### A Solid Oxygen based Ultra-cold Neutron Source at IUCF

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Nuclear Physics Summer School, August 1, 200

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This work is supported by NSF

# Outline

- What are Ultra-cold Neutrons (UCN)?
- How are UCN useful?
  - Neutron physics with UCN
    - Neutron beta-decay
    - Neutron EDM search
    - n-nbar oscillation
- How to make a lot of them?
  - Superthermal UCN production
    - Solid D<sub>2</sub>
    - Solid O<sub>2</sub>

### UCN

- $E < 335 \text{ neV} (Ni^{58})$
- T < 4 mK
- Velocity < 8m/s
- $\lambda > 500$  Å





# Why UCN?

UCN have advantages over higher energy neutrons (cold neutrons):

- UCN can be confined in a trap
  - Copper wall ~ B=2.8 T ~ h=1.7m
- Low background



- Long storage time
  - UCN can be stored up to the β-decay lifetime, a relatively long coherence time of measurements (for particle physics experiments).
- 100% neutron polarization
  - Provide motivation to shift from cold neutron beams to UCN for β-decay angular correlation experiments and EDM experiments.

Clean, high precision experiments with reduced, well controlled systematic effects.

# **Neutron** β-decay



### Lifetime

- Cold Neutron beam experiments: =  $N_0 / \dot{N}_d$ 
  - Absolute measurements of the neutron number and the decay particle flux.
- Bottled UCN:  $N(T) = N_0 e^{-T/t_\beta} \Rightarrow t_\beta = T \ln(N_0 / N(T))$ 
  - Ratio of the neutrons stored for different periods. It is a relative measurement.
  - Material bottle -- Mampe (887.6  $\pm$  3 s)
    - Wall loss depends strongly on the UCN spectrum.
    - Systematically limited.
  - Magnetic bottle -- hexapole bottle (876.7  $\pm$  10 s), NIST bottle.

Neutron measurements which address fundamental particle physics issues

- Neutron β-decay lifetime and angular correlations test the V-A theory and place direct constraints on extensions to charged current sector of the standard model.
- Permanent electric dipole moment (EDM) search
   T reversal symmetry & CP violation extensions to the standard model.

#### • N-Nbar oscillation search

place useful limits on (B-L) violating processes.

Motivated by the observed baryon asymmetry of the universe.

# $n \rightarrow nbar (\Delta B=2)$ Theoretical **Motivations**

- Baryon Asymmetry of the Universe (BAU)
  - Sakharov Criteria
    - Baryon Number Violation
    - CP & C violation (EDM search)
    - Departure from Thermal Equilibrium



- Unification of particles and interactions.
  - Processes predicted by some GUT models.
  - Proton decay.  $\Delta B=1$  (=  $\Delta L$ )

 $p \to e^+ \pi^0 > 5.7 \cdot 10^{33}$  years  $p \rightarrow \overline{\nu}K^+ > 2.0 \cdot 10^{33}$  years

- In nucleon disappearance, the conservation of angular momentum leads to the selection rule:
  - $n \rightarrow \overline{n} \qquad \Lambda B = 2$  $-\Delta B = \pm \Delta L$  or  $|\Delta (B-L)| = 0,2$
  - In SM,  $\Delta$ (B-L)=0 always.

 $- \Delta(B-L) \neq 0 \rightarrow \Delta B = - \Delta L \text{ or } \Delta B = 2 \text{ or } \Delta L = 2$ 

 $p \rightarrow \nu \nu e^+ \Delta B = -\Delta L$ 

 $n \rightarrow \nu \nu \overline{\nu} \Lambda B = -\Lambda L$ 

# n→nbar Oscillation

Unkown mixing interaction

Schrödinger equation

 $i\hbar \frac{\partial}{\partial t} \begin{pmatrix} \psi_n \\ \psi_{\bar{n}} \end{pmatrix} = \begin{pmatrix} E_n & \varepsilon \\ \varepsilon & E_{\bar{n}} \end{pmatrix} \begin{pmatrix} \psi_n \\ \psi_{\bar{n}} \end{pmatrix} , \quad \begin{pmatrix} \psi_n(0) \\ \psi_{\bar{n}}(0) \end{pmatrix} = \begin{pmatrix} 1 \\ 0 \end{pmatrix}$ 

$$|\psi_{\overline{n}}(t)|^{2} = \frac{4\varepsilon^{2}}{\omega^{2} + 4\varepsilon^{2}} \sin^{2}\left(\frac{1}{2}\sqrt{\omega^{2} + 4\varepsilon^{2}}t/\hbar\right)$$
$$\omega = (E_{n} - E_{\overline{n}}) = (m_{n} + \frac{p^{2}}{2m_{n}} + V_{n}) - (m_{\overline{n}} + \frac{p^{2}}{2m_{\overline{n}}} + V_{\overline{n}})$$

