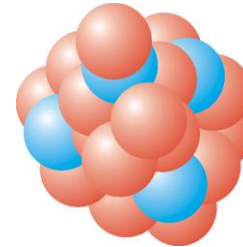


6 neutrons
6 protons (carbon)

^{12}C

Stable, found in nature

Exotic Nucleus:



16 neutrons
6 protons (carbon)

^{22}C

Radioactive, at the limit of
nuclear binding

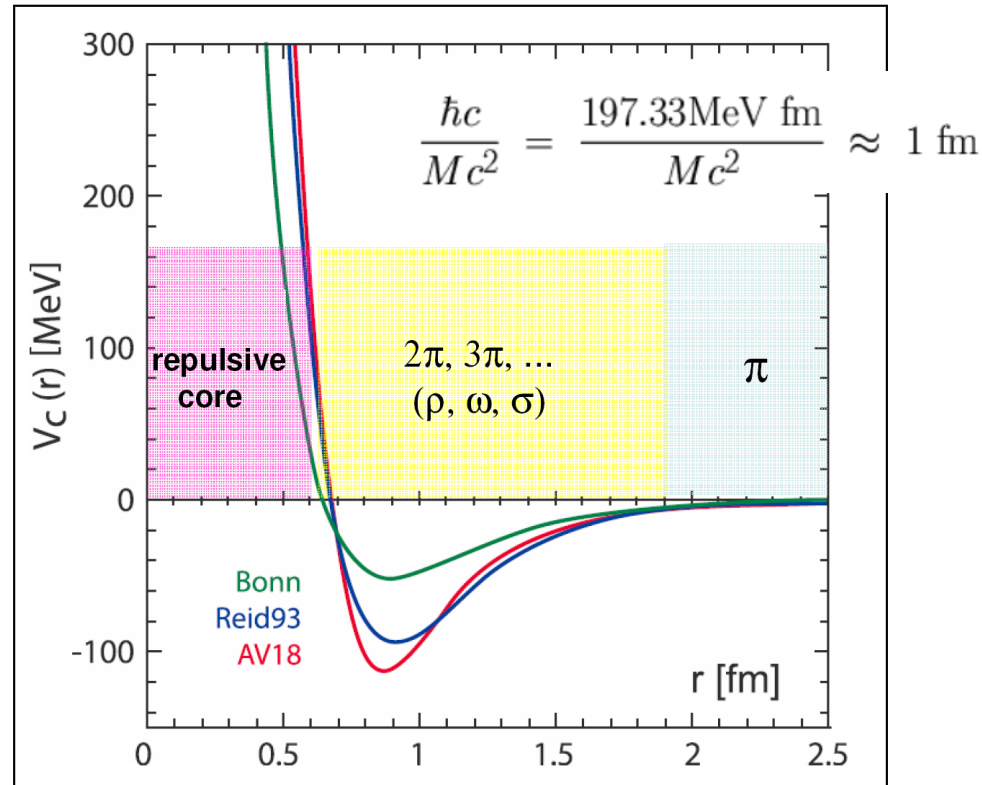
Characteristics of exotic nuclei: Excess of neutrons or protons, short half-life, neutron or proton dominated surface, low binding

The neutrons and protons are bound together by
the *strong or color force*

The strong force between the quarks in one proton and the quarks in another proton is strong enough to overcome the electromagnetic repulsion



Two ways of thinking about the strong force:
As a residual color interaction or
as the exchange of mesons



From T. Hatsuda (Oslo 2008)



Binding energy, mass and mass excess

mass $M(N, Z)$ of the neutral atom

Mass excess: $\Delta(N, Z) \equiv M(N, Z) - uA$.

Atomic mass unit u :

$$u = M(^{12}\text{C})/12 = 931.49386 \text{ MeV}/c^2.$$

Equivalent to $\Delta(^{12}\text{C})=0$

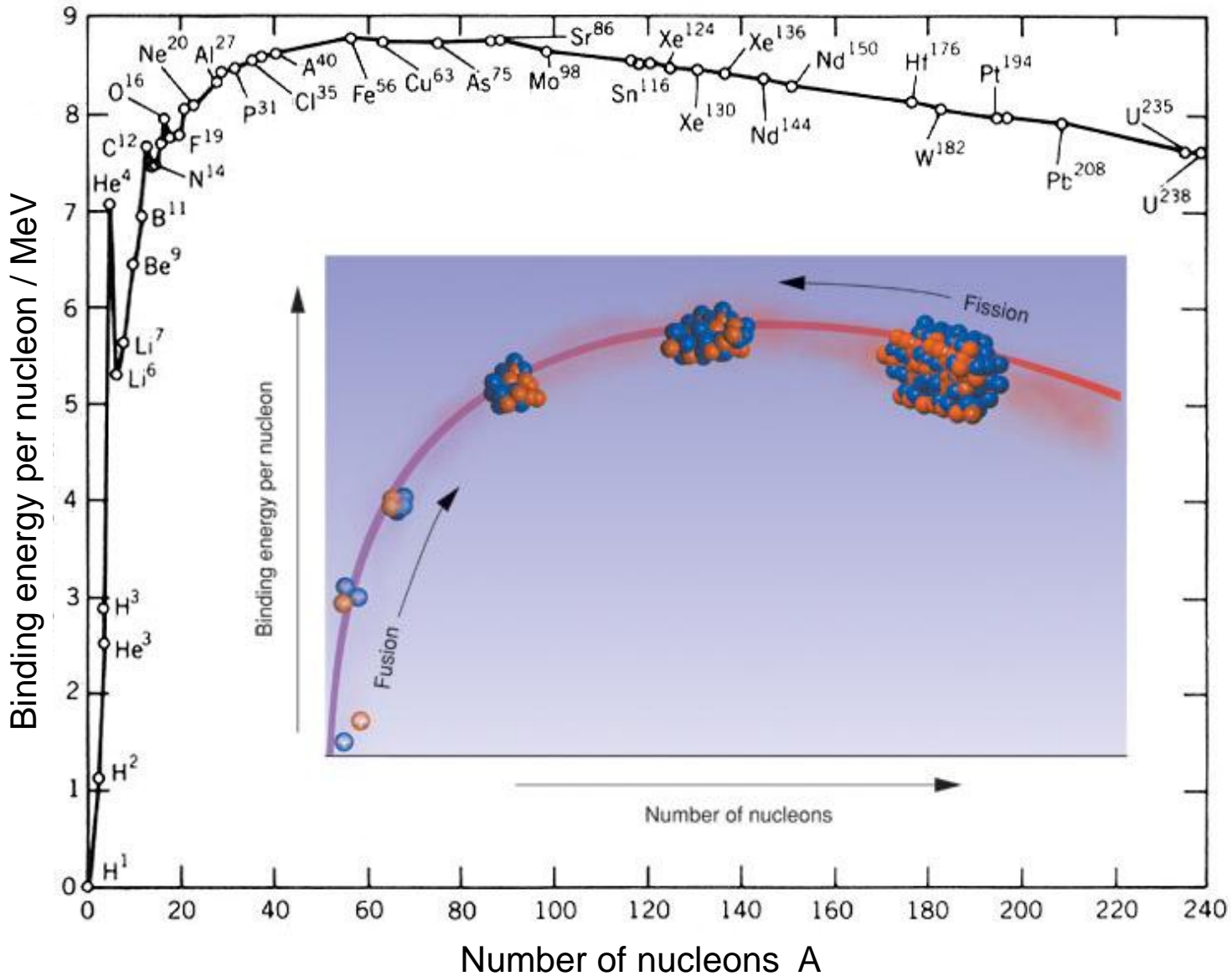
Binding energy:

$$B(N, Z) = ZM_Hc^2 + NM_n c^2 - M(N, Z)c^2$$

$$\Delta_H c^2 = 7.2890 \text{ MeV} \quad \Delta_n c^2 = 8.0713 \text{ MeV}$$

$$B(N, Z) = Z\Delta_H c^2 + N\Delta_n c^2 - \Delta(N, Z)c^2$$

Average binding energy of nuclei



A semi-empirical description of nuclear binding

- $B(Z,A) = + a_V A$ Volume term
- $+ a_S A^{2/3}$ Surface energy term
- $+ a_C Z^2/A^{1/3}$ Coulomb term
- $+ a_A (N-Z)^2/A$ Asymmetry term
- $- a_P/A^{3/4}$ Pairing term

$$R \sim A^{1/3}$$

$$a_V = -15.68 \text{ MeV}$$

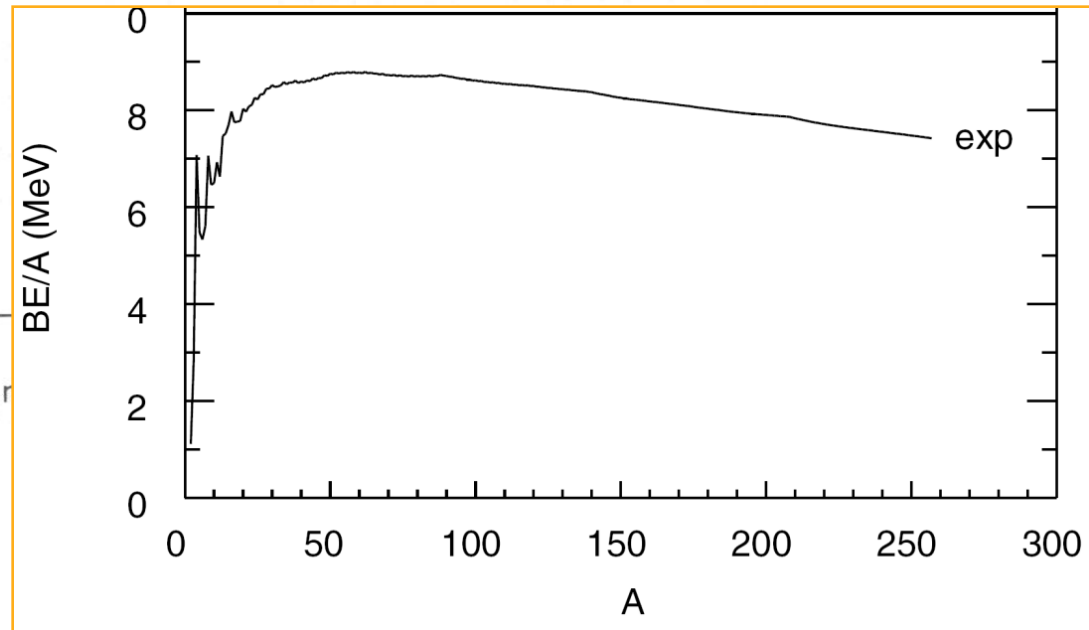
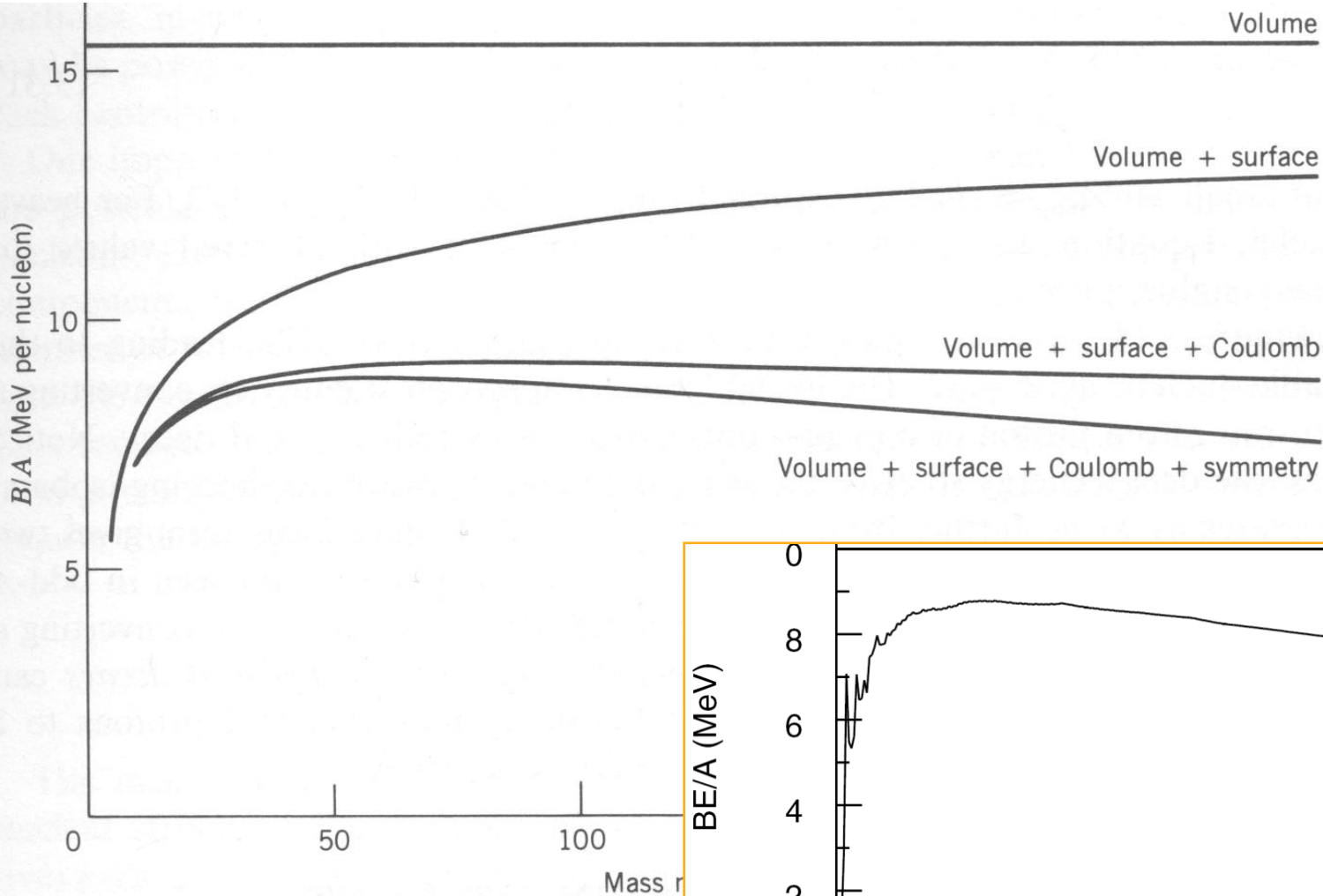
$$a_S = 18.56 \text{ MeV}$$

