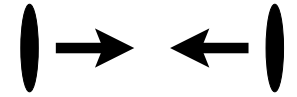


# QCD at High Temperature

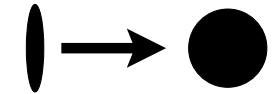
(Experiment)

# Kinematics

CMS:  $s = (p_1 + p_2)^2 = 4E_{CM}^2$



Lab:  $p_1 = (m, 0)$   $p_2 = (E_L, p_z) = (E_L, \sqrt{E_L^2 - m^2})$



$$s = (m + E_L)^2 - (E_L^2 - m^2) = 2m(E_L + m) \quad E_{CM} = \sqrt{mE_L/2}$$

SPS : 200 GeV (LAB)

$$E_{CM} = 10 \text{ GeV} \quad \gamma = 10$$

RHIC : 100 GeV (CMS)

$$E_{CM} = 100 \text{ GeV} \quad \gamma = 100$$

LHC : 2.75 TeV (CMS)

$$E_{CM} = 2.75 \text{ TeV} \quad \gamma = 2750$$

Rapidity:

$$y = \frac{1}{2} \log \left( \frac{E + p_z}{E - p_z} \right)$$

SPS :  $\Delta y = 6$     RHIC :  $\Delta y = 10.6$     LHC :  $\Delta y = 17.3$

## Bjorken Expansion

Experimental observation: At high energy ( $\Delta y \rightarrow \infty$ ) rapidity distributions of produced particles (in both pp and AA) are “flat”

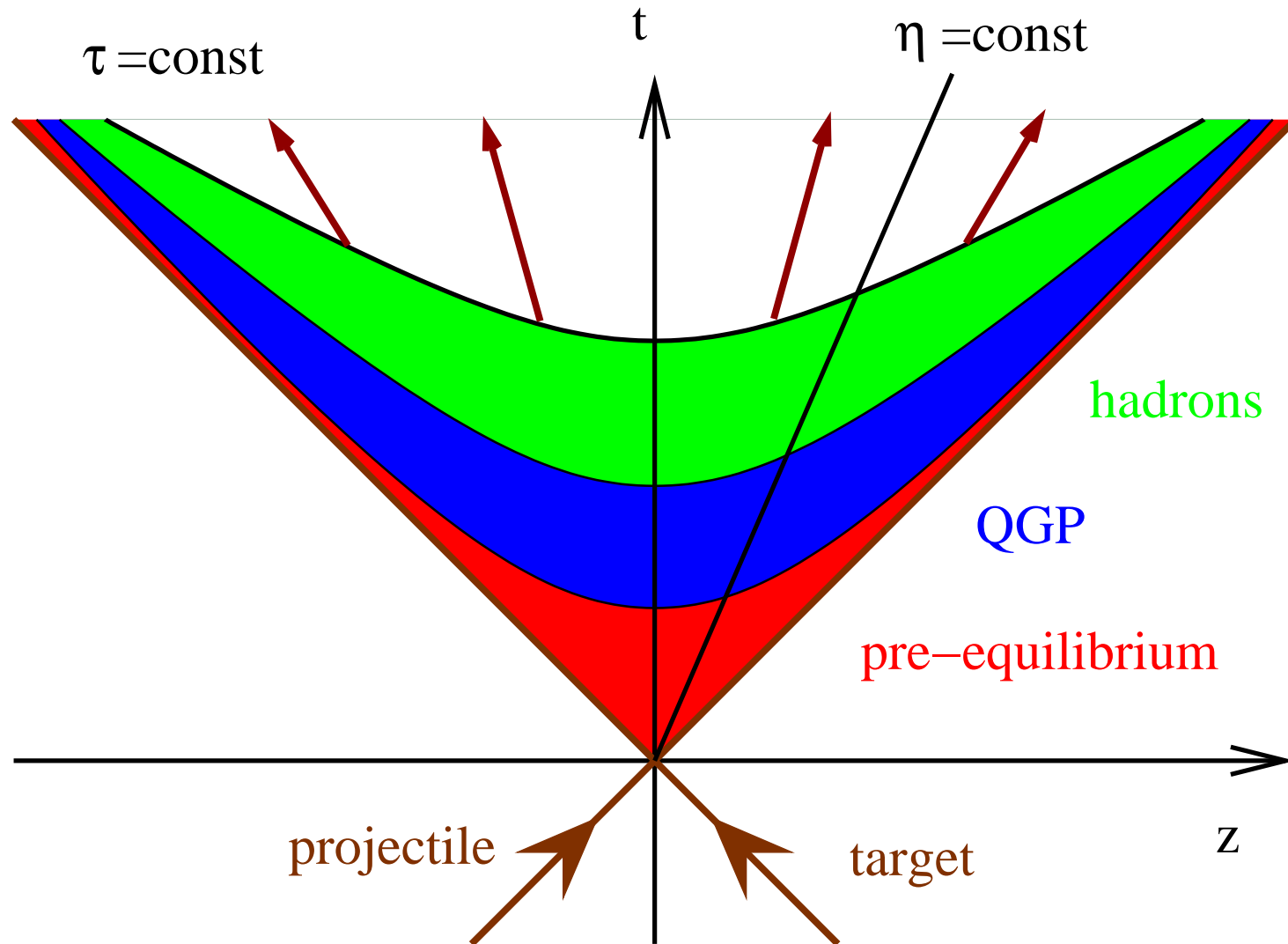
$$\frac{dN}{dy} \simeq \text{const}$$

Physics depends on proper time  $\tau = \sqrt{t^2 - z^2}$ , not on  $y$

All comoving ( $v = z/t$ ) observers are equivalent

Analogous to Hubble expansion

# Bjorken Expansion



# Bjorken Expansion: Hydrodynamics

Consider perfect relativistic fluid; 4-velocity  $u_\mu = (1, \vec{v})\gamma$

$$T_{\mu\nu} = (\epsilon + P)u_\mu u_\nu - P g_{\mu\nu}$$

Hydro = Conservation Laws ( $\partial^\mu T_{\mu\nu} = 0$ ) + Equ. of State ( $P = P(\epsilon)$ )

$$\partial^\mu T_{\mu\nu} = (\partial^\mu \epsilon + \partial^\mu P)u_\mu u_\nu + (\epsilon + P)((\partial^\mu u_\mu)u_\nu + u_\mu \partial^\mu u_\nu) - \partial_\nu P = 0$$

Contract with  $u_\nu$ , use  $u^2 = 1$

$$(\partial^\mu \epsilon + \partial^\mu P)u_\mu + (\epsilon + P)\partial^\mu u_\mu - u^\nu \partial_\nu P = 0$$

$$u_\mu \partial^\mu \epsilon + (\epsilon + P)\partial^\mu u_\mu = 0$$

Thermodynamic relations

$$d\epsilon = T ds$$

$$\epsilon + P = Ts$$

## Hydrodynamic equations

$$u^\mu (T \partial_\mu s) + (Ts) \partial^\mu u_\mu = 0$$

$$\boxed{\partial_\mu (s u^\mu) = 0} \quad \text{isentropic expansion}$$

Variables:  $t = \tau \cosh \alpha$ ,  $z = \tau \sinh \alpha$ .  $\Rightarrow u_\mu = (\cosh \alpha, 0, 0, \sinh \alpha)$

$$\partial^\mu (s u_\mu) = 0 \quad \Rightarrow \quad \frac{d}{d\tau} [\tau s(\tau)] = 0$$

Solution for ideal Bj hydrodynamics

$$\boxed{s(\tau) = \frac{s_0 \tau_0}{\tau}} \quad T = \frac{\text{const}}{\tau^{1/3}}$$

Exact boost invariance, no transverse expansion, no dissipation, ...