• Transition Probability: (if  $\omega t \ll 1$ )

$$P_{n\to\bar{n}}(t) = \psi_{\bar{n}}(t)^2 = \varepsilon^2 \cdot (t/\hbar)^2 = \left(\frac{t}{\tau_{n\bar{n}}}\right)^2 \qquad \tau_{n\bar{n}} = \frac{\hbar}{\varepsilon}$$

# Suppression of n-hoar transition $|\psi_{\bar{n}}(t)|^{2} = \frac{4\varepsilon^{2}}{\omega^{2} + 4\varepsilon^{2}} \sin^{2}(\frac{1}{2}\sqrt{\omega^{2} + 4\varepsilon^{2}}t/\hbar)$

Free neutron in a magnetic field

$$\omega = (E_n - E_{\overline{n}}) = (m_n + \frac{p^2}{2m_n} + V_n) - (m_{\overline{n}} + \frac{p^2}{2m_{\overline{n}}} + V_{\overline{n}}) = 2\mu \cdot B$$

Under earth field (0.5 gauss), 2µHB=6×10<sup>-12</sup>eV

- 
$$\mathcal{E}_{n\overline{n}} = \frac{h}{\tau_{n\overline{n}}} < 10^{-23} eV$$
 with  $\tau_{\text{free}} > 1.2 \times 10^8 s$ 

$$|\psi_{\bar{n}}(t)|^2 = \frac{4\varepsilon^2}{\omega^2} \sin^2(t/\tau_{Larmor}) = 10^{-23} \sin^2(t/2 \times 10^{-4}) \checkmark$$

- To measure  $\tau_{free}$ >1.2×10<sup>8</sup> s, the magnetic field has to be as small as 0.5×10<sup>-11</sup>gauss!
- For the neutron time-in-flight t=0.1s,
  - B < 5 mgauss.

$$\varepsilon_{n\overline{n}} = \frac{\hbar}{\overline{t}} < 10^{-14} eV$$



Y. Kamyshkov

#### Typical Detector for the UCN N→Nbar Experiment



# Figure of Merit

- Probability of  $n \rightarrow nbar$  event in free neutron

exp:  

$$N \cdot P = (\dot{N}_{cn}T) \cdot \left(\frac{\bar{t}}{\tau_{n\bar{n}}}\right)^2 \le 1 \implies \tau_{n\bar{n}} \ge \left(\sqrt{\dot{N}_{cn}T}\right)(\bar{t})\varepsilon_{ff}$$

 $\tau_{n\bar{n}} > (\sqrt{1.25 \cdot 10^{11} \cdot 2.4 \cdot 10^7}) \cdot 0.109 \cdot 0.48 = 8.6 \times 10^7 s$ 

- Probability of UCn  $\rightarrow$  UCnbar event:
  - Every wall collision destroy the phase coherence and reset the experiment.

$$N \cdot P = (\dot{N}_{ucn}T) \cdot \left(\frac{\bar{t}}{\tau_{n\bar{n}}}\right)^2 \cdot \left(\frac{\tau_{ucn}}{\bar{t}}\right) \le 1 \quad \Rightarrow \quad \tau_{n\bar{n}} \ge \left(\sqrt{\dot{N}_{ucn}T}\right) \sqrt{\tau_{ucn}\bar{t}} \varepsilon_{ff}$$
  
$$\tau_{n\bar{n}} > (\sqrt{\dot{N}_{ucn} \cdot 3 \cdot 10^7}) \cdot \sqrt{500 \cdot 1} \cdot 1 = 8.6 \cdot 10^7 s \quad \Rightarrow \quad \dot{N}_{ucn} > 5 \times 10^5 / s$$
  
$$\tau_{n\bar{n}} > (\sqrt{\dot{N}_{ucn} \cdot 3 \cdot 10^7}) \cdot \sqrt{500 \cdot 1} \cdot 1 = 1.0 \cdot 10^9 s \quad \Rightarrow \quad \dot{N}_{ucn} > 6.6 \times 10^7 / s$$

# Technical Challenges with Experiments using UCN:

### **Need more UCN flux!**

### Neutron Cooling



Masuda, Proceedings of the 3<sup>rd</sup> UCN workshop

## **Superthermal Process**

R. Golub and J. M. Pendlebury, Phys. Lett, A53, 133 (1975)

• Cold neutrons downscatter in the solid, giving up almost all their energy, becoming UCN.



