

Nuclear Searches

for Physics Beyond the Standard Model

- Experimental Searches
 - What are the experiments
 - Where to do experiments
- Nuclear and neutron beta decay
- Neutrons for Searches Beyond the Standard Model
 - Ultra-Cold Neutrons (UCN)

Neutral Weak Phenomena - $\sin^2\theta_w$

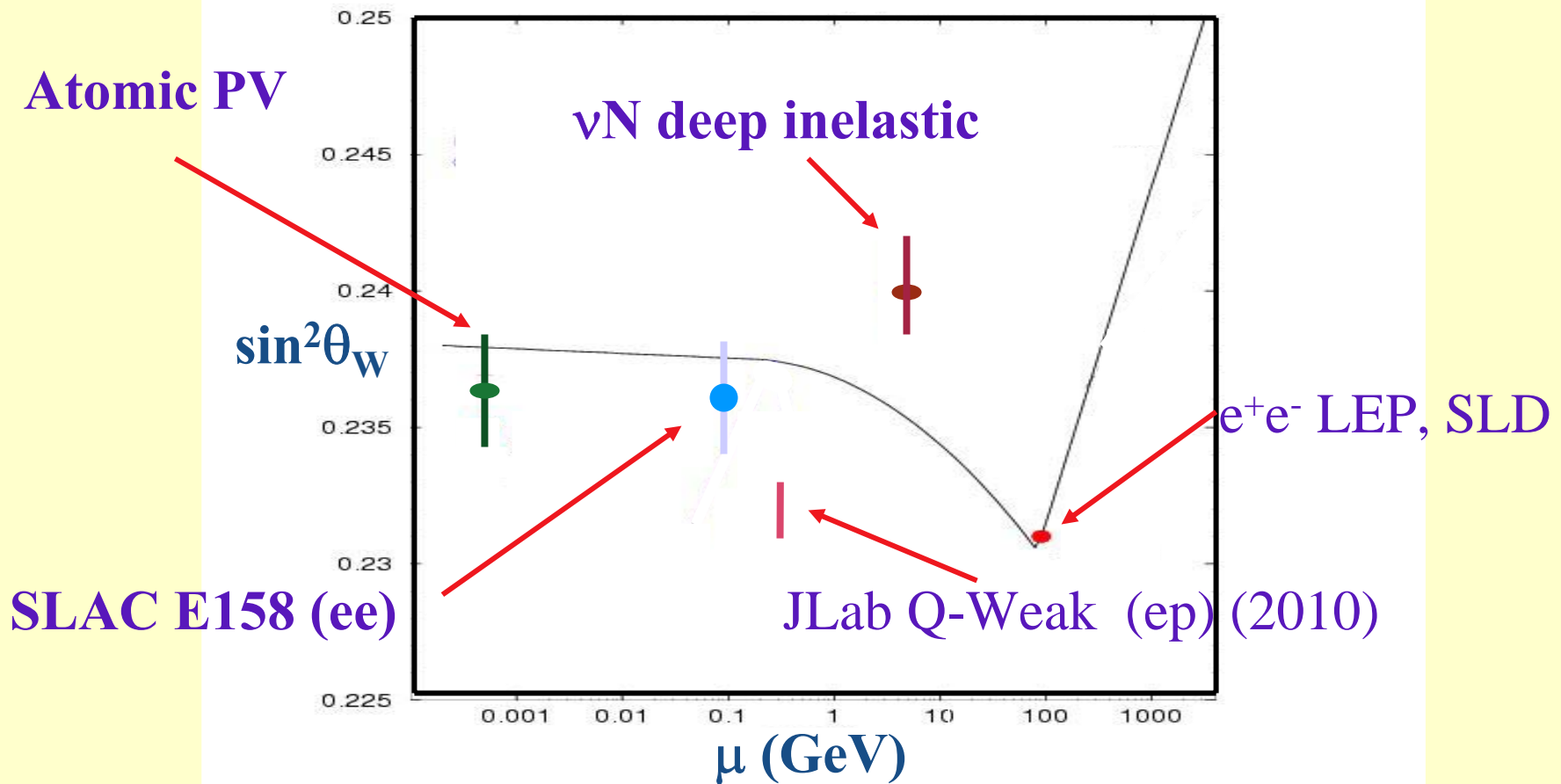
- Coupling constants in Standard Model are a function of the energy that it is being probed
 - Running of the couplings
- Standard Model predicts this energy dependence
- Can study by measuring neutral weak cross sections vs energy
 - $e^+ - e^-$, neutrino-nucleus, electron-nucleus

$$d\sigma = (\mathbf{EM} + \mathbf{Weak})^2 = \mathbf{EM}^2 + \mathbf{Weak}^2 + \mathbf{2EM \cdot Weak}$$

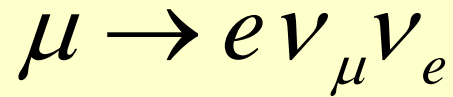
Parity Violating

Neutral Weak Phenomena - $\sin^2\theta_W$

Czarnecki, Marciano Erler.
Kurylov, Ramsey-Musolf



Muon Decay Distributions



- Energy dependence
- Angular dependence
- Called Michel parameters

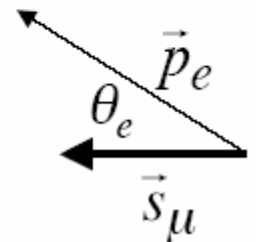
$$\frac{dN_e}{d\Omega_e dE_e} \propto x^2 \left[3 - 3x + \frac{2}{3} \rho(4x - 3) + 3\eta x_o \left(\frac{1-x}{x} \right) + P_{\mu} \xi \cos\theta_e \left(1 - x + \frac{2}{3} \delta(4x - 3) \right) \right]$$

Spectral shape in $x, \cos\theta_e$ is characterized in terms of four parameters -- ρ, η, ξ, δ