## Numerical Estimates

Total entropy in rapidity interval  $[y, y + \Delta y]$

$$S = s\pi R^2 z = s\pi R^2 \tau \Delta y = (s_0 \tau_0) \pi R^2 \Delta y$$

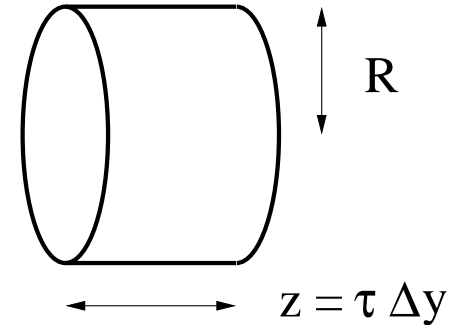
$$s_0 \tau_0 = \frac{1}{\pi R^2} \frac{S}{\Delta y}$$

Use  $S/N \simeq 3.6$

$$s_0 = \frac{3.6}{\pi R^2 \tau_0} \left( \frac{dN}{dy} \right)$$

$$\epsilon_0 = \frac{1}{\pi R^2 \tau_0} \left( \frac{dE_T}{dy} \right)$$

Bj estimate



Depends on initial time  $\tau_0$

RHIC: Au-Au collisions ( $\sqrt{s} = 200$  GeV)

$$\frac{dN}{dy} \simeq 998 \quad \tau_0 = 1 \text{ fm} \quad s_0 \simeq 33 \text{ fm}^{-3}$$

Use QGP equation of state  $s = 2g\pi^2 T^3 / 45$

$$T_0 \simeq 240 \text{ MeV} \quad \epsilon_0 \simeq (5 - 6) \text{ GeV}/\text{fm}^3$$

LHC: Factor  $\sim 2$  in multiplicity

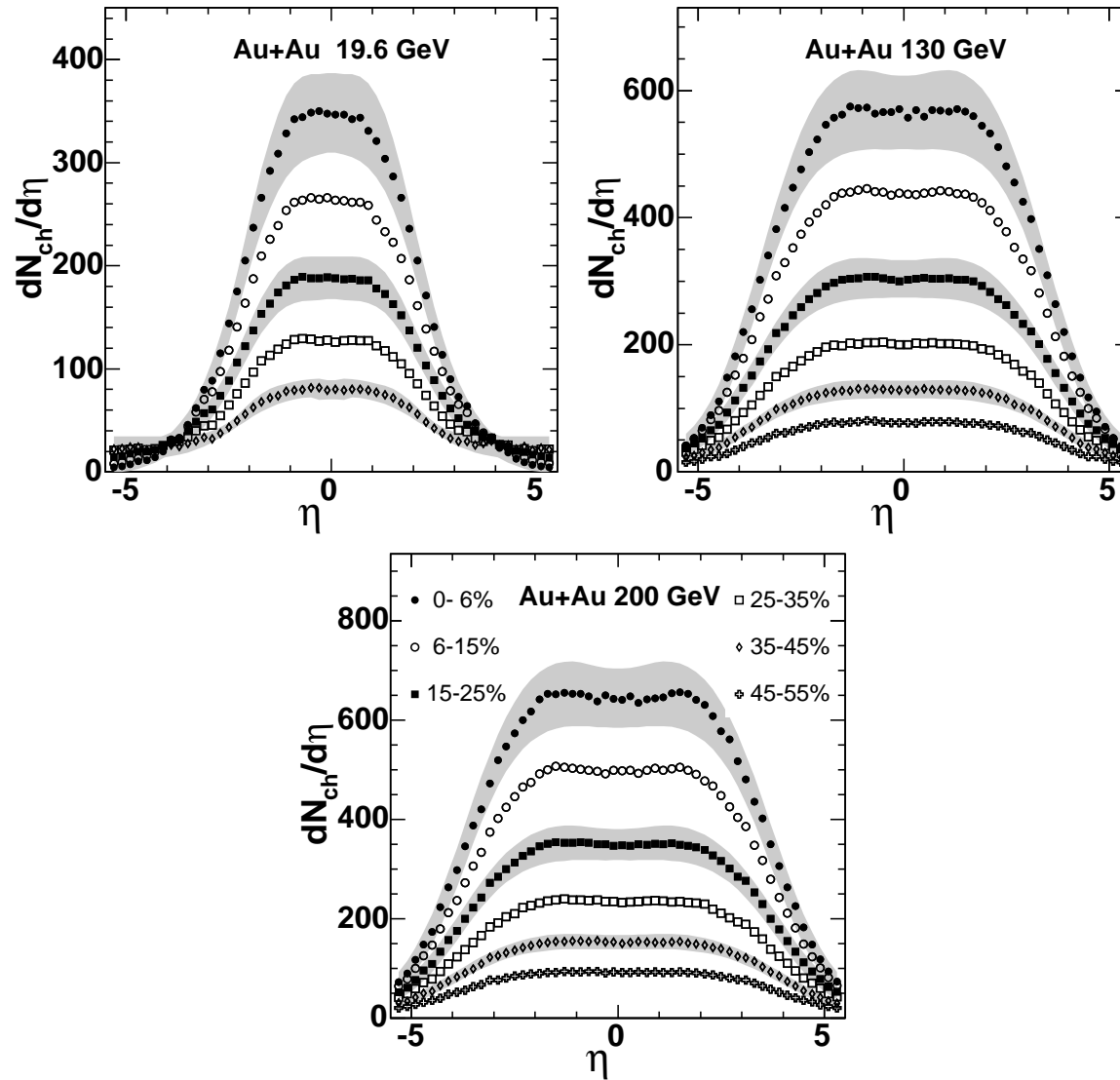
$$T_0 \simeq 300 \text{ MeV} \quad \epsilon_0 \simeq 15 \text{ GeV}/\text{fm}^3$$