 UCN upscattering (the reverse process) is suppressed by cooling the moderator to low temperatures.

### Dynamics of UCN Production --Defeat thermal equilibrium



- Lifetime of UCN in the source material is a critical parameter in the establishment of large UCN densities.
- Extract UCN out of the source before it is thermalized ⇒ Spallation N source + Seperation of the source and the storage + a UCN Valve

### Superfluid <sup>4</sup>He – UCN production

Isotropic superfluid <sup>4</sup>He

 Energy excitation is isotropic.
 Neutron scattering is isotropic.



• UCN can accumulate until the production rate = loss rate

$$\rho_{ucn} = P \times \tau = \left( \Phi_0 \sigma_{down} \right) \left( \frac{1}{n \sigma_{up} v} \right) \propto \frac{\sigma_{down}}{\sigma_{up}} = \frac{1 + n(\omega)}{n(\omega)} \sim \exp(\omega / T)$$
$$n(\omega) = \frac{1}{\exp(\omega / T) - 1}$$
Superthermal gain

### Superfluid <sup>4</sup>He – UCN loss

- UCN production rate:  $P = 7.2 \frac{d^2 \Phi}{d\lambda d\Omega} \frac{1}{\lambda_{wall}}$  UCN/cm<sup>3</sup>Hsec
- UCN density:  $\rho_{ucn} = P \times \tau \propto \sigma_{down} \left( \frac{1}{\sigma_{up}} + \frac{1}{\sigma_{\beta}} + \frac{1}{\sigma_{nucl.ab.}} + \dots \right)$
- The figure of merit:  $\sigma_s / \sigma_a$

Isotop	$\sigma_{\!\scriptscriptstyle coh}$	$\sigma_{inc}$	$\sigma_{a}$	$\sigma_s/\sigma_a$	purity	Debye T
<sup>2</sup> D	5.59	2.04	0.000519	$1.47 \times 10^{4}$	99.82	110
<sup>4</sup> He	1.13	0	0	$\infty$		20
<sup>15</sup> N	5.23	0.0005	0.000024	2.1×10 <sup>5</sup>	99.9999	80
<sup>16</sup> O	4.23	0	0.00010	2.2×10 <sup>4</sup>	99.95	104
<sup>208</sup> Pb	11.7	0	0.00049	$2.38 \times 10^4$	99.93	105

Solid Deuterium –UCN production (I)

- Incoherent contribution ( $\sigma_{inc} = 2.04$  barn) (due to the difference of singlet and triplet scattering)
  - No momentum delta function in the scattering cross section.

$$\sum_{q} \rightarrow \int d\omega Z(\omega)$$

All the Cold Neutron with energy smaller than the Debye T could become UCN through incoherent phonon creation.



### Solid D2 – UCN production (II)

- Coherent contribution ( $\sigma_{coh} = 5.59$  barn)
  - Momentum and energy conservations are still strictly hold.
  - The anisotropic dispersion relation broadens the range of conditions for single phonon creation process.
     In a cold neutron flux with a continuous spectrum, more neutrons could participate in the UCN production.





### Solid Deuterium - UCN Loss



# LANL UCN prototype source (2000)

### UCN lifetime in S-D<sub>2</sub>



C. Morris et al., Phy. Rev. Lett. 89, 272501 (2002)

- Superthermal temperature dependence.
- Para-D2 upscattering time:  $1.2 \pm 0.2$  ms.

### Volume Scan

# LANL UCN prototype source (2000)





UCN yield saturates above 200 c.c ⇒ mean free path = 8 cm

Resulted from UCN incoherent elastic scattering (random walk).

• No additional scattering due to the finite crystal effects.

### UCN Production Measurement --Bottle Technique LANL UCN prototype source



#### C. Morris *et al.*, Phy. Rev. Lett. 89, 272501 (2002) Los Alamos s-D<sub>2</sub> UCN Prototype Source



- Source has para-D<sub>2</sub>: 4%
- Bottled UCN density: 100 UCN/c.c. in a S.S. bottle 1 m away from the source. (world record)
- UCN Flux = 3.8×10<sup>4</sup> UCN/s
- Noticeable beam heating on solid deuterium.