P_{μ} is the muon polarization

$$x \equiv \frac{E_e}{E_e^{\max}}$$

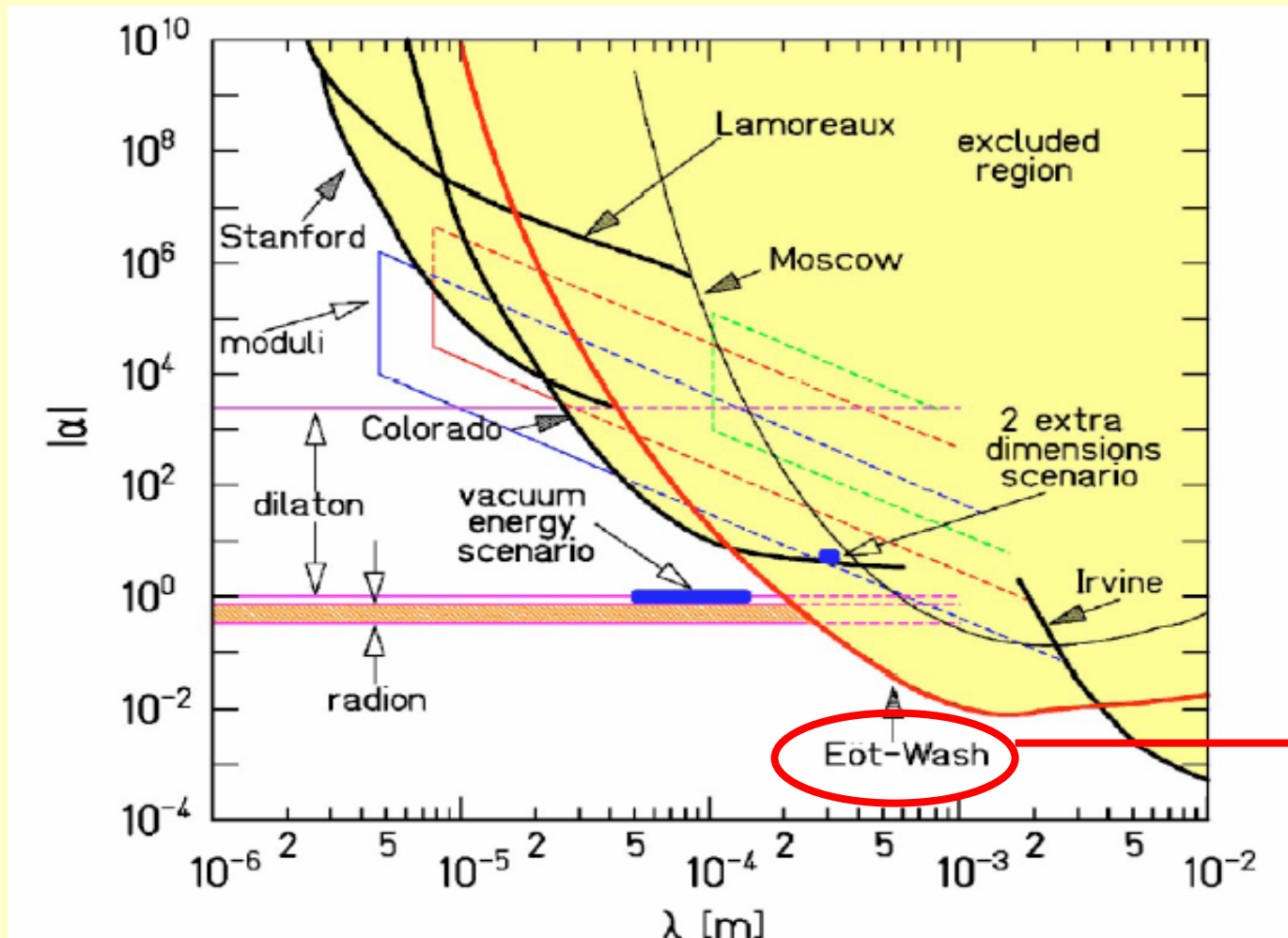
$$x_o \equiv \frac{m_e}{E_e^{\max}}$$



Summary of Michel Parameters

Parameter	Standard Model	Data
ρ	$\frac{3}{4}$	0.7509 ± 0.0010
η	0	0.001 ± 0.024
δ	$\frac{3}{4}$	0.7495 ± 0.0012
ξP_{μ}	1	$1.0027 \pm 0.0079 \pm 0.0030$

Short-distance Gravity



UW
Seattle

US Facilities for Fundamental Physics (not Neutrons)

- Low Energy Facilities
 - ATLAS, HRI BF, NSCL, ISAC, LBNL, Stonybrook & TAMU for 0^+-0^+ & other β -decay & EDM
- Medium Energy Facilities
 - JLAB for Q_{weak} , DIS-Parity
- Heavy-I on Facilities
 - BNL for $(g-2)_{\mu}$ and EDM
 - High Energy Facilities e.g. SLAC for Moeller asymmetry

Cold and Ultra-Cold Neutrons

- Cold neutrons $\sim 4 \text{ \AA}$
 - High flux from Cold Moderator (20K)
- Ultra-Cold Neutrons (UCN) $\sim 500 \text{ \AA}$
 - Boltzmann tail from Reactor cold moderator
 - Superthermal Converters
 - LHe, SD_2

What good are UCN?

- Can be trapped in certain materials
- Can be easily polarized ($B = 6\text{T}$)
- Allow precision measurements with free neutrons

US Facilities for Fundamental Physics (with Neutrons)