# BNL and RHIC

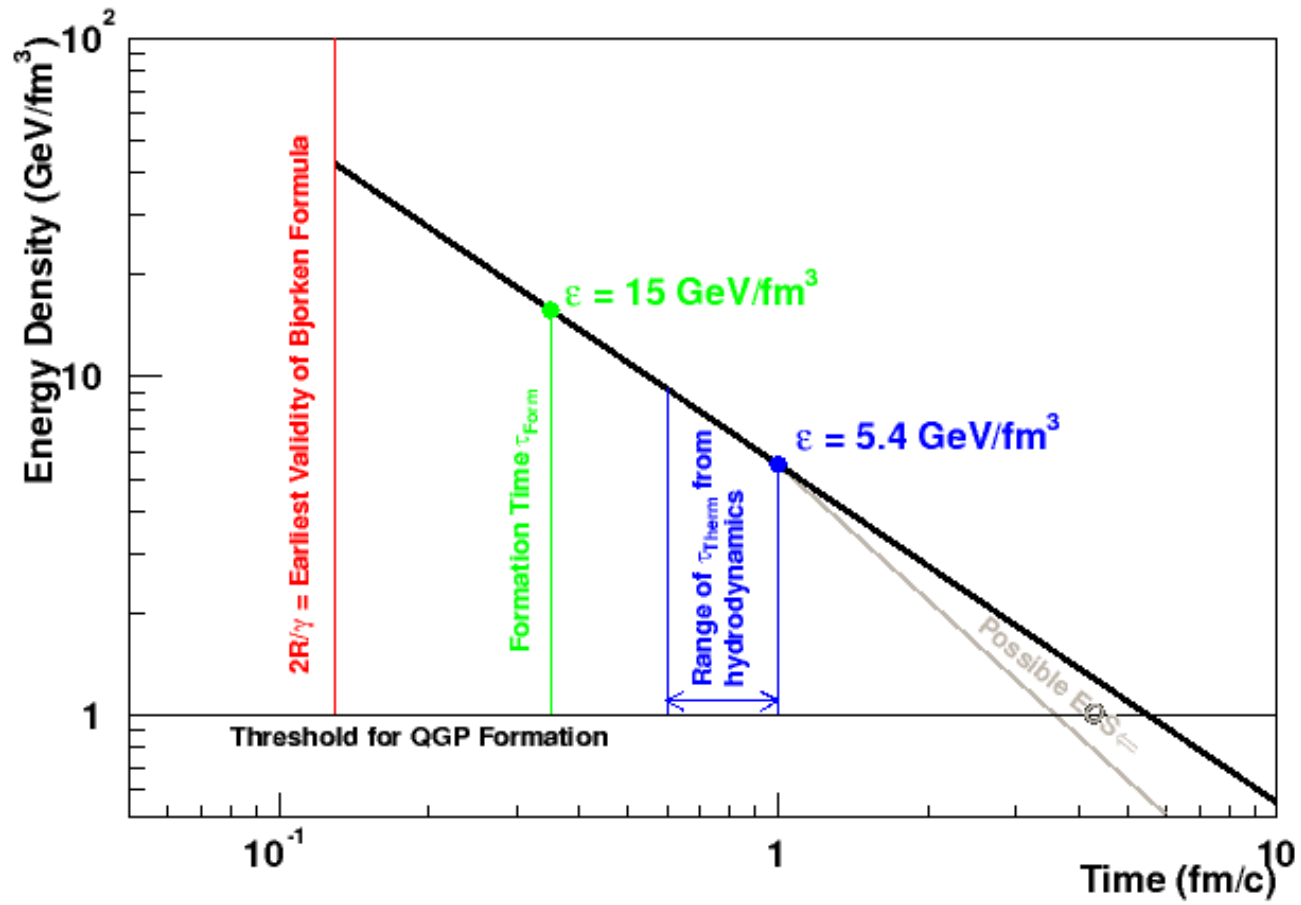


# Multiplicities



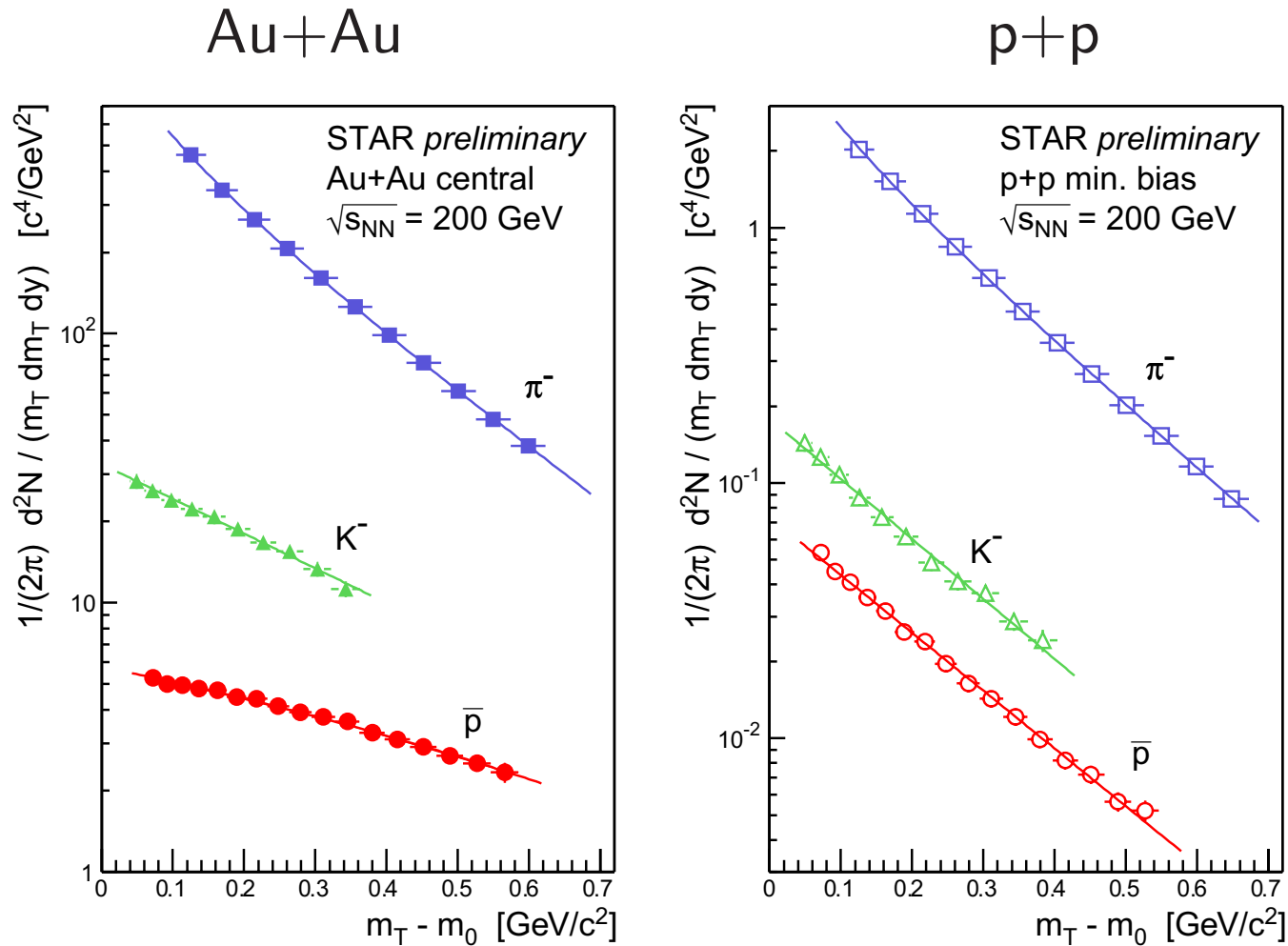
Phobos White Paper (2005)