$$a_C = 0.717 \text{ MeV}$$

$$a_A = 28.1 \text{ MeV}$$

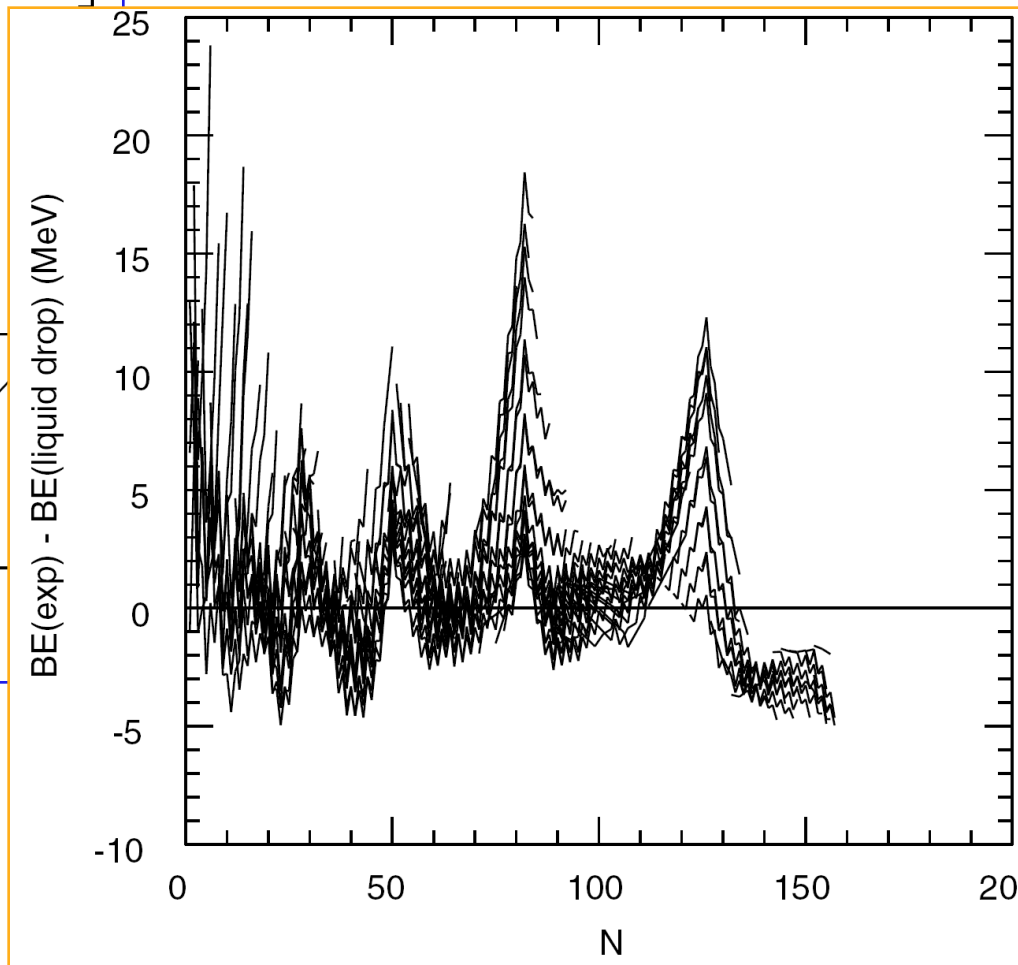
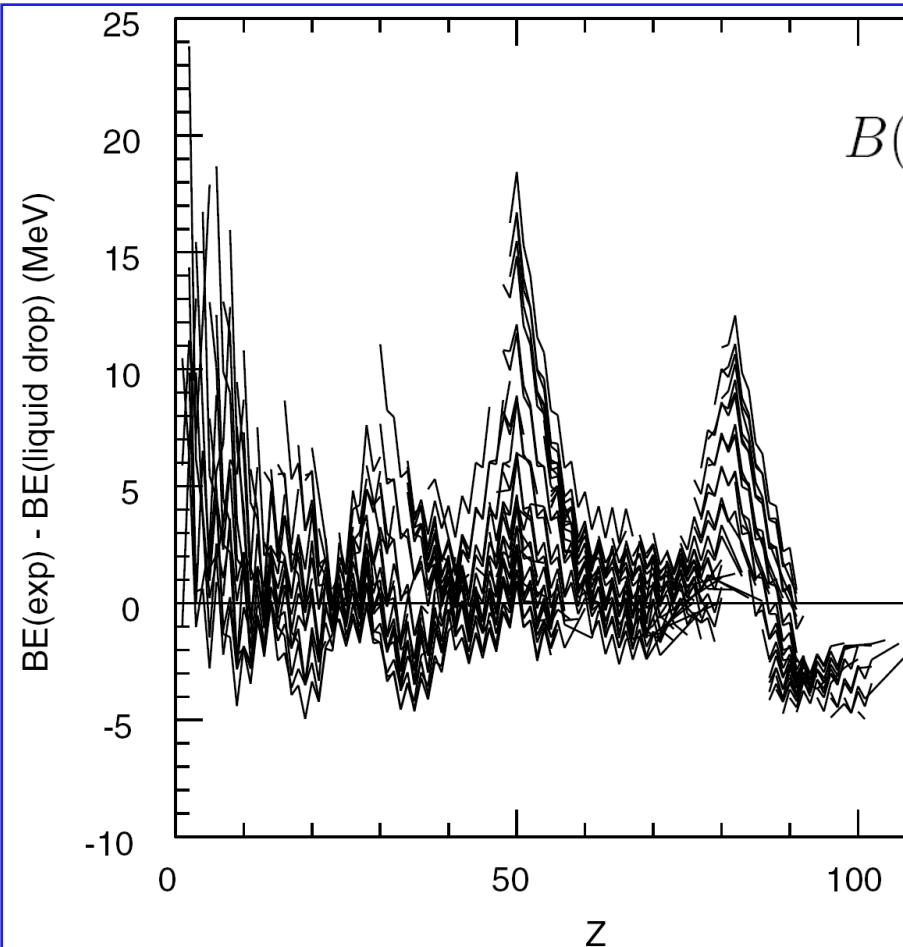
$$a_P = 34.0 \text{ MeV for even-even, } -34.0 \text{ MeV for odd-odd, } 0 \text{ for even-odd}$$

The different contributions to nuclear binding



Semi-empirical mass formula versus experimental reality

$$B(N, Z) = ZM_Hc^2 + NM_n c^2 - M(N, Z)c^2$$





Q values and nucleon separation energies

Q value of a process (${}^AZ_i \rightarrow {}^AZ_f$):

$$Q = \sum_i M(N_i, Z_i)c^2 - \sum_f M(N_f, Z_f)c^2 = \sum_f B(N_f, Z_f) - \sum_i B(N_i, Z_i)$$

Nucleon separation energies:

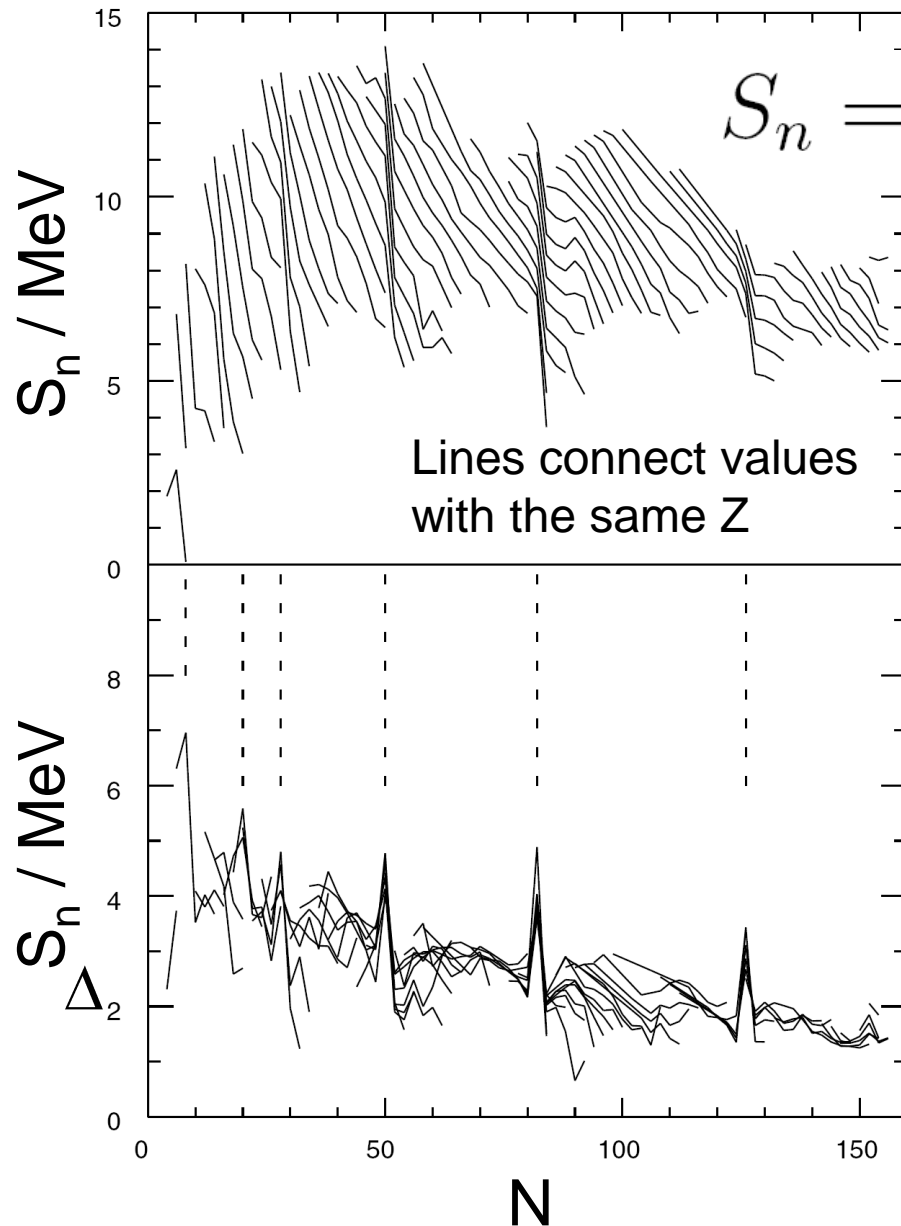
$$S_n = B(N, Z) - B(N - 1, Z),$$

$$S_p = B(N, Z) - B(N, Z - 1),$$

$$S_{2n} = B(N, Z) - B(N - 2, Z),$$

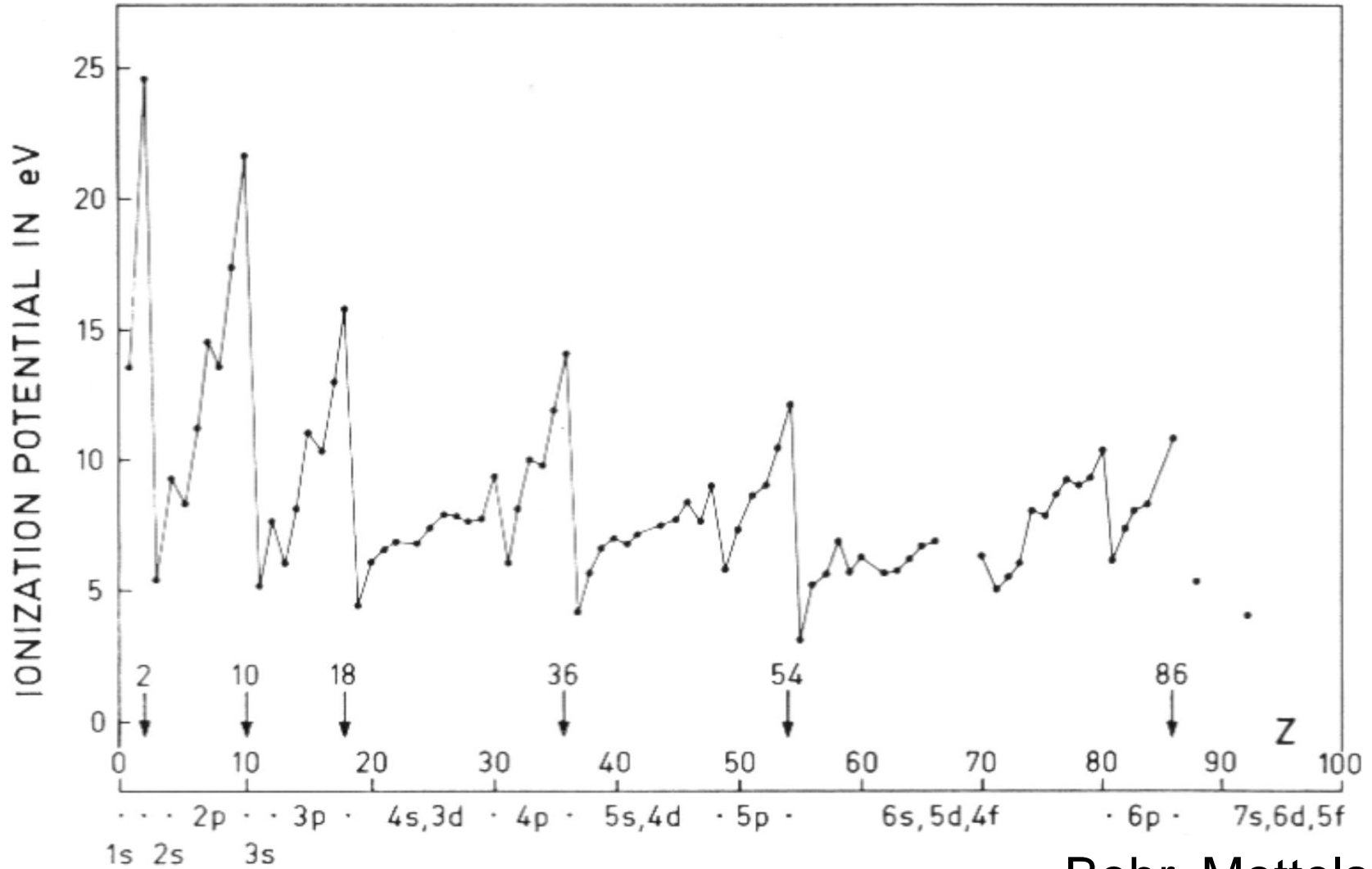
$$S_{2p} = B(N, Z) - B(N, Z - 2).$$

Nucleon separation energies



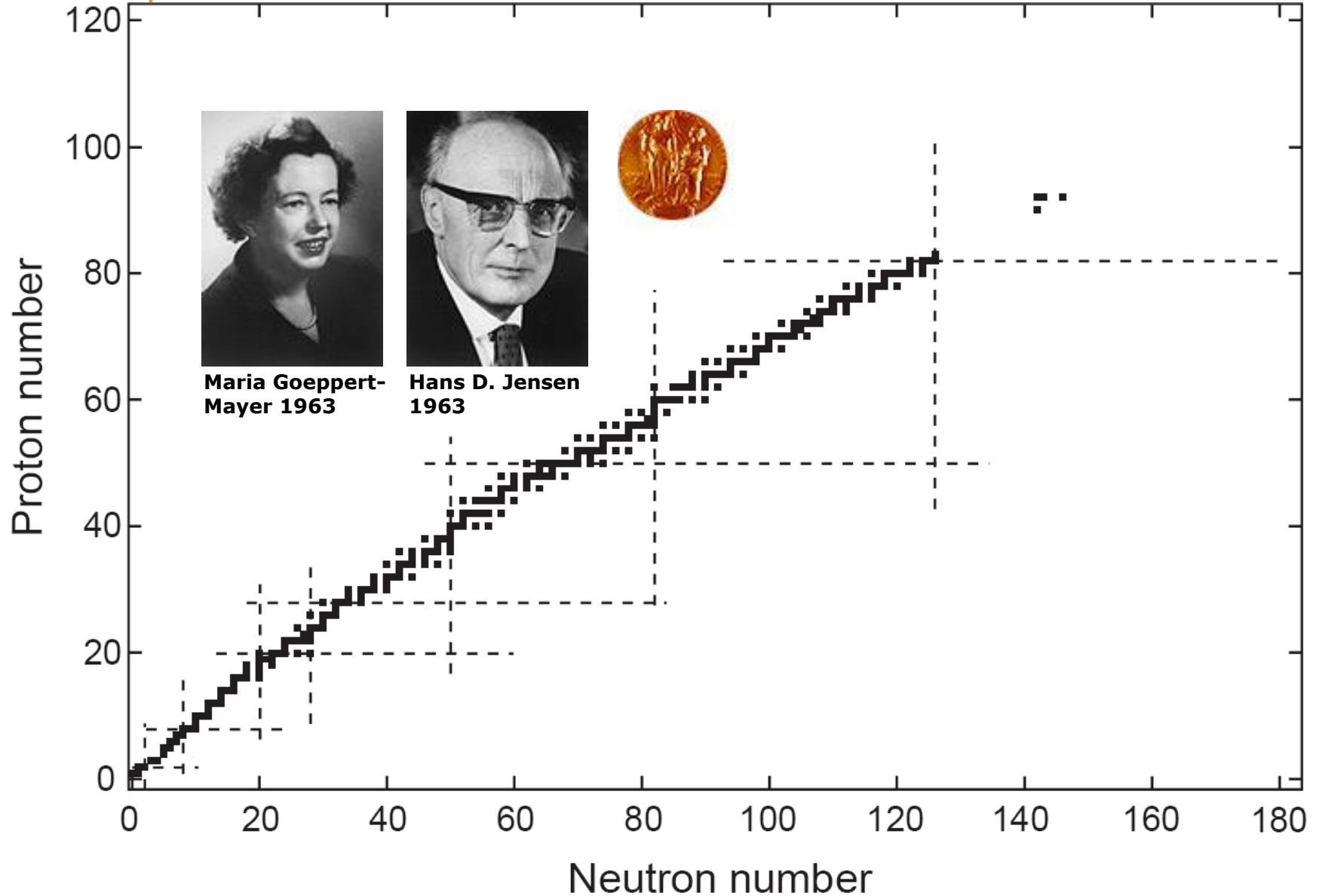
Huge changes in the separation energy at neutron numbers 8, 20, 28, 50, 82, 126
→ Those nuclei are particularly stable

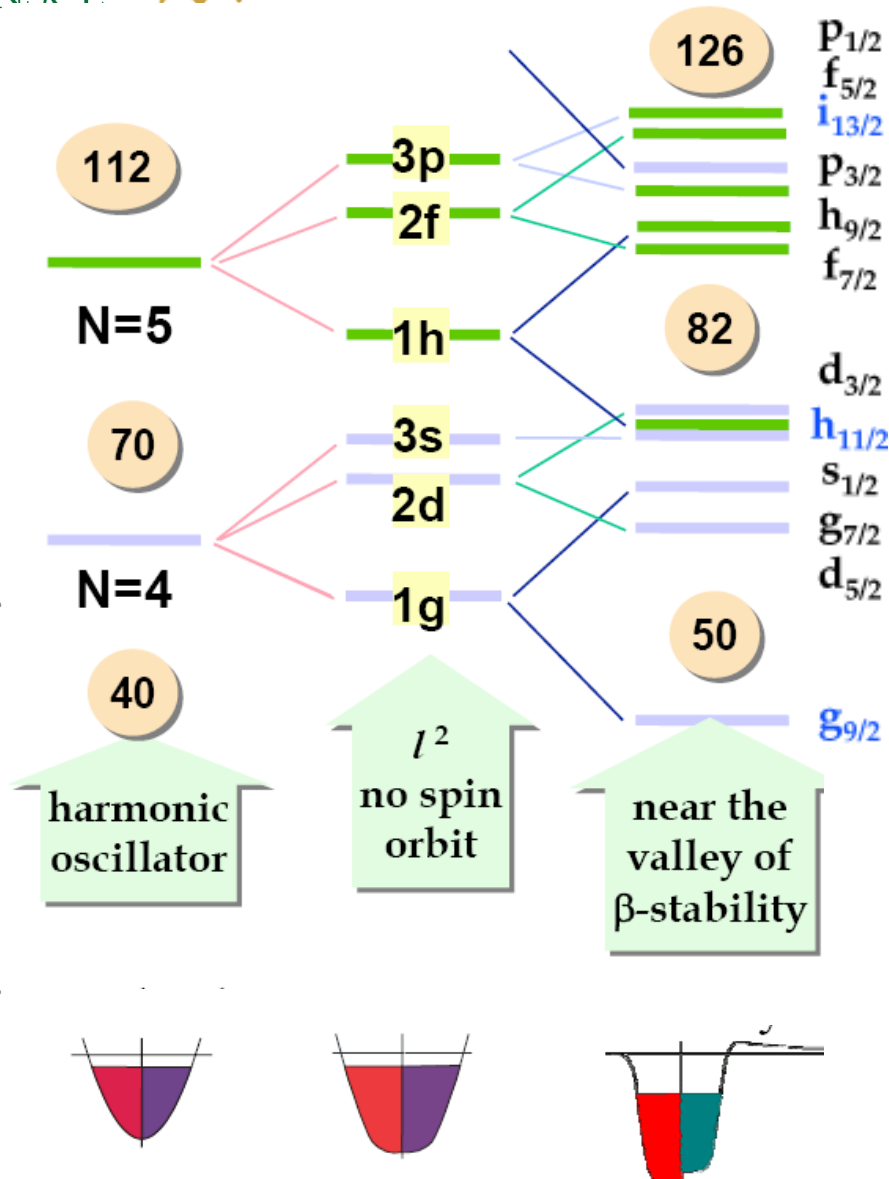
Seen that before ... compare to atomic shell structure!



Bohr, Mottelson

Shell structure – magic numbers



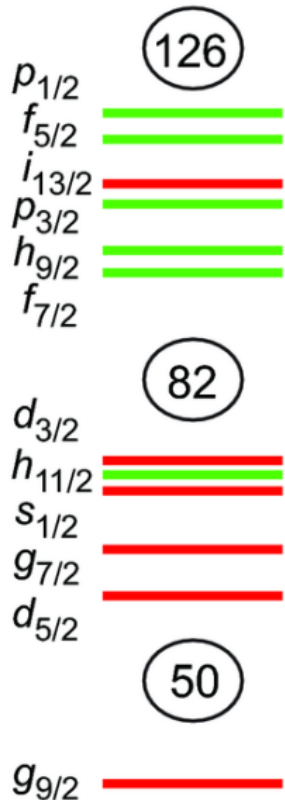


- Single-particle levels in nuclei**
 The single-particle levels of this fermionic system are grouped. Large, stabilizing gaps between groups of single-particle states occur at certain occupation numbers of the orbits with a “magic number” of protons and neutrons
- Magic numbers**
 Numbers of neutrons and protons in nuclei which correspond to particularly stable structures (2, 8, 20, 28, 50, 82, 126)
- $\ell = 0, 1, 2, 3, \dots$ $j = \ell \pm 1/2$
 s, p, d, f, ... Max. occupancy: $2j+1$
- Experimental signatures of nuclear shells**
 - low capture cross sections
 - little collectivity
 - more tightly bound than neighboring nuclei

$$H = H_0 + H_{res} = \sum_{i=1}^A \left[\frac{\mathbf{p}_i^2}{2m_i} + U_i(\mathbf{r}) \right] + H_{res}$$