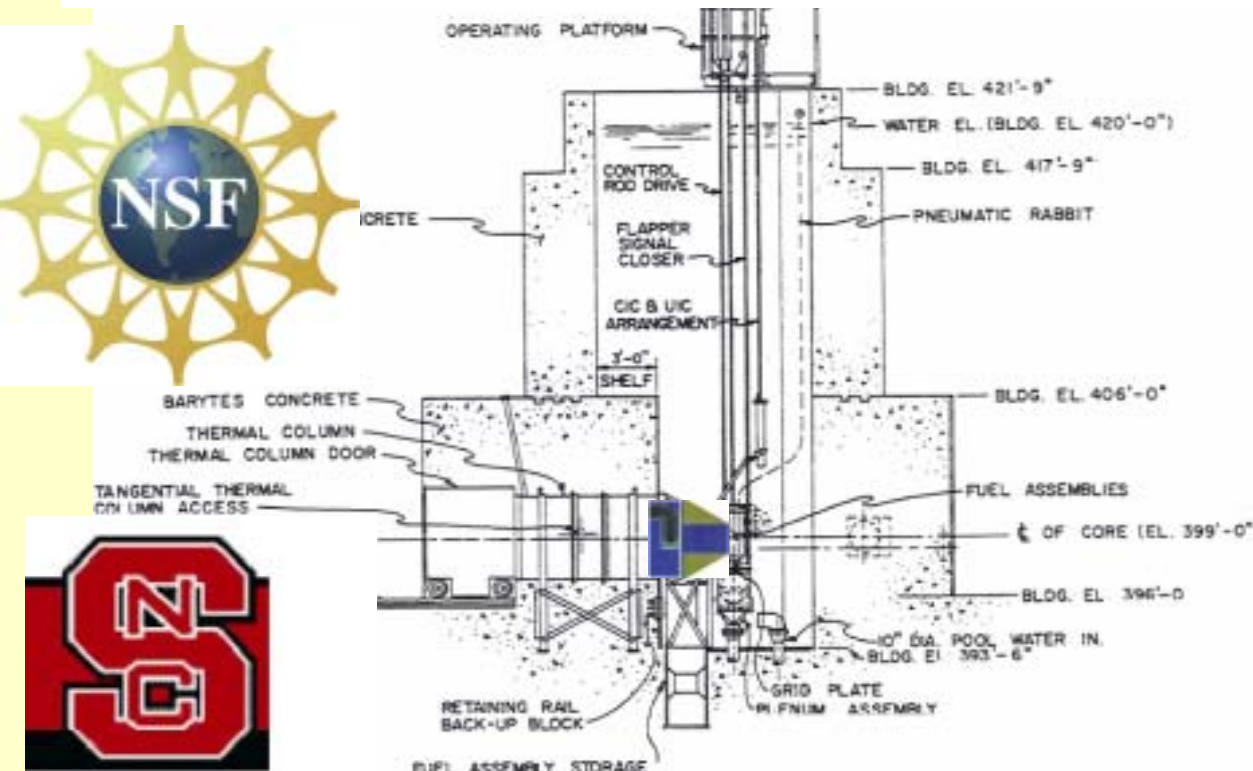
- LANSCE
 - Flight Path 12 (FP12) - Pulsed Cold Neutrons
 - Area B - UCN
- NIST
 - NG6 - Cold Neutron Lines
- SNS
 - Fundamental Neutron Physics Beamline (FNPB)
- Other Initiatives
 - PULSTAR Reactor @ NCSU
 - LENS Cold neutrons @ IUCF

Non-US Neutron Facilities

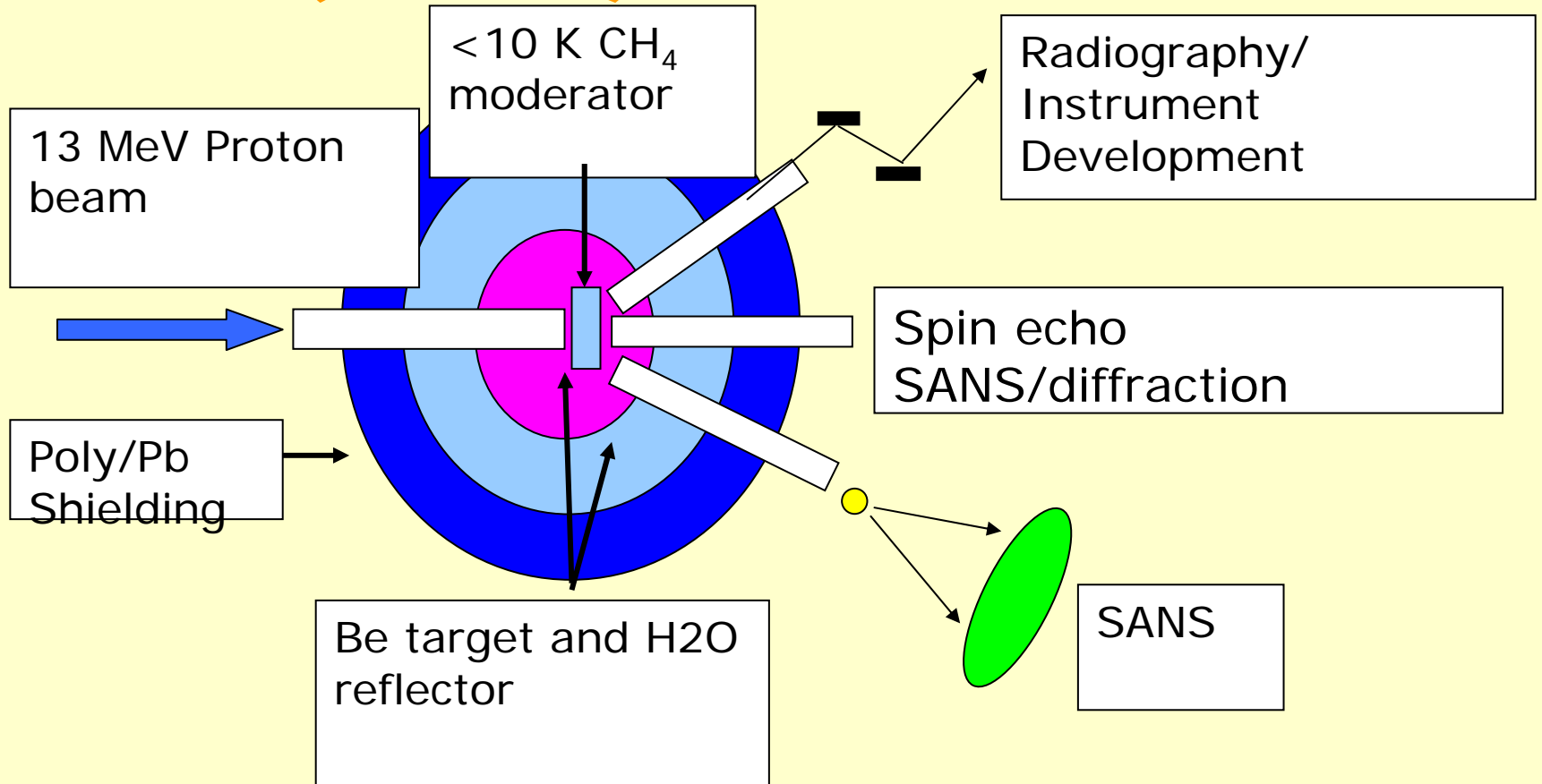
- ILL (Institut Laue-Langevain) – Cold & Ultra-Cold neutrons
 - Improved EDM experiment (Cryo-EDM)
 - Improved correlation experiments (<.5%)
- PSI – Spallation SD_2 UCN source
 - Proposal for EDM exists
- Germany - FRM-II 20MW research reactor
 - Working on SD_2 source insert for UCN
- Japan
 - Osaka UCN, JPARC?