# Bjorken Expansion



# Collective Behavior: Radial Flow

Radial expansion leads to blue-shifted spectra in Au+Au

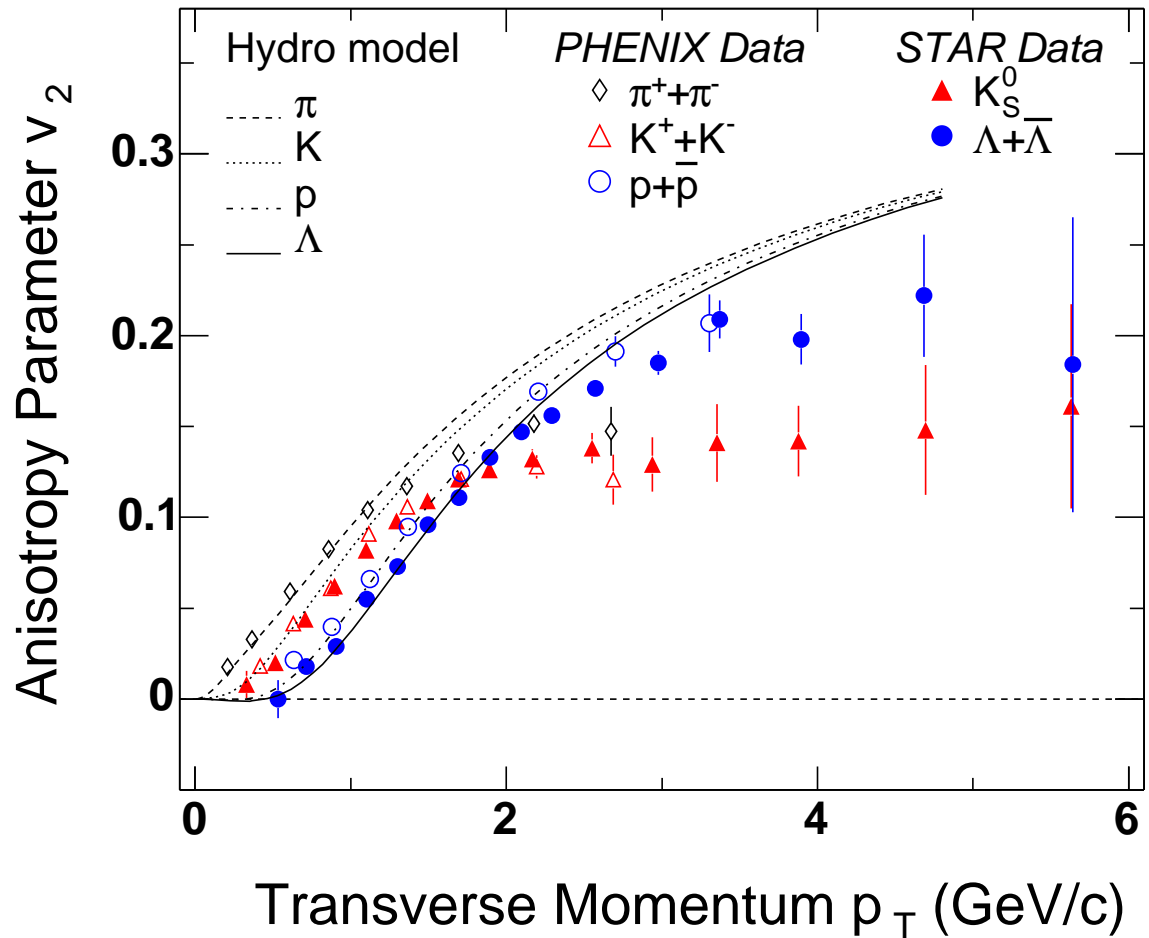
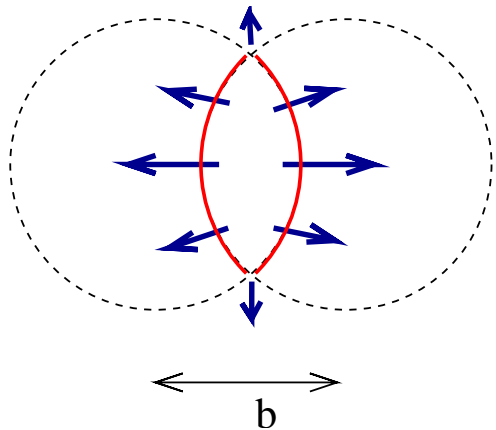


$$v_T \sim 0.6c!$$

$$m_T = \sqrt{p_T^2 + m^2}$$

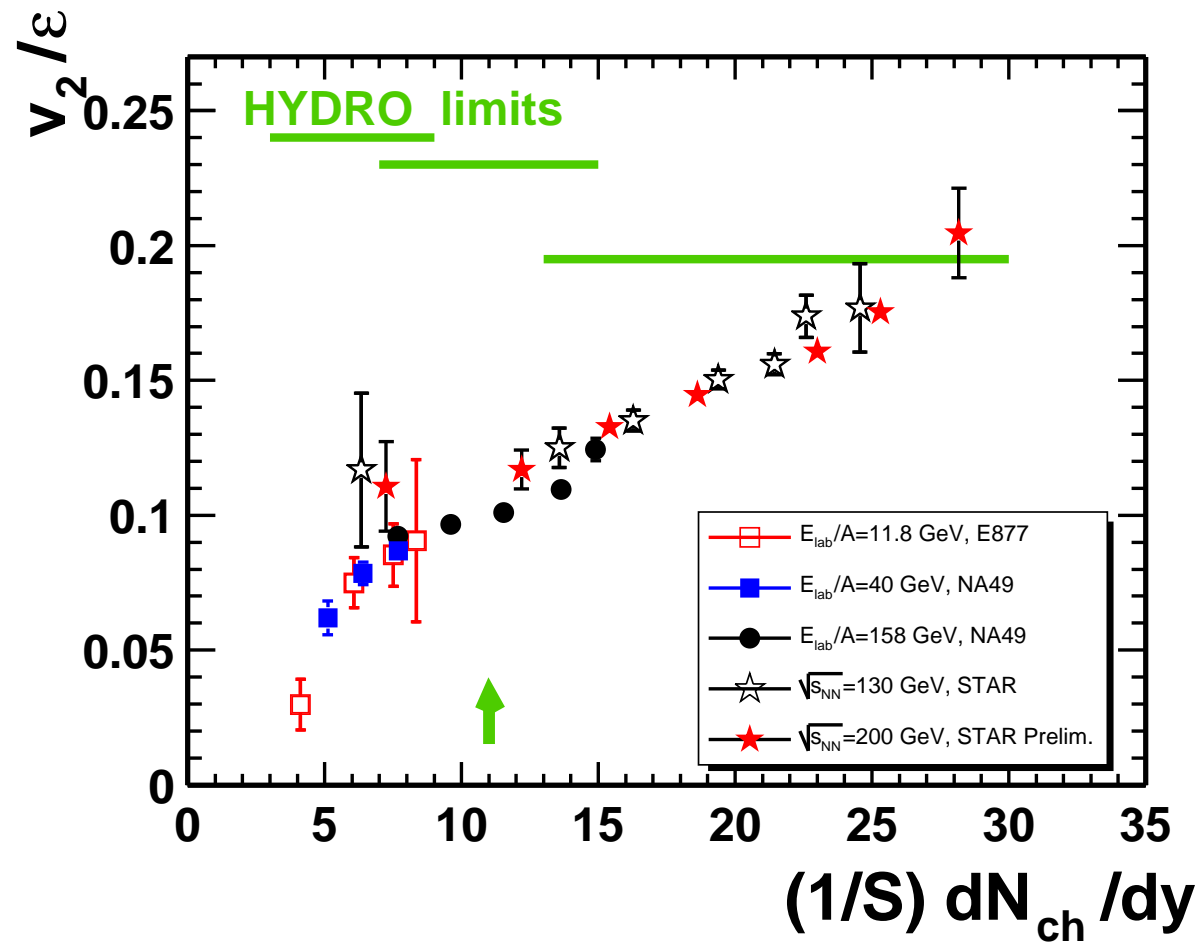
# Elliptic Flow

Hydrodynamic expansion converts  
 coordinate space  
 anisotropy  
 to momentum space  
 anisotropy



source: U. Heinz (2005)

# Elliptic Flow II



source: U. Heinz (2005)