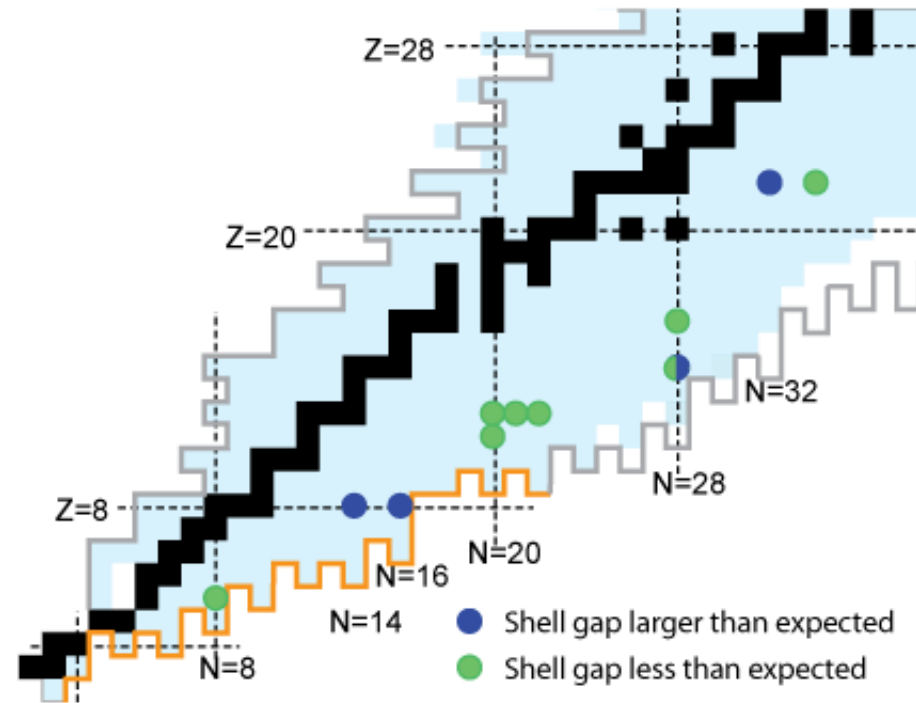
Nuclear Shell Structure

Near stability



towards
neutron-rich
nuclei

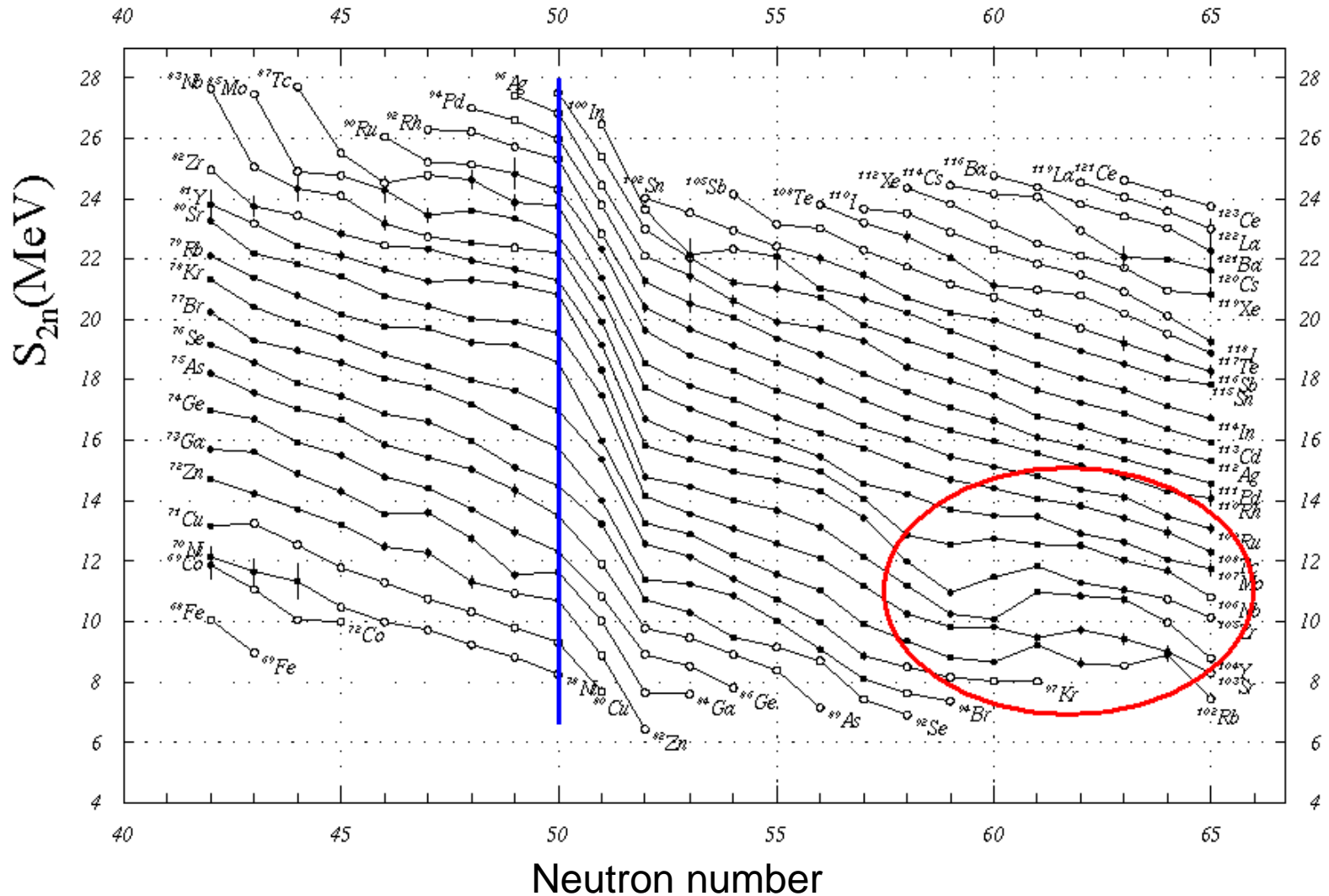
For $N \gg Z$



- Mean field near stability
- Strong spin-orbit term

- Mean field for $N \gg Z$?
- Reduced spin-orbit
- Diffuse density
- Tensor force

An indicator for changes in nuclear structure



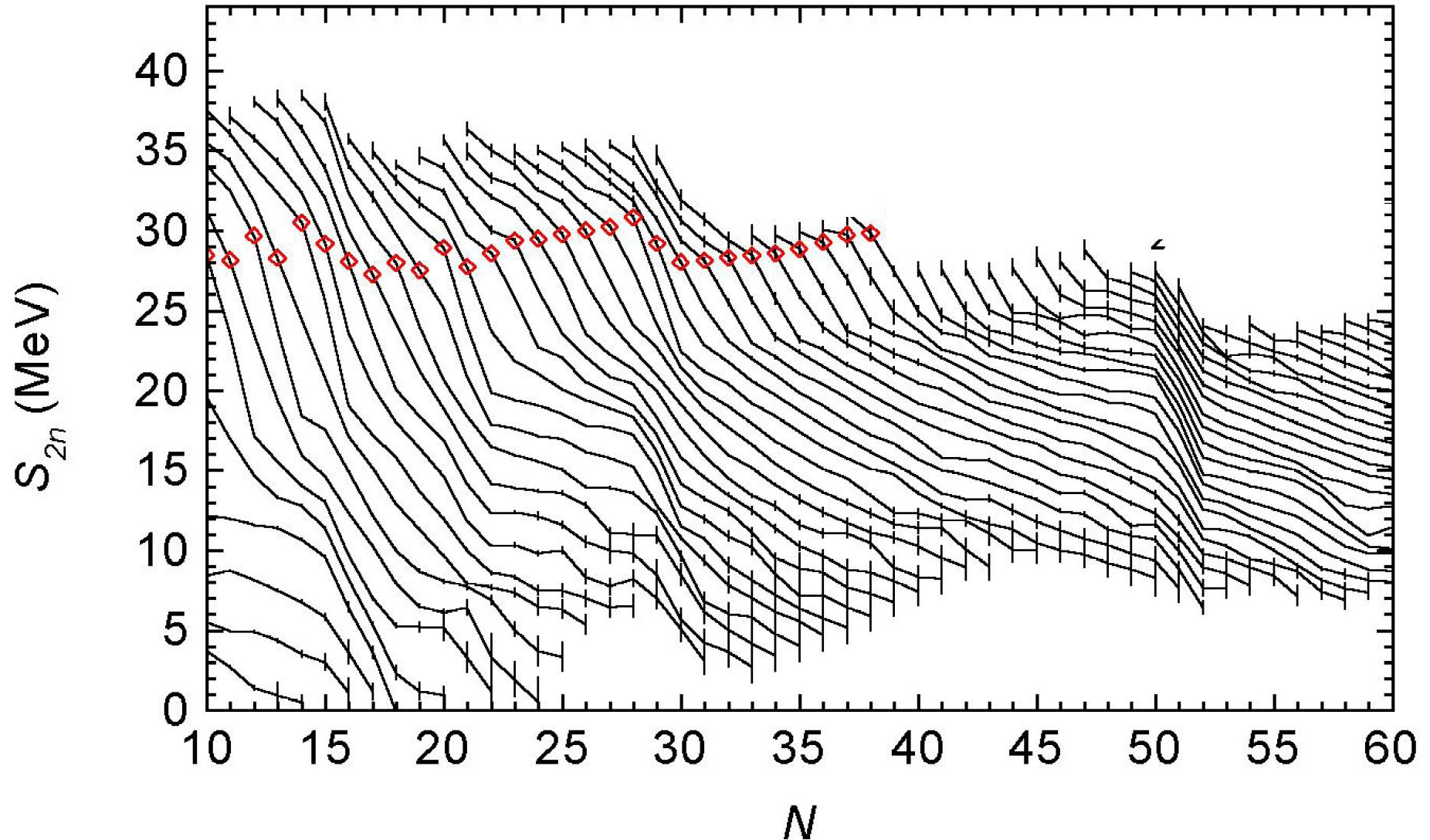
magic number

deformation



Masses – what are they good for?

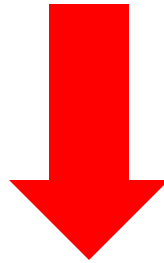
Nuclear structure





Observables

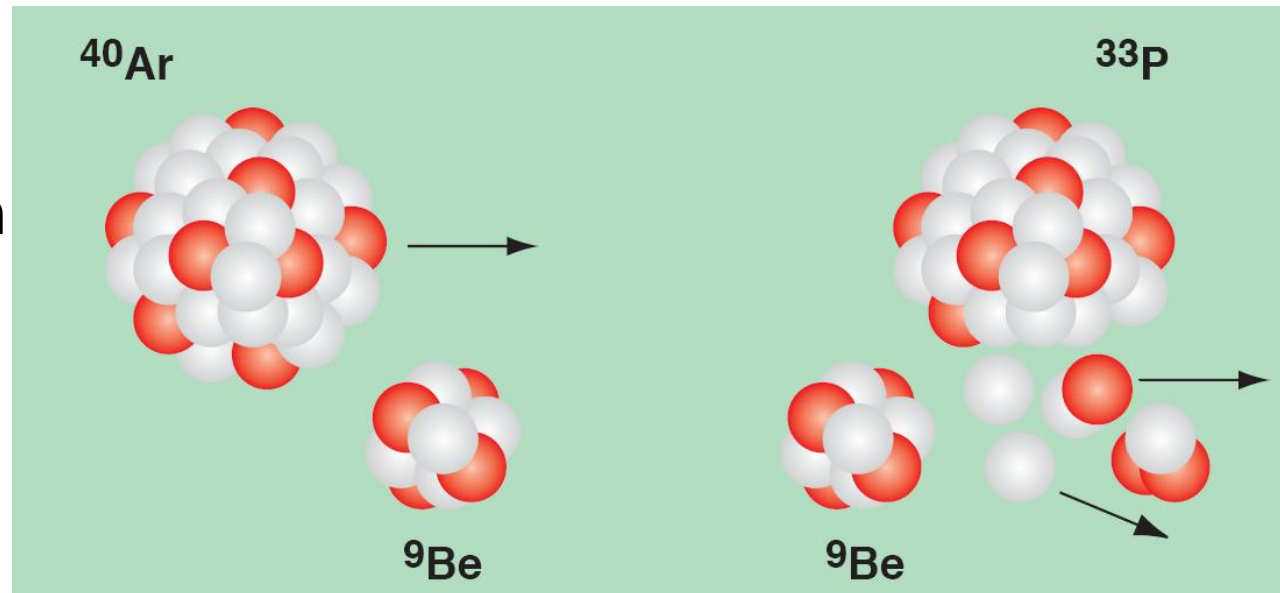
$$\begin{aligned} S_{2n} &= B(N,Z) - B(N-2,Z) \\ &= M(N-2,Z) + 2M_N - M(N,Z) \end{aligned}$$



Measure nuclear masses of exotic nuclei!

... but first we have to produce them!