NCSU PULSTAR UCN Source

- PULSTAR is a 1MW research reactor on NCSU campus
- UCN source expected to provide densities $> 1,000/\text{cm}^3$
- Dedicated to **nuclear physics** research
- Source construction funded by NSF (1.2M\$), operational by Jan. 2007
- Reactor operations funded by State of North Carolina, upgrades may be funded through DOE, INIE program



Low Energy Neutron Source (LENS) at Indiana



University-based pulsed cold neutron source

Vertical UCN beam possible, replace CH₄ → CH₄+solid O₂

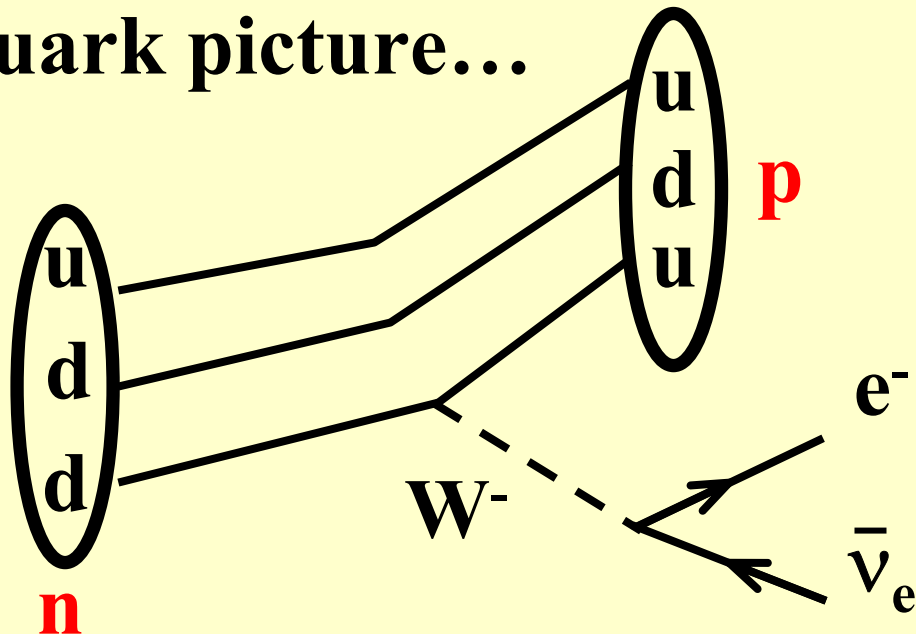
Neutron & Nuclear Beta Decay

$$\mathbf{n} \Rightarrow \mathbf{p} + \mathbf{e}^- + \bar{\nu}_e$$

$$\tau_n \approx 15 \text{ min.}, \quad t_{\frac{1}{2}} \approx 10 \text{ min.}$$

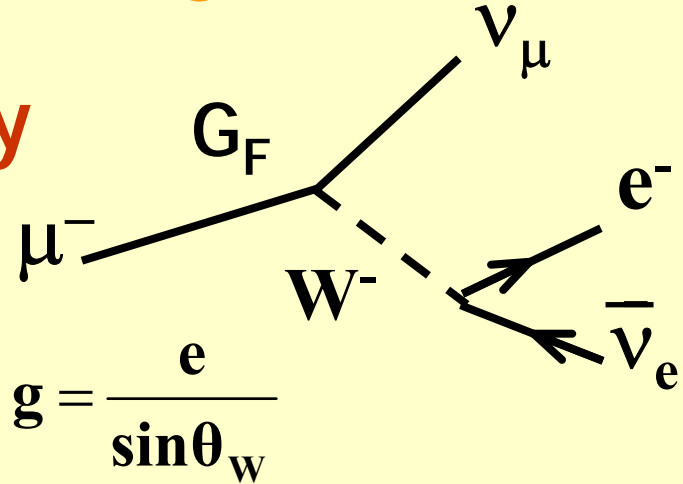
$$\sum_{\mathbf{p}, \mathbf{e}, \nu} \text{K.E.} = 0.78 \text{ MeV}$$

or in quark picture...



Weak Decays

- Consider muon decay



$$G_F = \frac{\sqrt{2}g^2}{8M_W^2}; \text{ since } m_\mu \ll M_W \text{ and } g = \frac{e}{\sin\theta_w}$$

$$= 1.16637(1) \cdot 10^{-5} (\hbar c)^3 \text{ GeV}^{-2}$$

- Can calculate muon lifetime

$$\frac{1}{\tau_\mu} = \int d\Gamma_\mu \propto \int |M|^2 d\text{LIPS}$$

LIPS = Lorentz Invariant Phase Space

$$M = \frac{G_F}{\sqrt{2}} \bar{u}_{\nu_\mu} \gamma^\mu (1 - \gamma^5) u_\mu \bar{u}_e \gamma_\mu (1 - \gamma^5) v_{\nu_e}$$

Weak Decays

- Ignoring small kinematic corrections

$$\tau_{\mu} \approx \frac{192\pi^3 \hbar^7}{G_F^2 m_{\mu}^5 c^4} = 2.187 \times 10^{-6} \text{ sec}$$

$$\tau_{\mu}^{\text{exp}} = 2.19703(4) \times 10^{-6} \text{ sec}$$

- Key feature of Electroweak Standard Model is **UNIVERSALITY** - lepton and quark weak interactions are identical ... modulo the CKM matrix. Thus

$$M_{n-p} = \frac{1}{\sqrt{2}} \bar{u}_p \gamma^{\mu} (G_V - G_A \gamma^5) u_n \bar{u}_e \gamma_{\mu} (1 - \gamma^5) \nu_{\nu_e}$$

$$\tau_n \approx \frac{2\pi^3 \hbar^7}{G_F^2 m_e^5 c^4} = 8,611 \text{ sec} \quad \tau_n^{\text{exp}} = 885(1) \text{ sec}$$

Neutron Decay in the Standard Model

$$\tau_n = \left(\frac{2\pi^3 \hbar^7}{m_e^5 c^4} \right) \left(\frac{1}{G_F^2 |V_{ud}|^2} \right) \left[\frac{1}{1 + 3(G_A/G_V)^2} \right] \frac{1}{f(1 + \Delta_R)}$$

G_F = Fermi Constant (known from μ decay)

V_{ud} = up-down quark weak coupling (more later)

G_A = Axial vector weak coupling constant

G_V = Vector weak coupling constant

f = phase space integral

Δ_R = Electroweak radiative correction

Note: Z^0 Boson ($M=91$ GeV) gives 2% correction!