### **Source Candidates**

Isotope	$\sigma_{\rm coh}$	$\sigma_{inc}$	$\sigma_{abs}$	$\sigma_{tot}^{\prime} \sigma_{abs}^{\prime}$	purity	T <sub>Debye</sub>
<sup>2</sup> D	5.59	2.04	5.2e-4	<b>1.47e+4</b>	99.82	110
<sup>4</sup> He	1.13	0	0	8		20
$^{15}$ N	5.23	5e-4	2.4e-5	2.1e+5	99.9999	80
<sup>16</sup> O	4.23	0	1.0e-4	2.2e+4	99.95	104
<sup>208</sup> Pb	11.7	0	4.9e-4	2.4e+4	99.93	105

**Too Heavy !** 

**Too Expensive !** 

### Solid Oxygen as a UCN Source

- Electronic spin S=1 in O<sub>2</sub> molecules
- Nuclear spin = 0 in <sup>8</sup>O
- Anti-ferromagnetic ordering α-phase, T < 24K.</li>

P.W. Stephens and C.F. Majkrzak, Phys. Rev. B 33, 1 (1986)

#### **UCN Production in S-O<sub>2</sub>**

- Produce UCN through magnon excitations.
  - Magnetic scattering length ~ 5.4 fm.
- Null incoherent scattering length.
- Small nuclear absorption probability.



 $\Rightarrow$  A very large source possible.

# Neutron Scattering in Solid O<sub>2</sub>

Spin(n) -Spin(e) coupling

$$V(r) = -\mu_N \cdot H = -\gamma \mu_N \sigma \cdot \left\{ \nabla \times \frac{\mu_e \times r}{r^3} \right\}$$

$$V(k) = \gamma_0 \sum_{l} \sigma \cdot \widetilde{k} \times (\widetilde{S}_l \times \widetilde{k}) e^{ik \cdot r_l}$$

(Spin)×(Translation)

1

Elastic Bragg + Magnon Scatt. + Magneto-vibrational Scatt. + both magnon, phonon

### Neutron-Magnon Scattering in S-O<sub>2</sub>

Density of States (Magnon)



- AF magnon: scattering amplitude prefers low momentum transfers.
  - Needs a colder neutron spectrum for the optimum UCN production.
- Magnon production energy gap ~ 0.8meV
  - Magnons are partially frozen at T < 8K.</li>
  - Significantly reduce the UCN upscattering rate.

### UCN production in Solid Oxygen



### Low Energy Neutron Source (LENS) at IUCF (floor plan-2007)



### Target Moderator Reflector (TMR)

#### LENS, IUCF



#### Target/Moderator/Reflector (TMR) Assembly



### **Neutron Production at LENS**

**Neutron Source:** 

based on low-energy (p,n) and (p,pn) reactions ( $E_p < 14$ MeV) in Be.

Yield ~ 0.01 n/p @13 MeV ~ 0.002 n/p @ 7MeV



#### **Time line:**

• Phase I (Early 2005: 7MeV 10mA, 1% DF; **10<sup>12</sup> n/s**).

• Phase II (Fall 2006: 7Mev, 50mA, 5% DF; **2x10<sup>13</sup> n/s**)

– Eventual power (13MeV, 50mA, 5% DF; **10<sup>14</sup> n/s**)

#### LENS produces its first "cold" neutron beam

*Mark Leuschner, et. al.* April 26, 2005

- methane moderator (1 cm thick, 3.6 K.)
- two component spectrum (25 K, 157 K.)
- The moderator is very thin, so the neutrons that pass through it do not quite reach thermal equilibrium.
- Spin temperature is too high. (YunChang Shin)



 $4 \mathrm{K} \mathrm{CH}_4$ 



# UCN source at LENS

#### • Features

- (p,n) reactions make fewer fast neutrons and gammas in source  $\Rightarrow$  might be able to run CW. (Has the benefit of a reactor source)
- LENS will have a variable pulse width (from  $<5 \ \mu s$  to 1.0 ms or more).
- In long-pulse mode, LENS will have a time-averaged cold neutron intensity comparable to an existing national user facility (IPNS).