- Transfer reactions
- Fusion-evaporation
- Fission
- Fragmentation



- Target fragmentation (HRIBF, TRIUMF, SPIRAL, ISOLDE)
- Projectile fragmentation (NSCL, GSI, RIKEN, GANIL)

Indirect

- Decay measurements and kinematics in two-body reactions

reactions:



$$Q = M_A + M_a - M_b - M_B$$

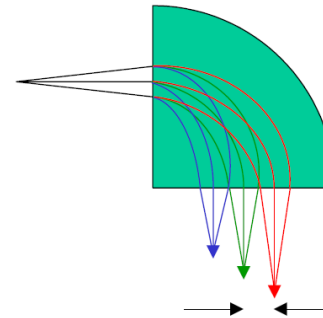
decays:



$$Q_\alpha = M_B - M_A$$

Direct

- Conventional mass spectrometry
 - Cern PS, Chalk River
- Time-of-flight
 - spectrometer (SPEG, TOFI, S800)
 - Multi-turn (cyclotrons, storage rings)
- Frequency measurements
 - Penning traps
 - Storage rings



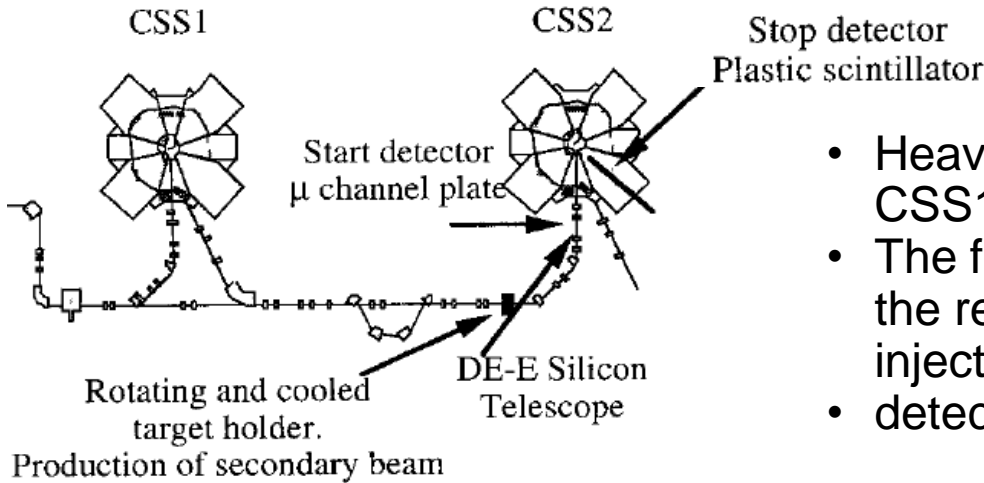
Mass separator
(spectrograph,
spectrometer)

Dispersion

$$D = \Delta x / \Delta m$$

Adapted from D. Lunney

TOF mass measurement – Cyclotrons at GANIL



A~100 nuclei: $^{50}\text{Cr}+^{58}\text{Ni}$ at 250 MeV

- Heavy-ion primary beam delivered by CSS1 with a few MeV/nucleon
- The fusion-evaporation products formed in the reaction with the production target are injected into the CSS2 and accelerated
- detected in a silicon-detector telescope

$$\delta m/m = \delta t/t$$

Mass Measurement of ^{100}Sn

M. Chartier,¹ G. Auger,¹ W. Mittig,¹ A. Lépine-Szily,² L. K. Fifield,³ J. M. Casandjian,¹ M. Chabert,¹ J. Fermé,¹
A. Gillibert,⁴ M. Lewitowicz,¹ M. Mac Cormick,¹ M. H. Moscatello,¹ O. H. Odland,⁵ N. A. Orr,⁶ G. Politi,⁷
C. Spitaels,¹ and A. C. C. Villari¹

$$\text{M.E.}(^{100}\text{Cd}) = -74.180 \pm 0.200(\text{syst}) \text{ MeV},$$

$$\begin{aligned} \text{M.E.}(^{100}\text{In}) &= -64.650 \pm 0.300(\text{syst}) \\ &\pm 0.100(\text{stat}) \text{ MeV}, \end{aligned}$$

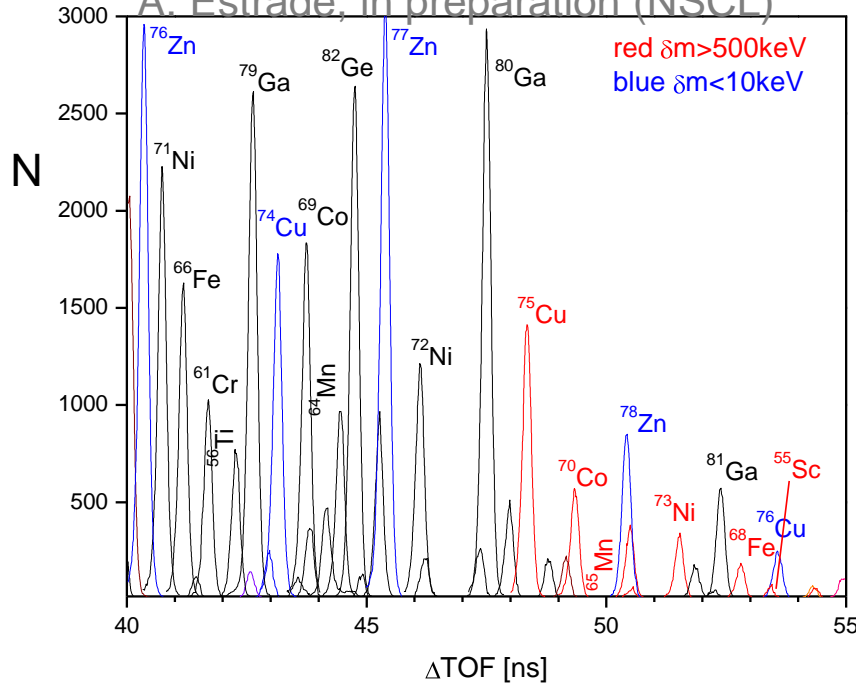
$$\begin{aligned} \text{M.E.}(^{100}\text{Sn}) &= -57.770 \pm 0.300(\text{syst}) \\ &\pm 0.900(\text{stat}) \text{ MeV}. \end{aligned}$$

$$\frac{B}{\omega/h} = \gamma \frac{m}{q} = \frac{B\rho}{v}, \quad \omega_c = \omega/h$$

$$\frac{\delta T_{\text{turn}}}{T_{\text{turn}}} = \frac{\delta m/q}{m/q} \quad h = \# \text{rf periods/turn}$$

TOF mass measurements – Spectrographs at NSCL

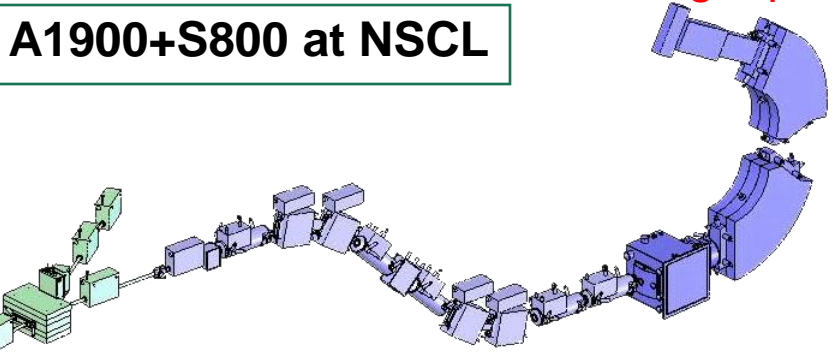
A. Estrade, in preparation (NSCL)



TOF mass measurements on neutron-rich isotopes
goal: $\delta m = 0.2 \text{ MeV}$ for $A \sim 70$
 $\rightarrow \delta m/m = 2 \times 10^{-6}$

TOF stop
58m flight path

A1900+S800 at NSCL

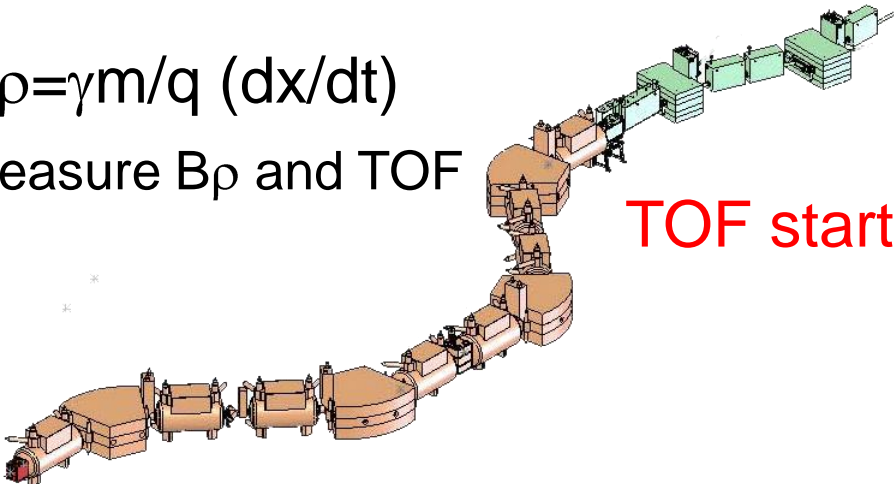


$B\rho = \gamma m/q \text{ (dx/dt)}$

Measure $B\rho$ and TOF

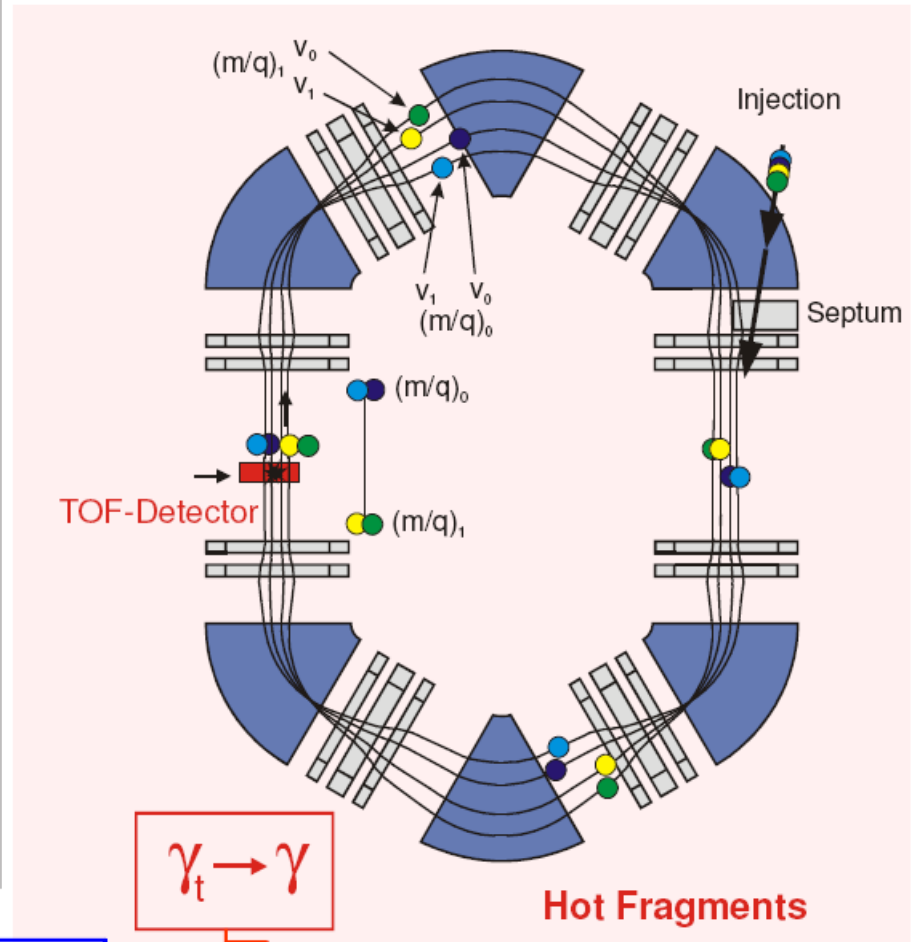
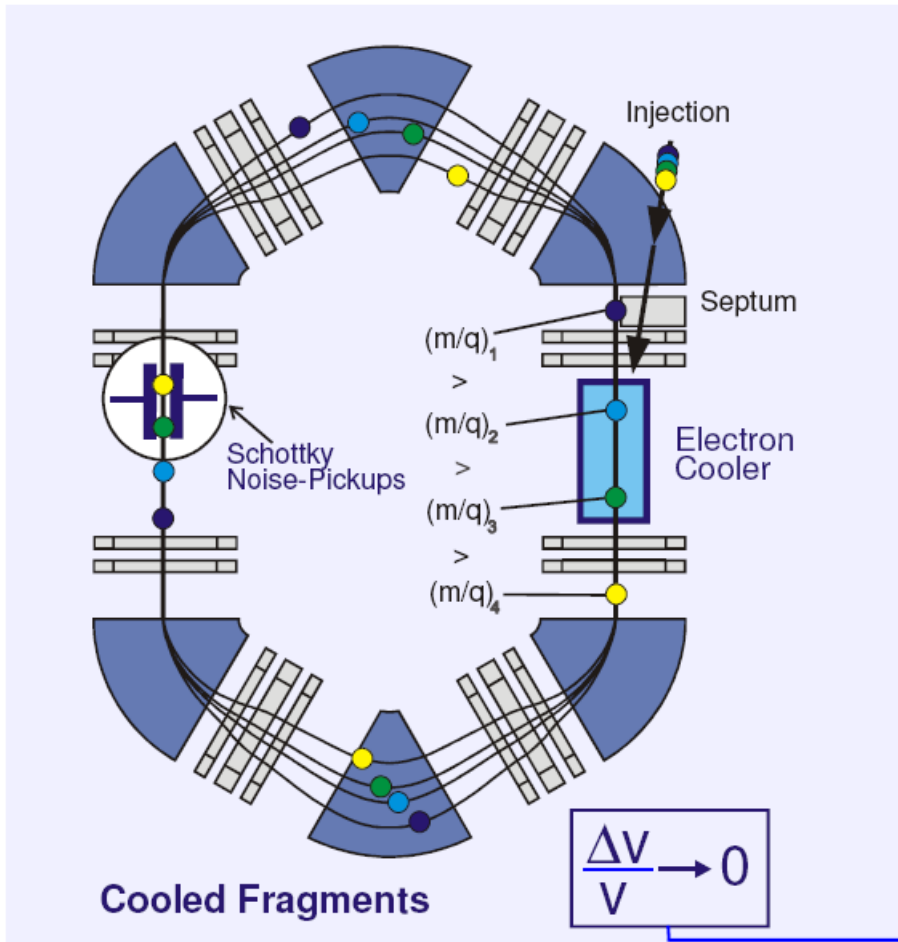
TOF start