G_A/G_V from parity violating decay asymmetry in n decay

$$A = \frac{-2\lambda(1+\lambda)}{1+3\lambda^2}, \quad \lambda = \frac{G_A}{G_V}$$

Polarized Neutron decay

$$d\Gamma = \Gamma_n \left(1 + \mathbf{b} \frac{\mathbf{m}_e}{E_e} + \mathbf{a} \frac{\vec{\mathbf{p}}_e \cdot \vec{\mathbf{p}}_v}{E_e E_v} + \mathbf{A} \frac{\vec{\sigma}_n \cdot \vec{\mathbf{p}}_e}{E_e} + \mathbf{B} \frac{\vec{\sigma}_n \cdot \vec{\mathbf{p}}_v}{E_v} + \mathbf{D} \vec{\sigma}_n \cdot \frac{\vec{\mathbf{p}}_e \times \vec{\mathbf{p}}_v}{E_e} \right)$$

- $\Gamma_n = 1/\tau_n$ total decay rate (depends on G_A and G_V)
- Correlations \mathbf{a} , \mathbf{A} and \mathbf{B} depend on G_A and G_V
- Coefficient \mathbf{b} requires S or T interaction
- Correlation \mathbf{D} violates Time Reversal Invariance

Must measure two observables to extract G_A
and G_V (eg. G_n and \mathbf{A})

Precision neutron decay measurements

- Neutron lifetime essential in Big-Bang Nucleosynthesis Calculations
- Can provide most precise measurement of V_{ud}

Weak eigenstates

$$\begin{pmatrix} \mathbf{d}_w \\ \mathbf{s}_w \\ \mathbf{b}_w \end{pmatrix} = \begin{pmatrix} V_{ud} & V_{us} & V_{ub} \\ V_{cd} & V_{cs} & V_{cb} \\ V_{td} & V_{ts} & V_{tb} \end{pmatrix} \begin{pmatrix} \mathbf{d} \\ \mathbf{s} \\ \mathbf{b} \end{pmatrix}$$

Mass eigenstates

Single complex phase
is possible
(gives "CP" Violation
-more later)

- $< 0.3\%$ measurements can be sensitive to new physics (from loops in electroweak field theory)

a.k.a. Radiative Corrections

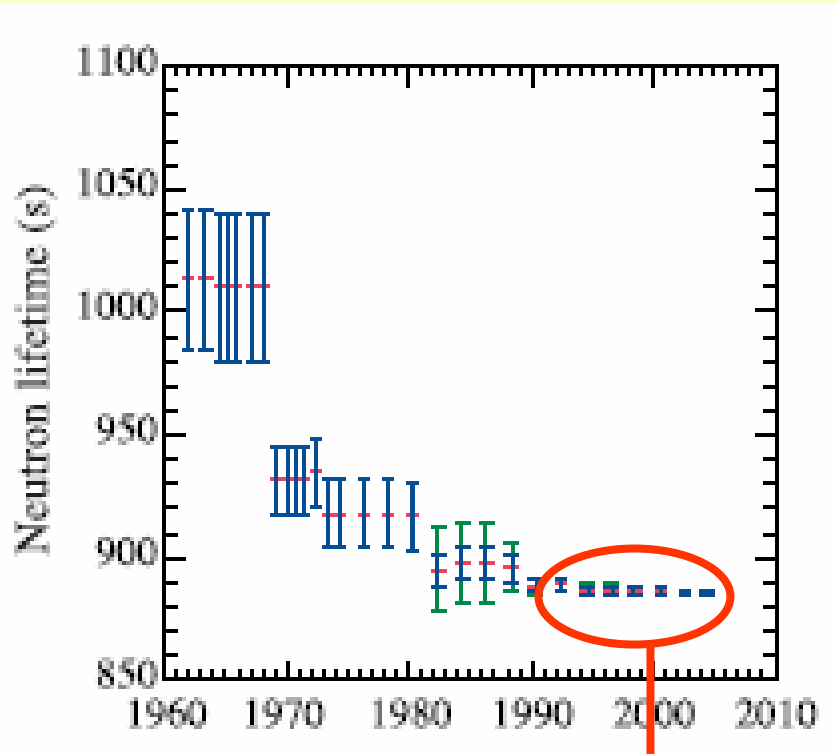
$T \simeq 0.1$ MeV; nuclei can then begin to form without being immediately photo-dissociated again. Only 2-body reactions such as $D(p, \gamma)^3\text{He}$, $^3\text{He}(D, p)^4\text{He}$, are important because the density has become rather low by this time.

Nearly all the surviving neutrons when nucleosynthesis begins end up bound in the most stable light element ^4He . Heavier nuclei do not form in any significant quantity both because of the absence of stable nuclei with mass number 5 or 8 (which impedes nucleosynthesis via $n^4\text{He}$, $p^4\text{He}$ or $^4\text{He}^4\text{He}$ reactions) and the large Coulomb barriers for reactions such as $T(^4\text{He}, \gamma)^7\text{Li}$ and $^3\text{He}(^4\text{He}, \gamma)^7\text{Be}$. Hence the primordial mass fraction of ^4He , conventionally referred to as Y_{P} , can be estimated by the simple counting argument