# Elliptic Flow III: Viscosity

$$T_{\mu\nu} \rightarrow T_{\mu\nu} + \eta(\nabla_{\mu}u_{\nu} + \nabla_{\nu}u_{\mu} - \text{trace})$$

perturbative QCD

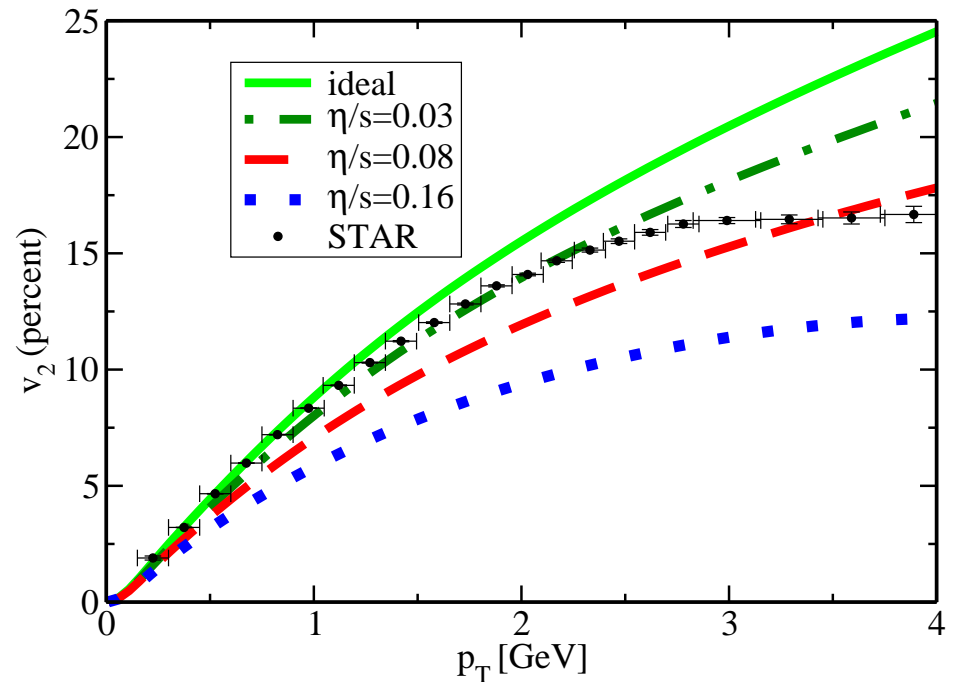
$$\frac{\eta}{s} = \frac{5.12}{g^4 \log(g^{-1})} \sim 1$$

Arnold, Moore, Yaffe

universal bound?

$$\frac{\eta}{s} \geq \frac{1}{4\pi}$$

Son, Starinets



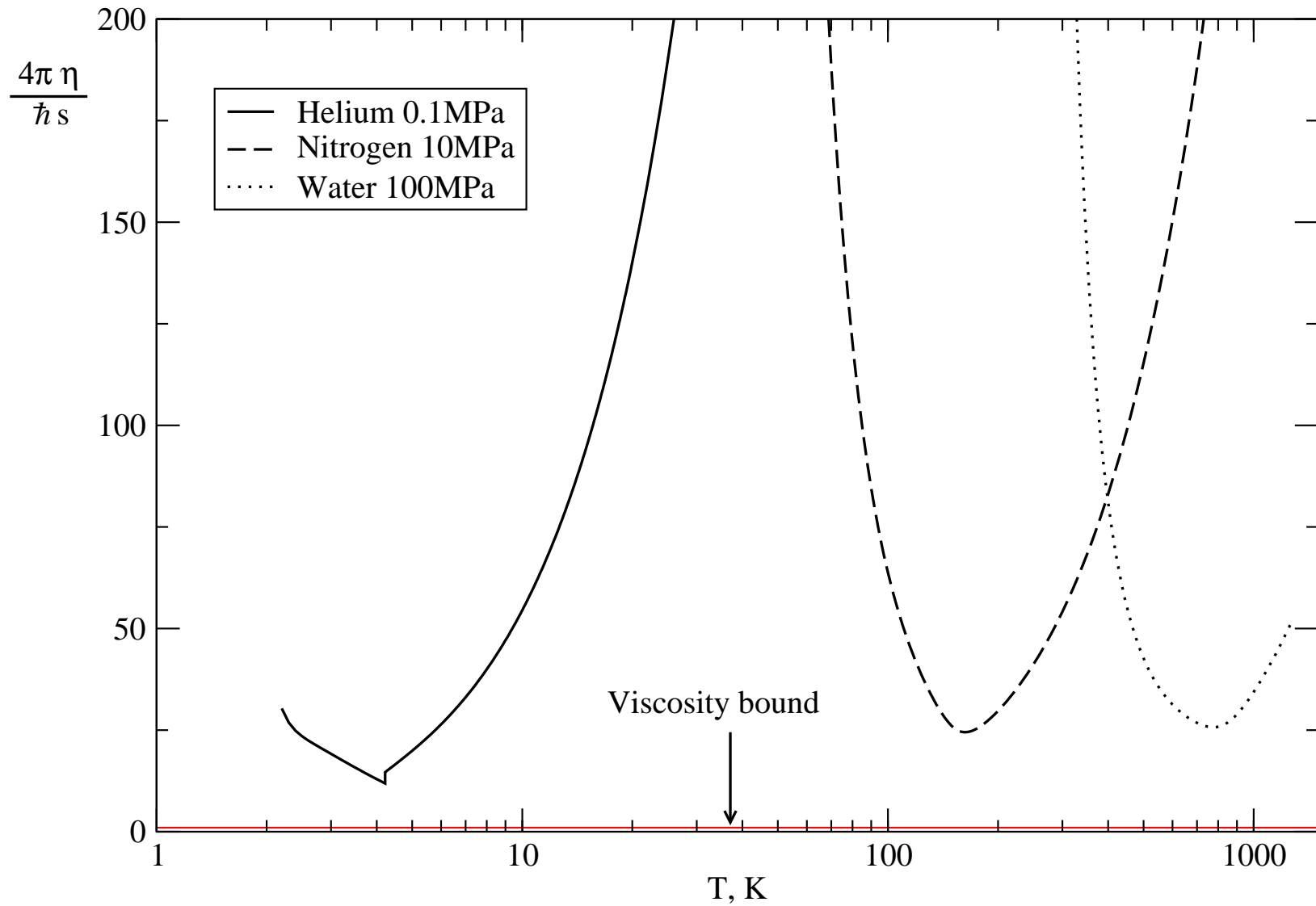
Romatschke (2007), Teaney (2003)