- Measure many masses simultaneously
- Mass accuracy: $\Delta m/m \sim 10^{-6}$
- Beam rate: particles/min (e.g 10000 particles total for $\delta m \sim 200 \text{ keV}$ for $A \sim 100$)



SCHOTTKY MASS SPECTROMETRY

ISOCHRONOUS MASS SPECTROMETRY



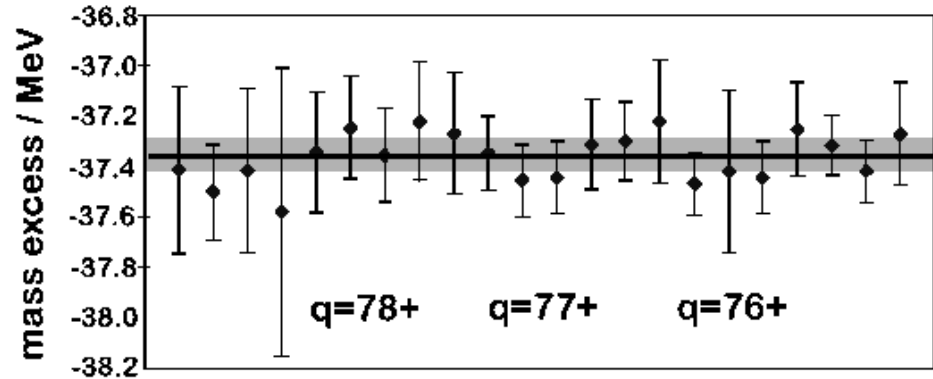
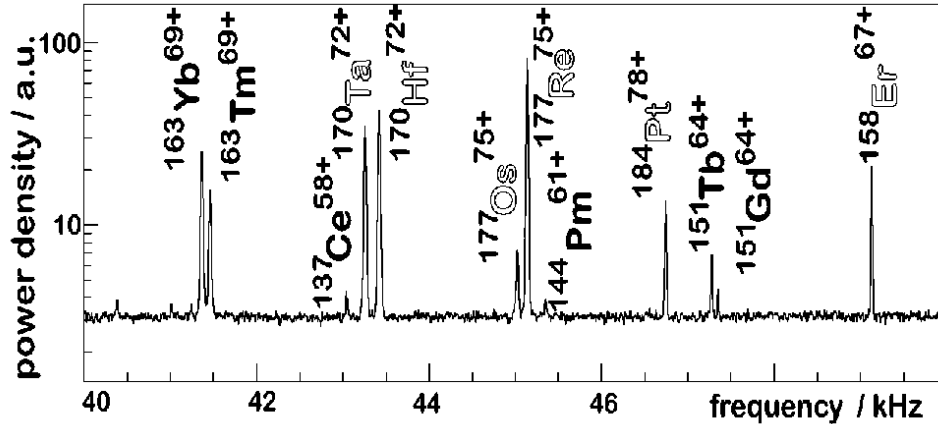
$T_{1/2} > 1 \text{ s}$

$$\frac{\Delta f}{f} = -\frac{1}{\gamma_t^2} \frac{\Delta(m/q)}{m/q} + \frac{\Delta V}{V} \left(1 - \frac{\gamma^2}{\gamma_t^2}\right)$$

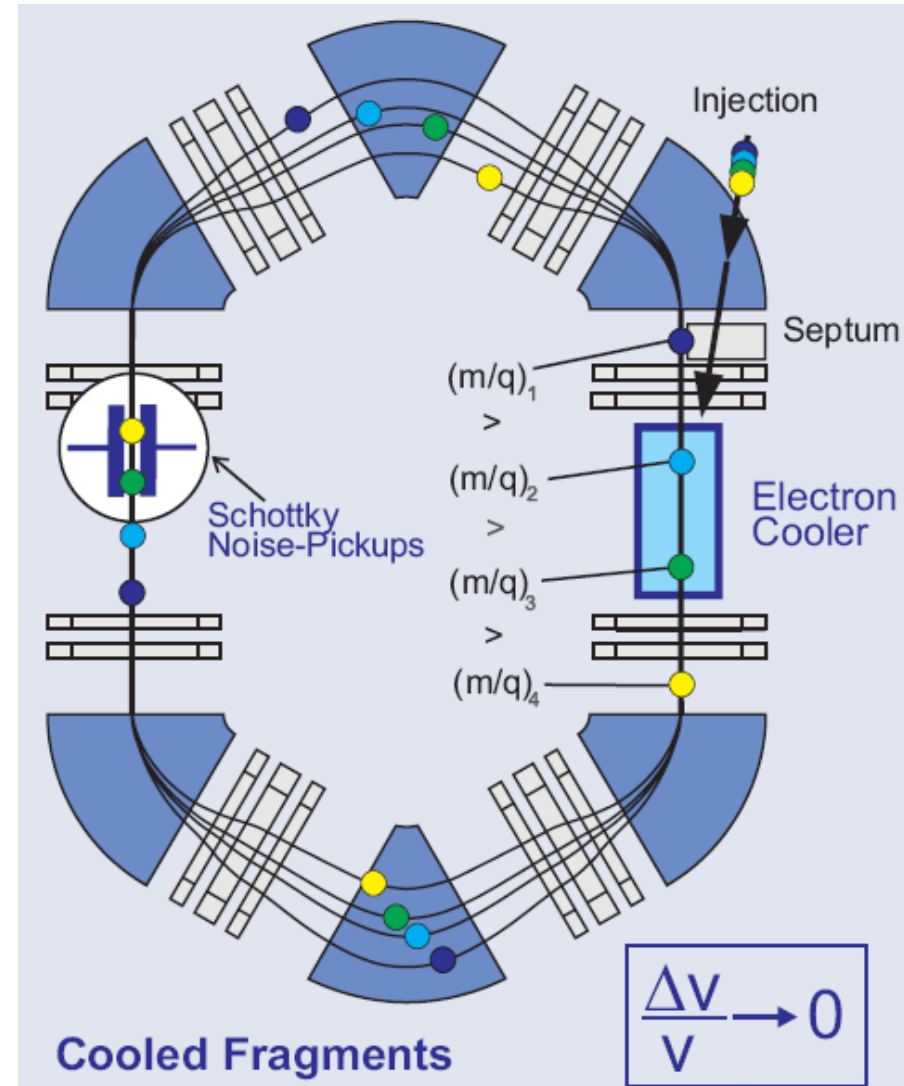
$T_{1/2} > 10 \mu\text{s}$

γ_t : relative change in path length by turn relative to change in Bp

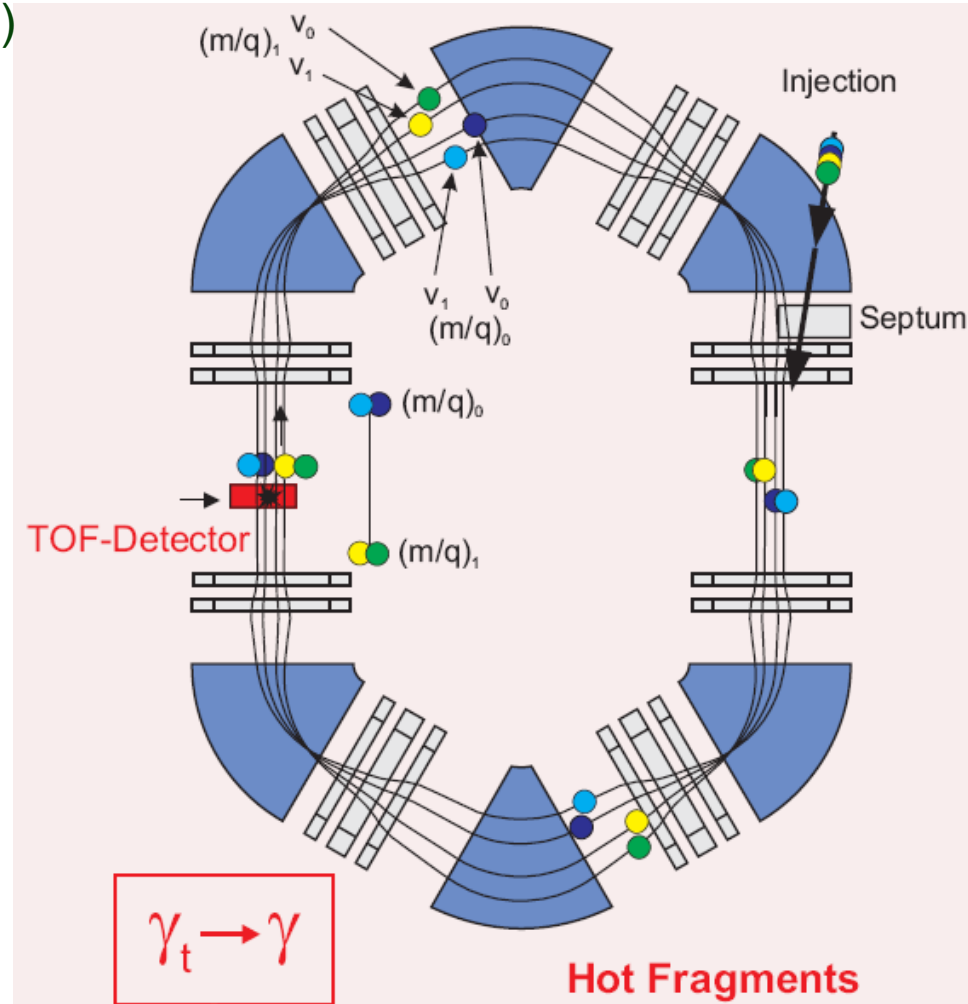
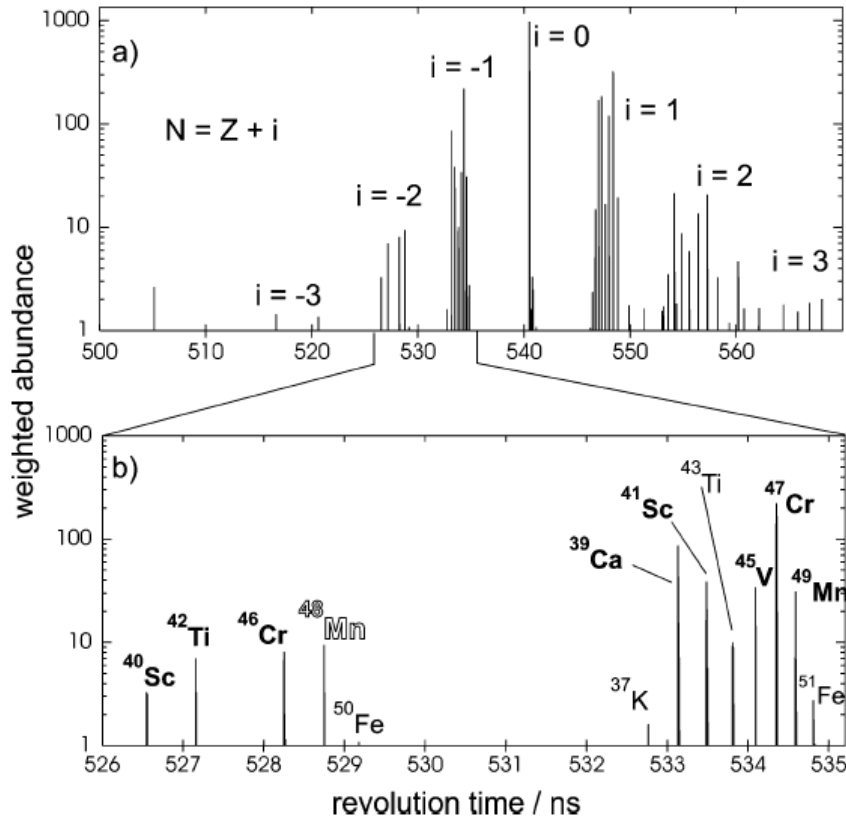
- Schottky spectrometry in storage ring (GSI), e.g. ^{184}Pt



Mass excess for ^{184}Pt as determined in several runs using different reference isotopes and in different ionic charge states q . ($dm/m=5 \cdot 10^{-7}$)