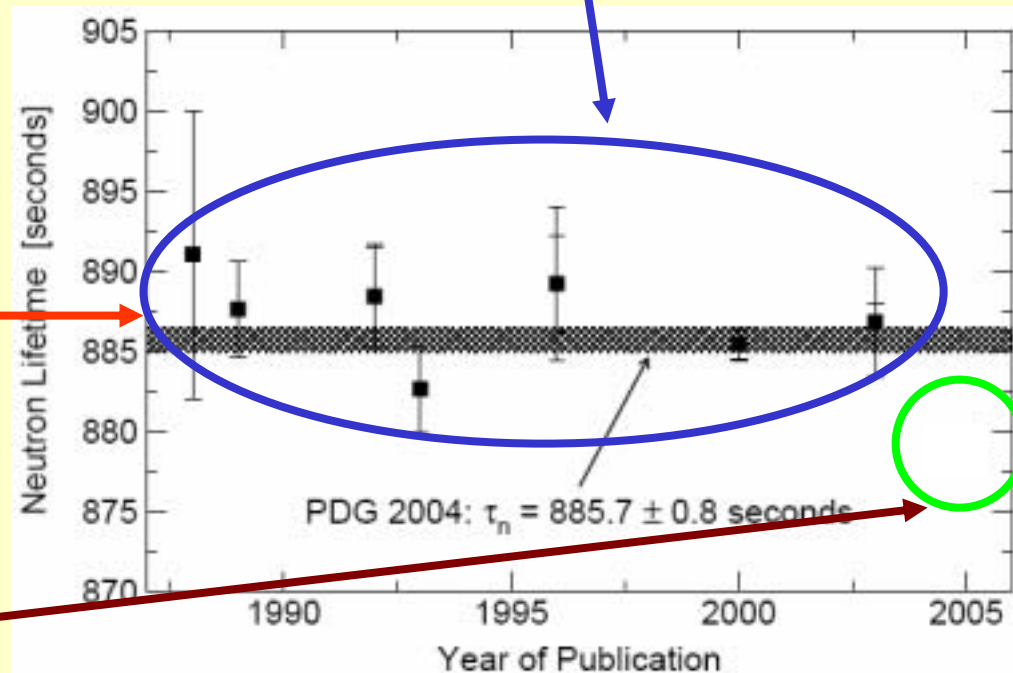
$$Y_{\text{P}} = \frac{2(n/p)}{1 + n/p} \simeq 0.25 . \quad \text{Primordial He Abundance} \quad (20.1)$$

There is little sensitivity here to the actual nuclear reaction rates, which are however important in determining the other “left-over” abundances: D and ^3He at the level of a few times 10^{-5} by number relative to H, and $^7\text{Li}/\text{H}$ at the level of about 10^{-10} (when η_{10} is in the range 1–10). These values can be understood in terms of approximate analytic arguments [8]. The experimental parameter most important in determining Y_{P} is the neutron lifetime, τ_n , which normalizes (the inverse of) Γ_{n-p} . (This is not fully determined by G_{P} alone since neutrons and protons also have strong interactions, the effects of which cannot be calculated very precisely.) The experimental uncertainty in τ_n used to be a source of concern but has recently been reduced substantially: $\tau_n = 885.7 \pm 0.8$ s.