# A (Most) Perfect Fluid?



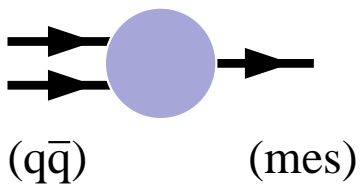
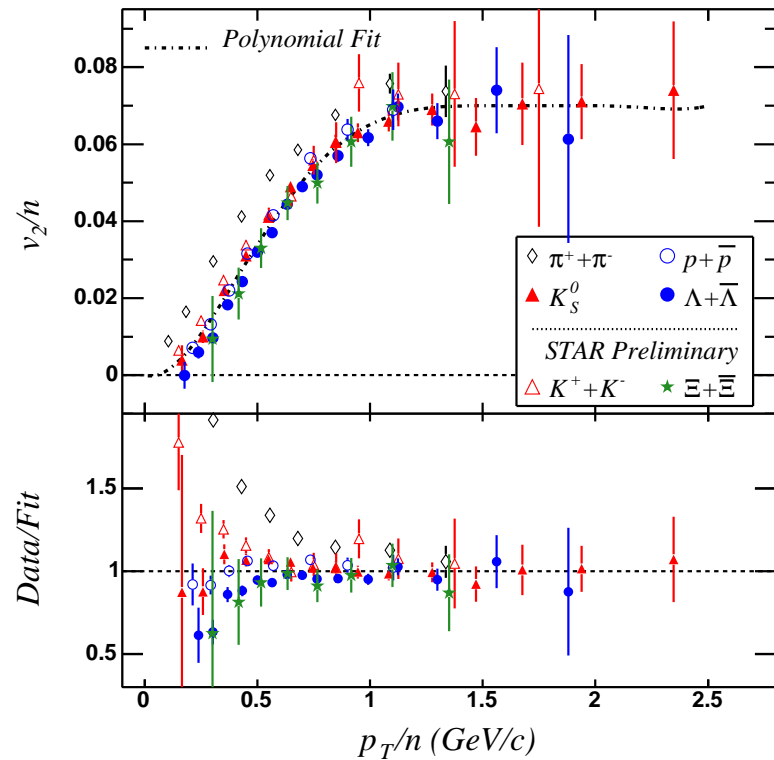
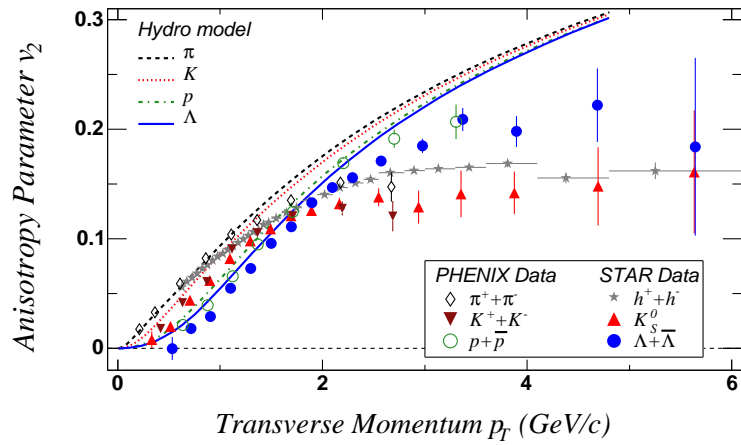


# A (Most) Perfect Fluid?

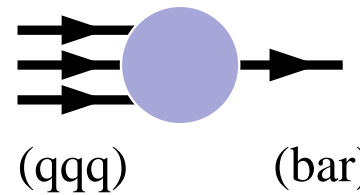


# Elliptic Flow IV: Recombination

“quark number” scaling of elliptic flow



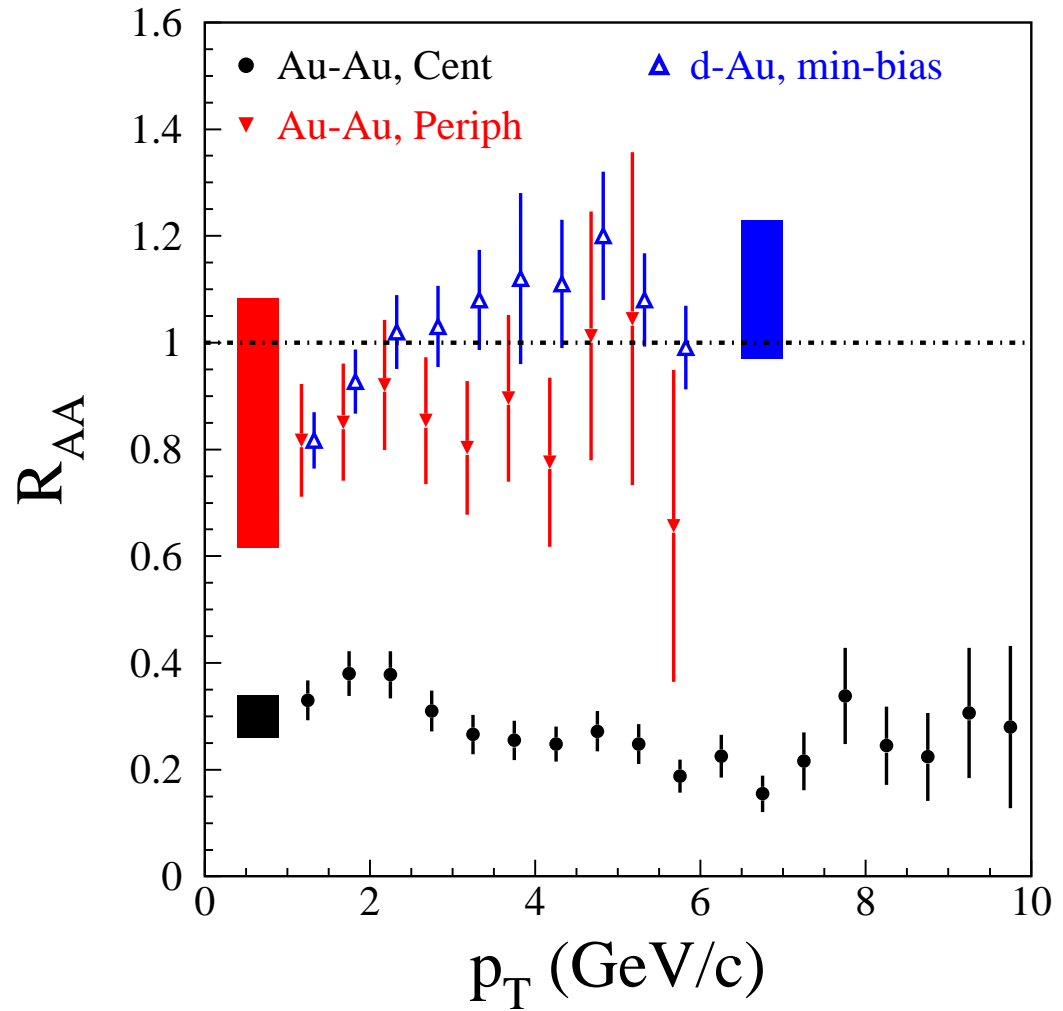
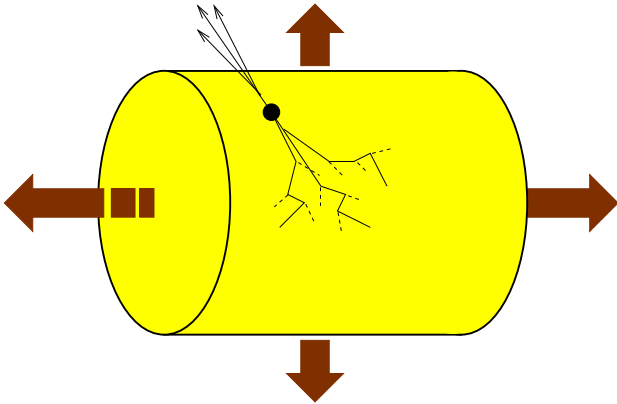
$$p_{\perp}^{mes} = 2p_{\perp}^{qu}$$