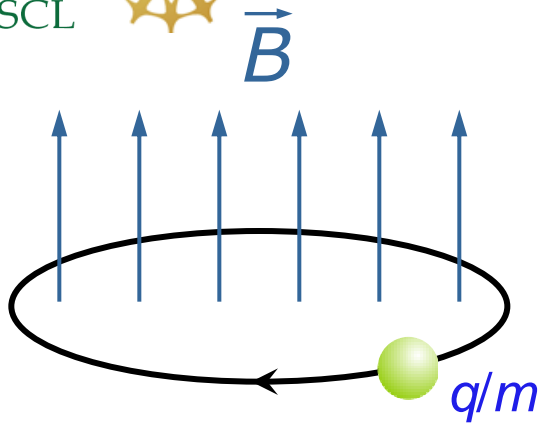


- Mass measurement of short-lived ^{44}V , ^{48}Mn , ^{41}Ti and ^{45}Cr (X-ray burst models)



Accuracy of $\delta m = 100\text{-}500$ keV was achieved (lifetimes ~ 100 ms)

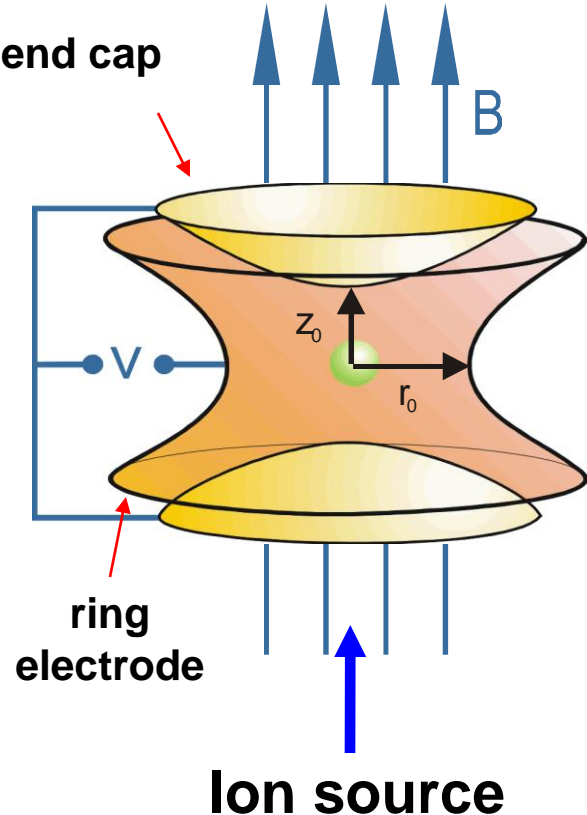
Mass measurements with Penning traps



Mass measurement via determination of cyclotron frequency

$$f_c = \frac{1}{2\pi} \cdot \frac{q}{m} \cdot B$$

from characteristic motion of stored ions



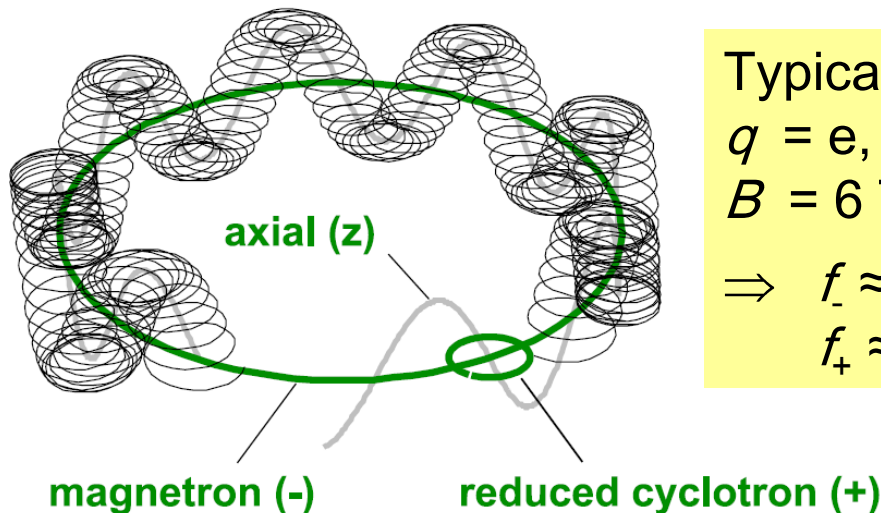
PENNING trap

- Strong homogeneous magnetic field of known strength B provides radial confinement
- Weak electric 3D quadrupole field provides axial confinement

Mass measurements with Penning traps

Motion of an ion is the superposition of three characteristic harmonic motions:

- axial motion (frequency f_z)
- magnetron motion (frequency f_-)
- modified cyclotron motion (frequency f_+)



Typical frequencies

$$q = e, \quad m = 100 \text{ u}, \quad B = 6 \text{ T}$$

$$\Rightarrow f_- \approx 1 \text{ kHz}$$

$$f_+ \approx 1 \text{ MHz}$$

Excite the cyclotron motion with multipolar RF (Goal: excite the cyclotron motion to resonance)



Transform radial to axial energy (gradient dB/dz) and eject ions

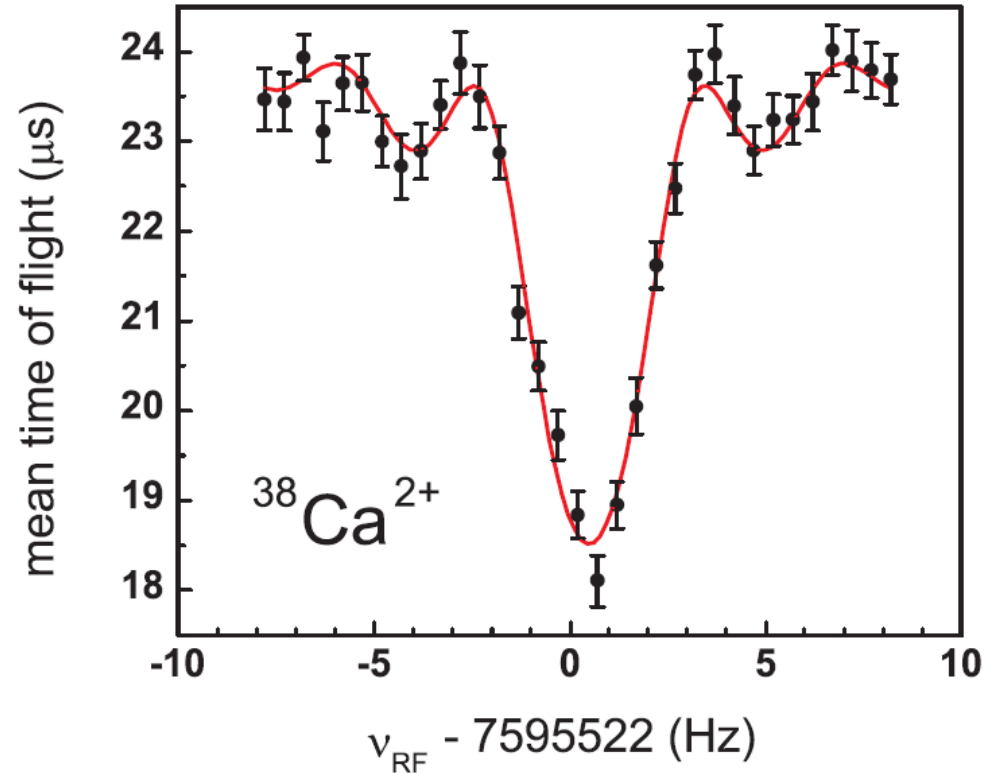
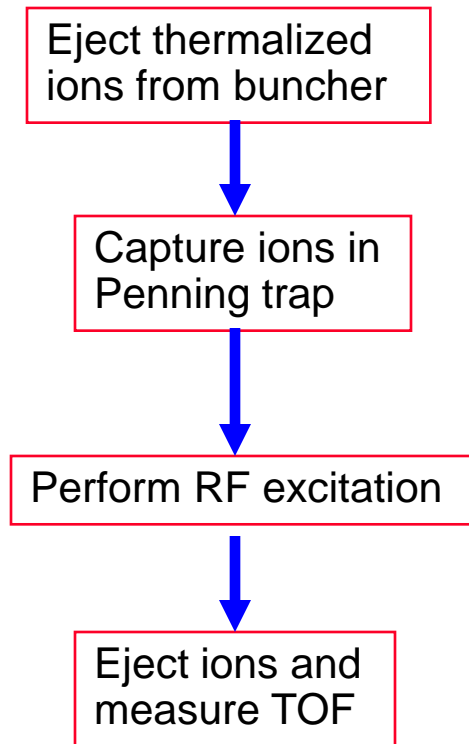


Measure time of flight (TOF) - the shorter TOF, the closer is the excitation frequency to the resonance

The frequencies of the radial motions obey the relation

$$f_+ + f_- = f_c = \frac{1}{2\pi} \cdot \frac{q}{m} \cdot B$$

Mass measurements with Penning traps

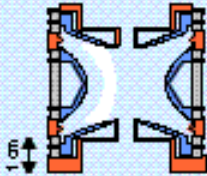


$$ME = -22058.53(28) \text{ keV}$$

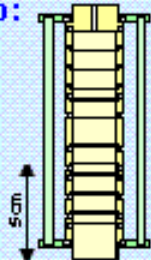
$$\delta m = 280 \text{ eV}$$

The Triple Trap Mass Spectrometer ISOLTRAP

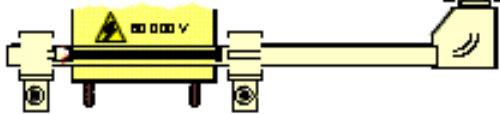
Precision trap:
Mass measurement



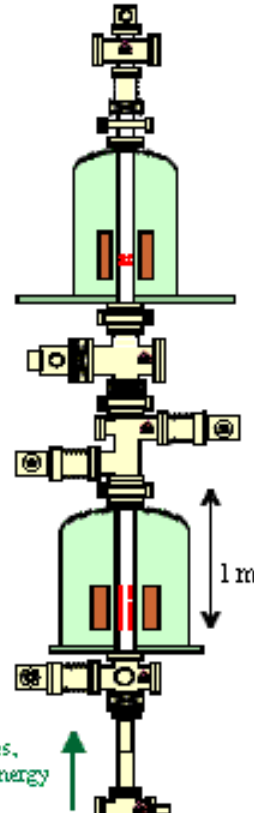
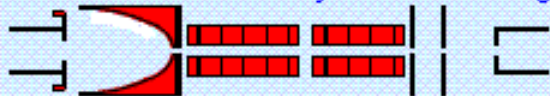
Cooler trap:
mass selective cooling



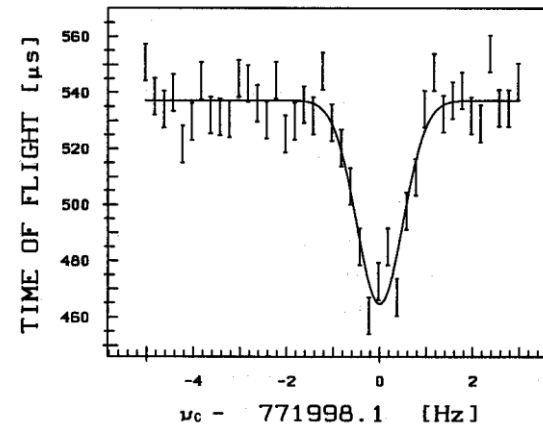
60keV ISOLDE
ion beam



Beam Buncher: - accumulation of continuous ion beam
- bunched ejection at low energy