Neutron Lifetime versus Year

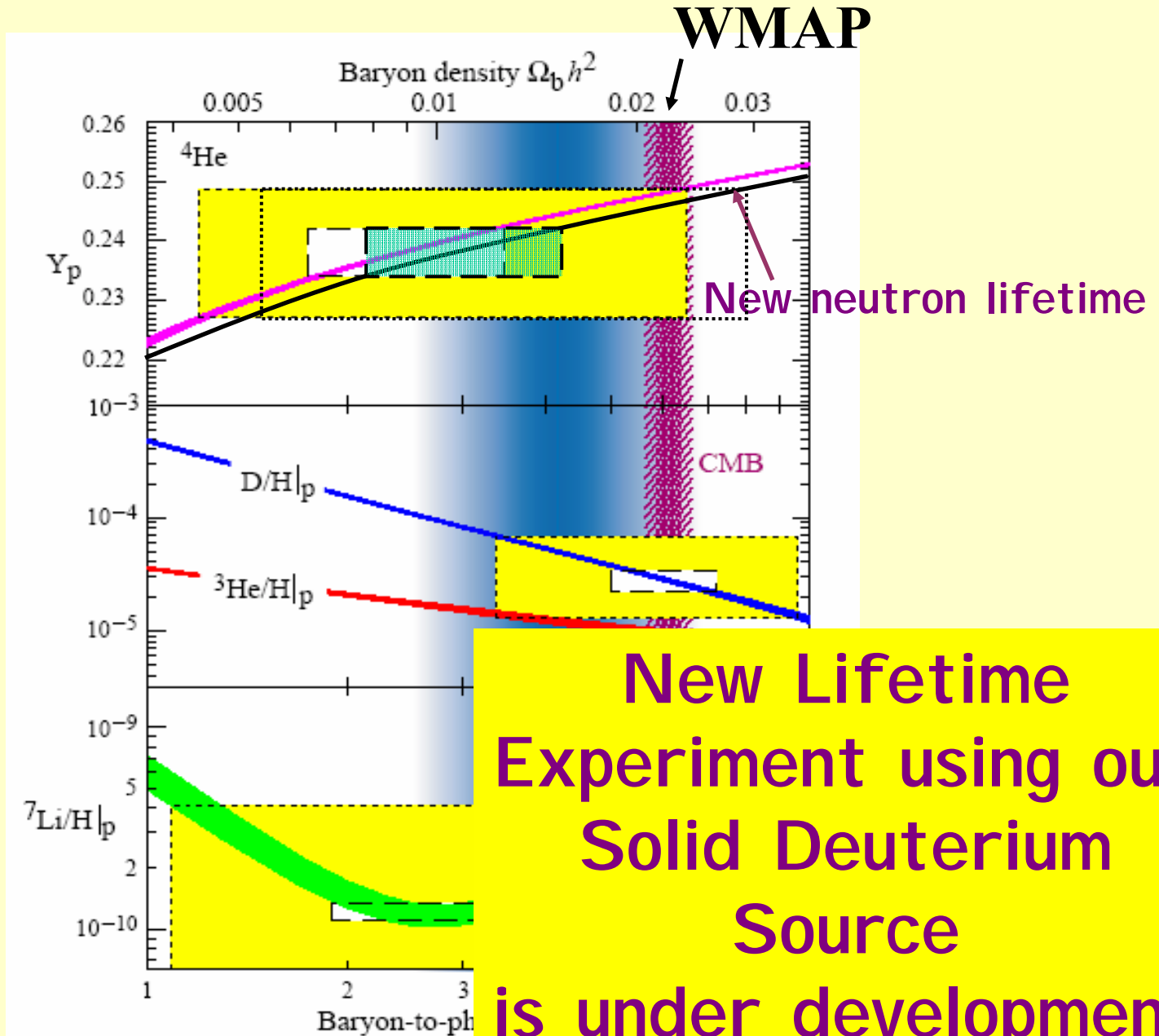


Data points used by Particle Data Group (PDG) 2004 for averaging



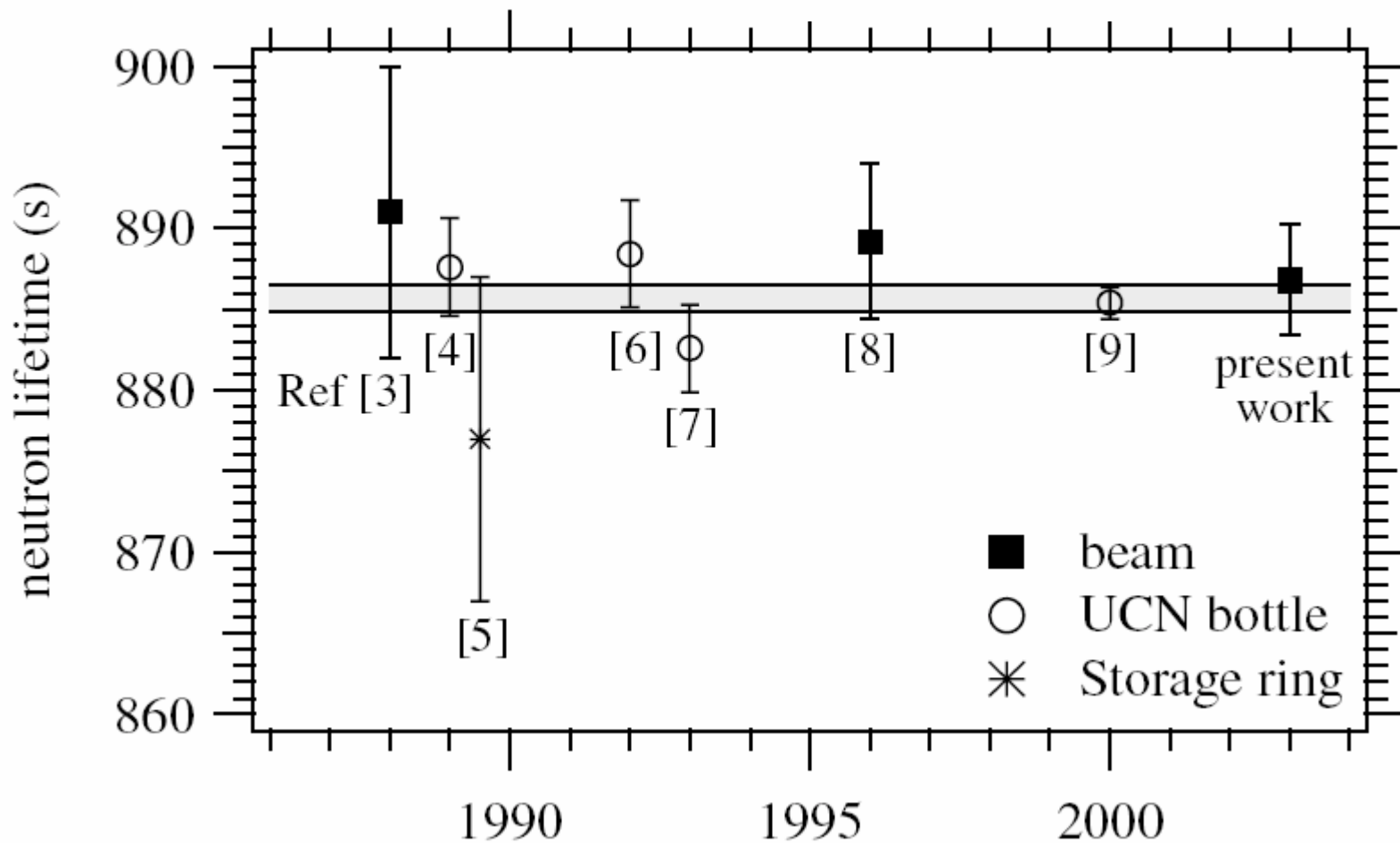
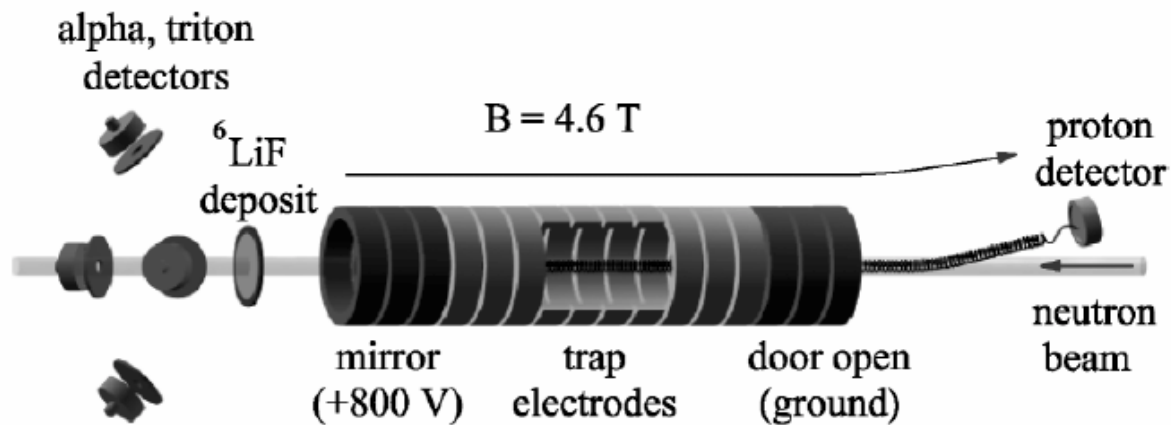
Serebrov *et al.*,
Phys. Lett. B 605, 72 (2005)
($878.5 \pm 0.7 \pm 0.3$) seconds