$$p_{\perp}^{bar} = 3p_{\perp}^{qu}$$

# Jet Quenching

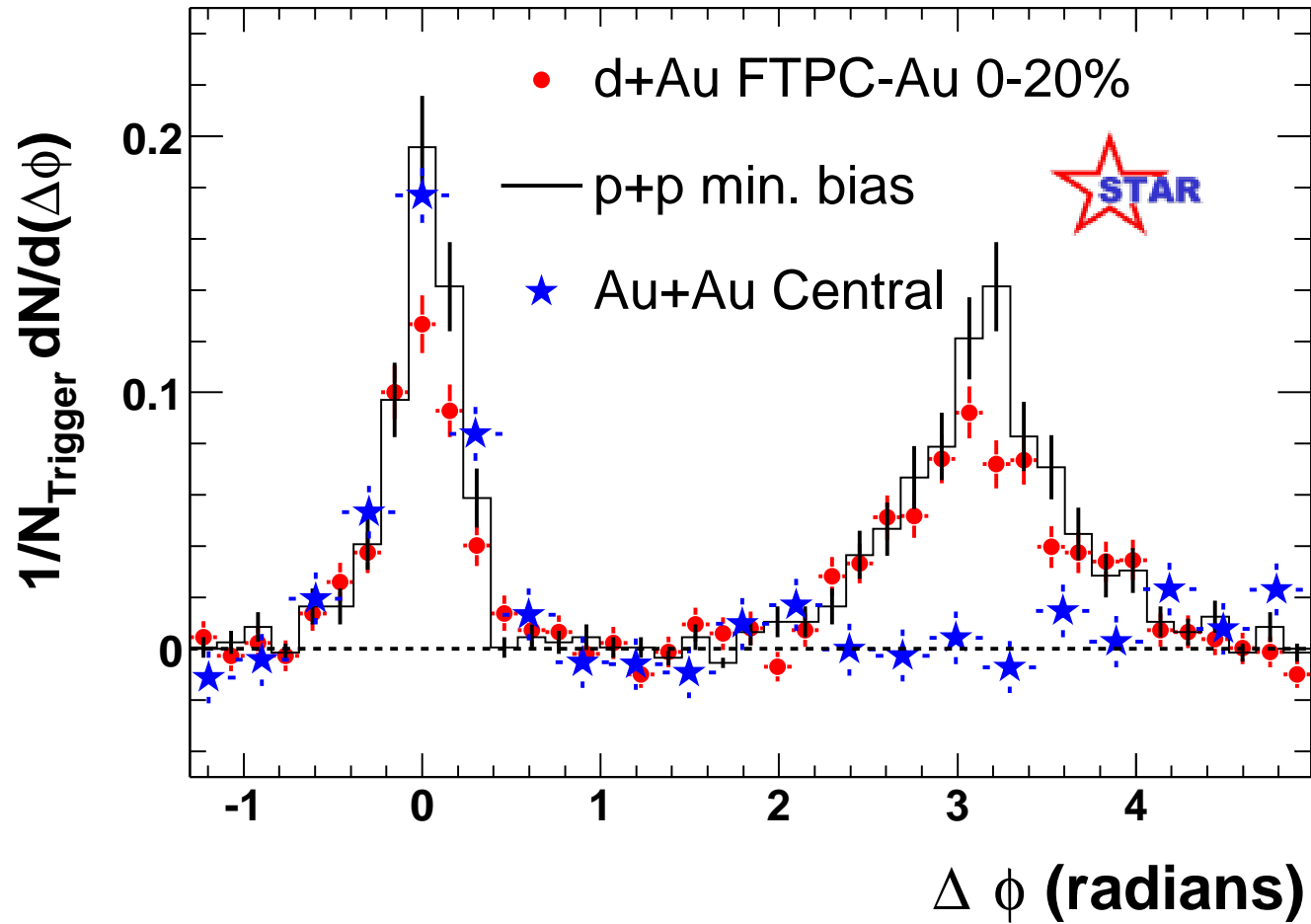
$$R_{AA} = \frac{n_{AA}}{N_{coll}n_{pp}}$$



source: Phenix White Paper (2005)

# Jet Quenching II

Disappearance of away-side jet

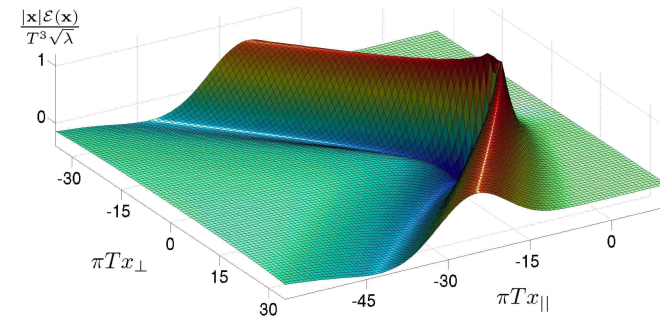
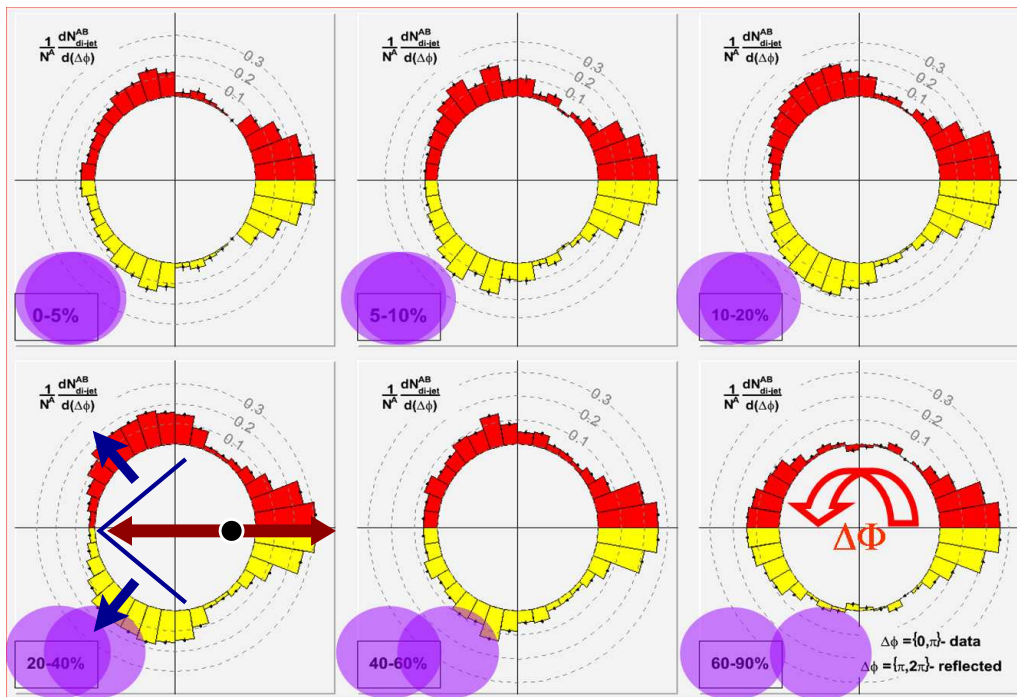


source: Star White Paper (2005)