- e.g. $^{123,124,126}\text{Ba}$

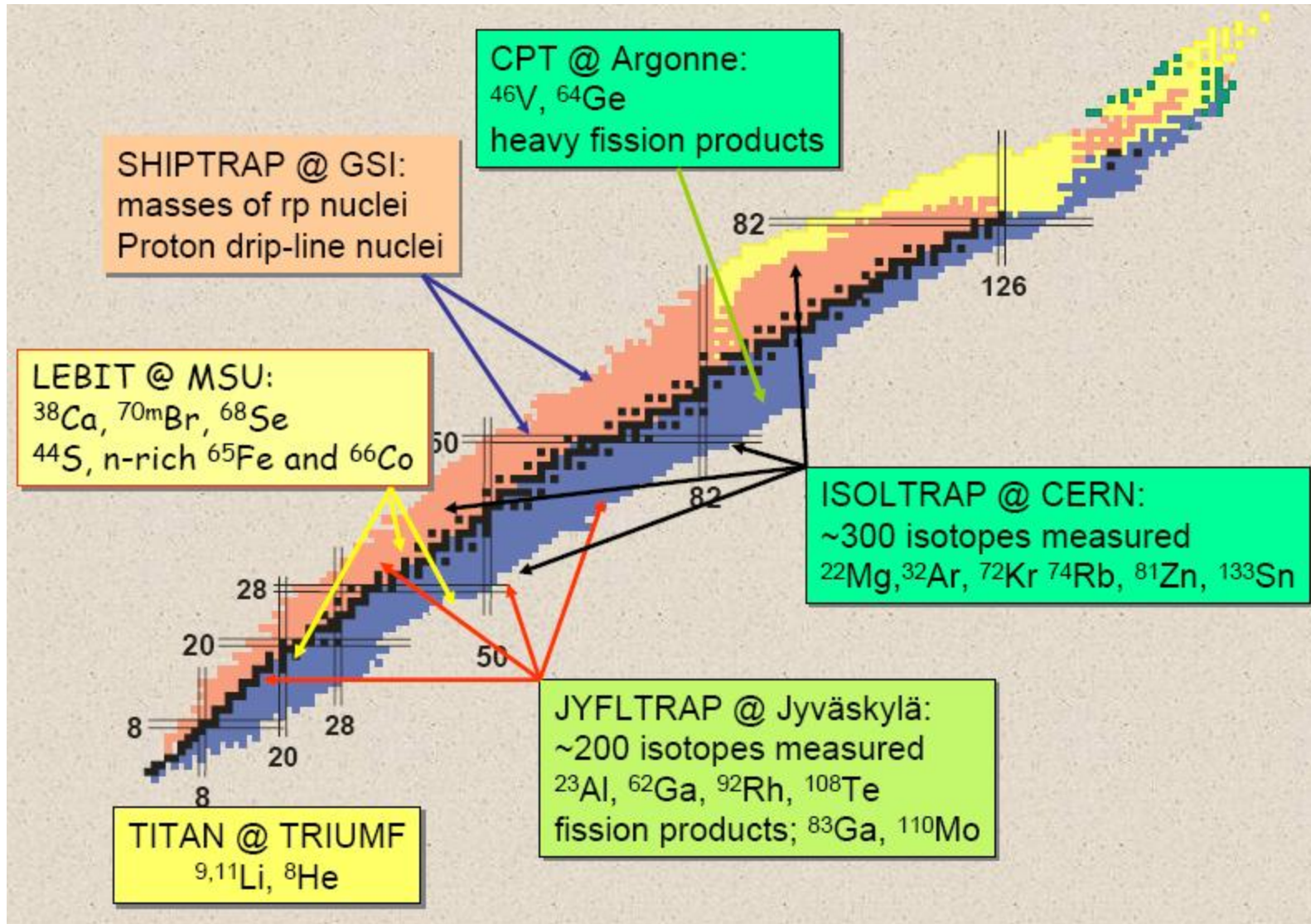


F. Ames et al., NPA 651, 3 (1999)

isoltrap.web.cern.ch/isoltrap/

Many other traps around the world, see e.g. EMIS 14
www.triumf.ca/emis14abs/abstracts.html#Topic06

Trap measurements – Overview





Masses – what are they good for?

- **Structure information**
 - Shell closures and deformation from separation energies ($\delta m/m < 10^{-5}$)
- **Astrophysics (Nucleosynthesis)**
 - r process ($\delta m/m < 10^{-5}$, $\delta m < 10$ keV)
 - rp process ($\delta m/m \sim 10^{-7}$)
- **Fundamental interactions and symmetries** ($\delta m/m < 10^{-8}$)
 - CVC
 - CKM



Theoretical description of masses

Algebraic: Garvey Kelson (**GK**), sum and difference relations between masses:

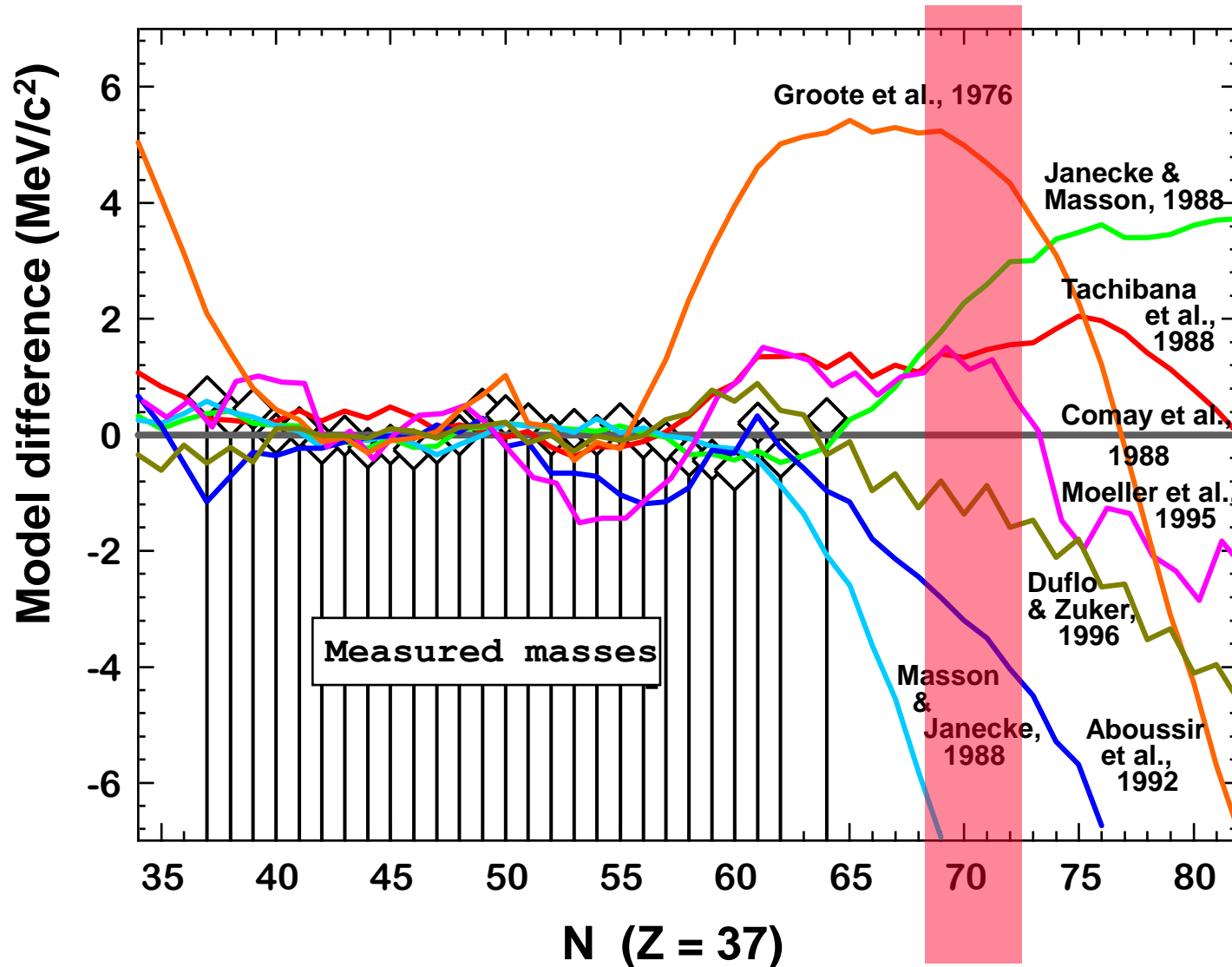
$$M(N+2, Z-2) - M(N, Z) + M(N, Z-1) - \\ M(N+1, Z-2) + M(N+1, Z) - M(N+2, Z-1) = 0$$

Microscopic-macroscopic: For example the finite-range droplet model **FRDM** (31 parameters, largely fit to known masses), bulk part from liquid drop model (macroscopic) + shell and pairing corrections (microscopic) (**extrapolation is dangerous**)

Microscopic: Relativistic mean-field (**RMF**) and Hartree-Fock Bogoliubov (**HFB**), RMF is based on meson/photon exchange Lagrangian, HFB uses Skyrme or Gogny effective nucleon-nucleon interactions (**computationally demanding**)

Masses – what are they good for?

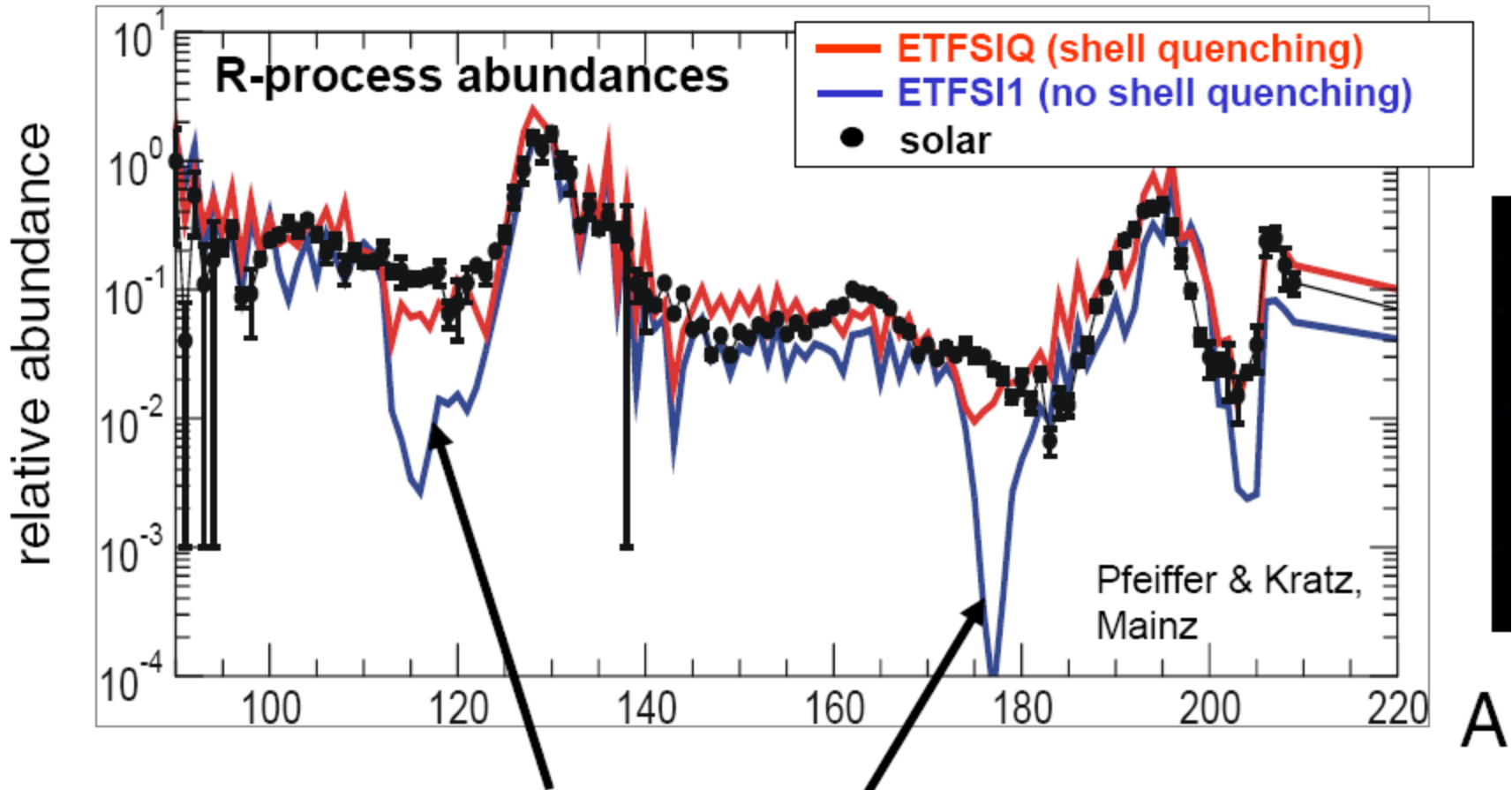
Constrain theory



Needed for r-process

Masses – what are they good for?

Nuclear astrophysics



Difference due to shell quenching for neutron-rich nuclei, or a problem with astrophysical model?



Masses – what are they good for?

Fundamental interactions/symmetries

Physics beyond the Standard Model

(required precision: as good as possible, at least: $\delta m/m < 10^{-8}$)

- Conserved vector current (CVC) hypothesis
- Unitarity of the Cabbibo-Kobayashi-Maskawa (CKM) matrix

Details: lecture by Tim Chupp



Takeaway

- Nuclear masses and resulting nucleon separation energies are an indicator of nuclear structure and indicate changes in the shell structure in the exotic regime
- Exotic nuclei can be produced with different methods and reactions
- Masses of short-lived nuclei can be measured in different ways
 - Time of flight mass measurements
 - Storage rings
 - Penning traps
- Masses are important input for nuclear astrophysics and the study of fundamental symmetries



Related review articles

Masses

- Traps for rare isotopes, G. Bollen, Lect. Notes Phys. 651, 169 (2004)
- Measurements of mass and beta lifetime of stored exotic nuclei, F. Bosch, Lect. Notes Phys. 651, 137 (2004)
- Recent trends in the determination of nuclear masses , D. Lunney, J.M. Pearson, C. Thibault, Rev. Mod. Phys. 75, 1021 (2003)
- Mass measurements of short-lived nuclides with ion traps, G. Bollen, NPA 693, 3 (2001)
- Precision nuclear measurements with ion traps, G. Savard and G. Werth, Annu. Rev. Nucl. Sci. 50, 119 (2000)
- Mass measurement far from stability, W. Mittig and A. Lepine-Szily, Annu. Rev. Nucl. Sci. 47, 27 (1997)