#### Cold moderators (4~22K)

- Rad damage from source is low for LENS
- Kr-doped CH4 (free rotor), 2-methyl-pyridine (CH3 free rotor),

#### • UCN flux estimate (in S-O<sub>2</sub>)

- Cold neutron flux:  $2 \times 10^{10}$  CN/cm<sup>2</sup>-s (Patrick McChesney) (proton: 13 MeV, 2.5mA(avg), with 4K poly moderator, T<sub>CN</sub>=35K)
- UCN density: 300 UCN/cc, UCN fluence: > 3 ×10<sup>5</sup> UCN/s (with 500 cc S-O<sub>2</sub>)
- Heat load ~ 1 W

# Challenges



#### • 1.8 K cryogenic.

Larger gamma heating and smaller thermal conductivit,

than  $S-D_2 \Rightarrow$  challenges on cryogenic engineering

Requires a fast thermal break (50K to 2K) over a few cm.

#### Shortened UCN mean free path

- **Polycrystalline sample formation**  $\Rightarrow$  Effects on UCN mean free path?
- Ozone formation ⇒ additional incoherent scattering, resulting in a reduced mean free path. (Low radiation level at LENS helps.)

### Probe the Magnon Mechanism using a B field



C. Uyeda at. al., J. Phys. Soc. Jpn. 54, 1107 (1985)

- An external magnetic field to perturb the magnon dispersion curve
  - Change the density of states.
  - Optimize UCN production.
- Spin flop transition around 7 Tesla  $\Rightarrow$  AF magnons turned off.
- Definitive demonstration of the magnon mechanism.

### Experiments of UCN Production from S-O<sub>2</sub>

- PSI, Switzerland
- FunSpin beamline in SINQ
  - $\Phi_{CN} = (4.5 \pm 1.0) \times 10^7 / \text{cm}^2 \text{s-mA}$
  - with 1.2mA on SINQ target
  - PSI UCN group has demonstrated UCN production in s-D<sub>2.</sub>
  - UCN count rate in the detector of 0.4/s, with a S/N ratio of 40 to 1.

F. Atchison, et al., Phys. Rev. C 71, 054601 (2005)

- Similar count rate is expected using solid oxygen
  - assuming the UCN extraction is not hindered by the very different physical properties of oxygen solid (compared with s-D<sub>2</sub>).

### Cold Neutron Transmission (TOF)

Scattering Probability in O<sub>2</sub> 0.8 Liquid O, @55K Probability 6.0 γ-O<sub>2</sub>@50K α-0**,@**8K 0.2 β-O**,@30**K 10<sup>-3</sup> 10<sup>-2</sup>  $10^{-1}$ 10 Energy (eV)

- Flight path =2.83m.
- Neutron Chopper.
- Scattering probability
  - $I_0(E)-I(E)/I_0(E)$
- Features:
  - Bragg edges
  - Additional Bragg peak in alpha
     phase. (indicate the presence of a magnetic structure.)

# UCN Production in Solid O<sub>2</sub>



- No superthermal temperature dependence.
  - Indicates unknown source of UCN loss.
- UCN yield is correlated with how the crystal is prepared.
- The UCN yield (best number) is ~ 3 times less than s-D<sub>2</sub>.
- A peak in the  $\alpha$ - $\beta$  phase transition. (critical scattering?)

# UCN Production in D<sub>2</sub> and CD<sub>4</sub>

- From  $D_2$  and  $CD_4$ .
- Signature temperature dependence of a superthermal source.





### Conclusions

- Magnons in the AF phase of S-O<sub>2</sub> offer an additional channel for inelastic neutron scattering.
  - UCN production rate in S-O<sub>2</sub>~ (1-2) × in S-D<sub>2</sub>.
  - UCN lifetime in S-O<sub>2</sub> ~  $10 \times$  in S-D<sub>2</sub>.
  - Larger source possible. (at least  $10 \times S-D_2$ )
  - UCN current output from S-O<sub>2</sub> (at least)  $100 \times \text{from S-D}_2$
- UCN Program at IUCF
  - LENS provides a unique opportunity to study and develop a S-O<sub>2</sub> based UCN source.
  - Currently, we are setting up to study crystal growth. (YunChang Shin, John Ullman)
  - Planning to study magnetic field influence on the crystal preparation, energy dispersion curves, etc..
- Broader impacts
  - A high UCN flux will make a UCN nnbar experiment possible.