# Jet Quenching III: The Mach Cone

azimuthal multiplicity  $dN/d\phi$   
 (high energy trigger particle at  $\phi = 0$ )

wake of a fast quark  
 in  $\mathcal{N} = 4$  plasma



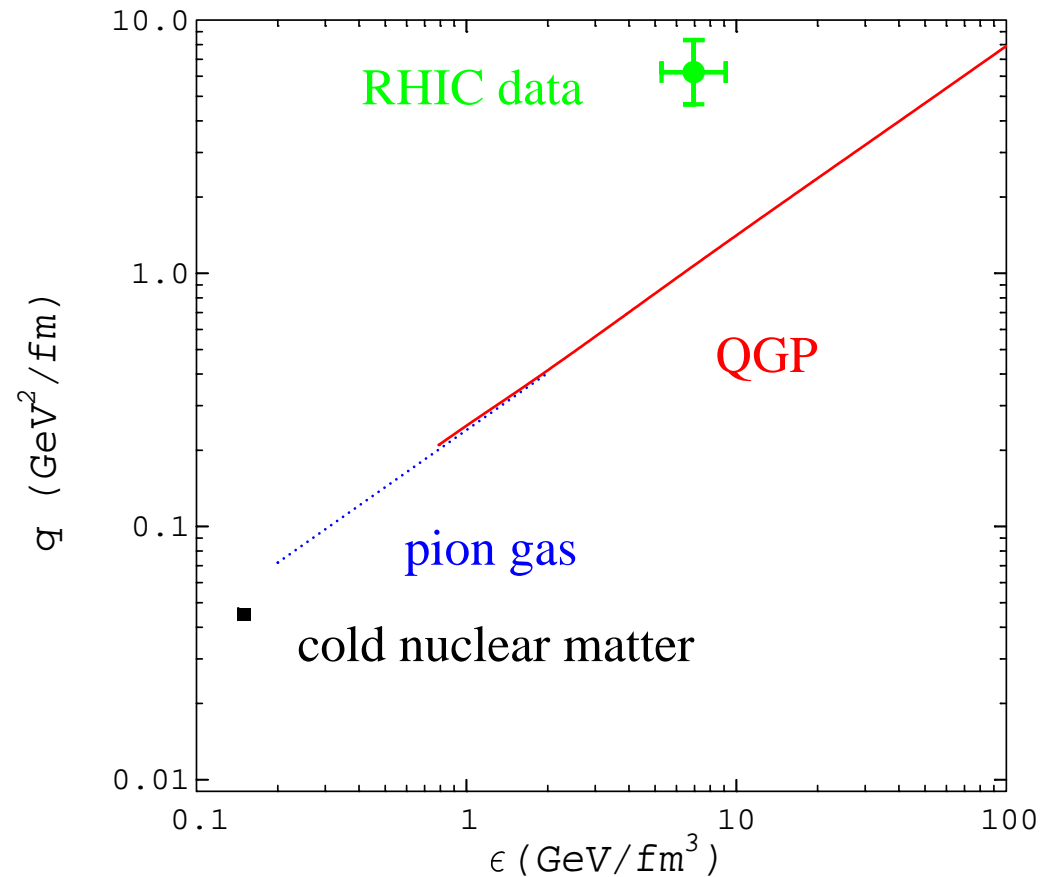
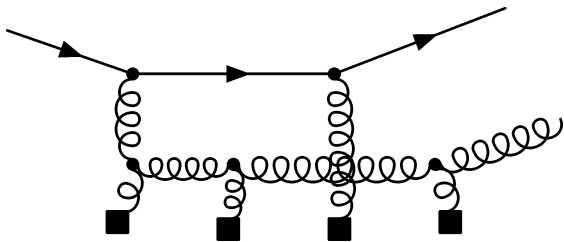
Chesler and Yaffe (2007)

source: Phenix (PRL, 2006), W. Zajt (2007)

# Jet Quenching: Theory

energy loss governed by

$$\hat{q} = \rho \int q_{\perp}^2 dq_{\perp}^2 \frac{d\sigma}{dq_{\perp}^2}$$

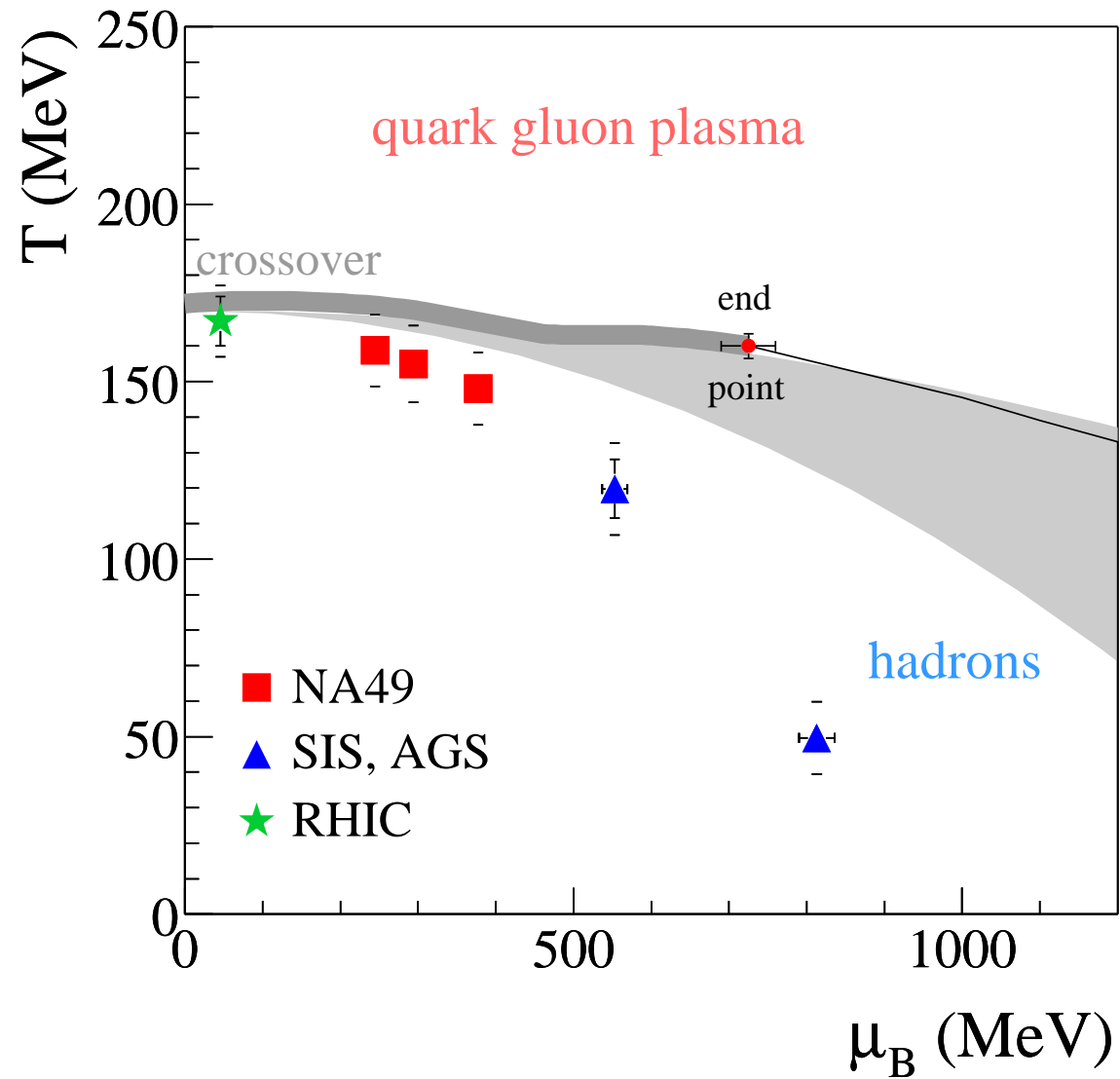


larger than pQCD predicts?

also: large energy loss of heavy quarks

[some recent doubts about  $\hat{q}$ , see P. Stankus seminar], source: R. Baier (2004)

# Phase Diagram: Freezeout



## Summary (Experiment)

Matter equilibrates quickly and behaves collectively

Little Bang, not little fizzle

Initial energy density in excess of  $10 \text{ GeV}/\text{fm}^3$

Conditions for Plasma achieved

Evidence for strongly interacting Plasma (“sQGP”)

Fast equilibration  $\tau_0 \ll 1 \text{ fm}$

Large elliptic flow, “perfect fluid”

Strong energy loss of leading partons



# The Future: LHC