Big-Bang Nucleosynthesis Constraints



New Lifetime
Experiment using our
Solid Deuterium
Source
is under development

Neutron τ_n @ NIST



New Neutron Lifetime Measurement UCN @ ILL

Need More Lifetime experiments

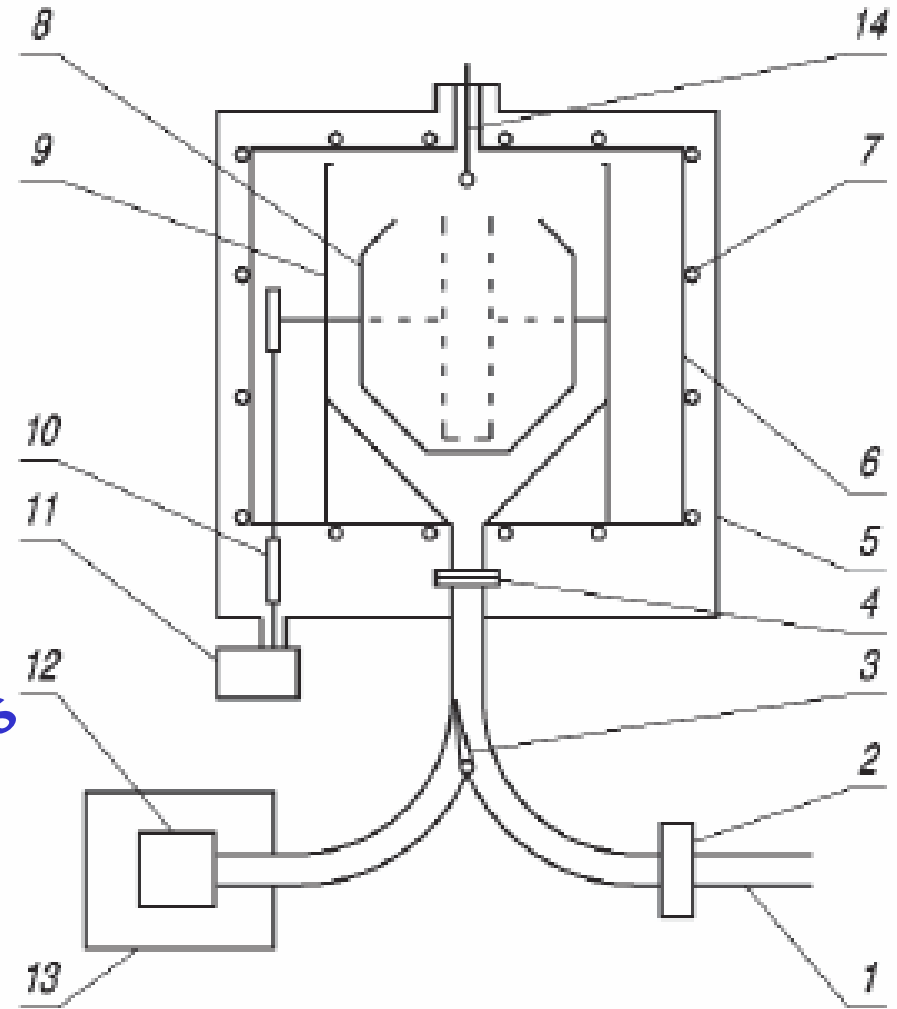
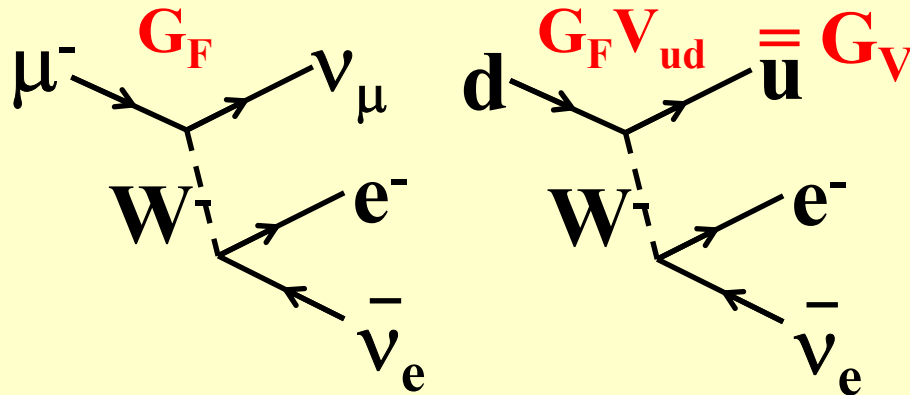


Fig. 1. The Scheme of "Gravitrapp", the gravitational UCN storage system. 1: neutron guide from UCN Turbine; 2: UCN inlet valve; 3: beam distribution flap valve (shown in the filling position); 4: connection unit; 5: "high" vacuum volume; 6: "rough" vacuum volume; 7: cooling coils; 8: UCN storage trap (the narrow cylindrical trap is shown by a dashed line); 9: cryostat; 10: mechanics for trap rotation; 11: stepping motor; 12: UCN detector; 13: detector shielding; 14: evaporator.

V_{ud} from β -decay

- G_V is related to Fermi coupling G_F via V_{ud}



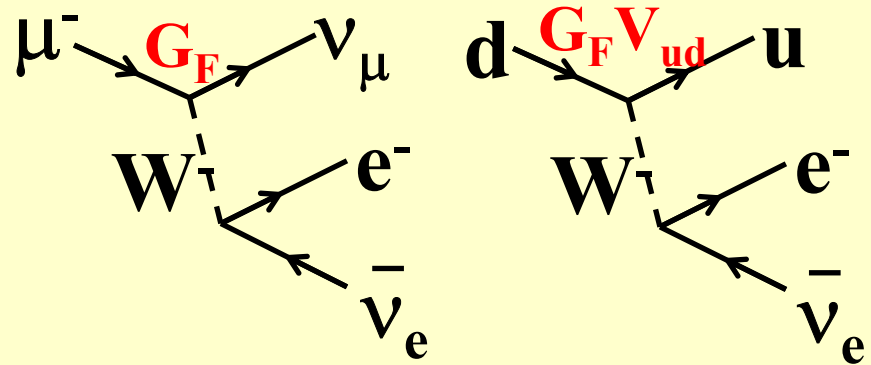
- G_V measured in β -decay ($u \rightarrow d$)
 - $0^+ - 0^+$ nuclear decay (must include nuclear corrections)
 - neutron decay (must extract G_V and G_A)

Sensitivity to New Physics?

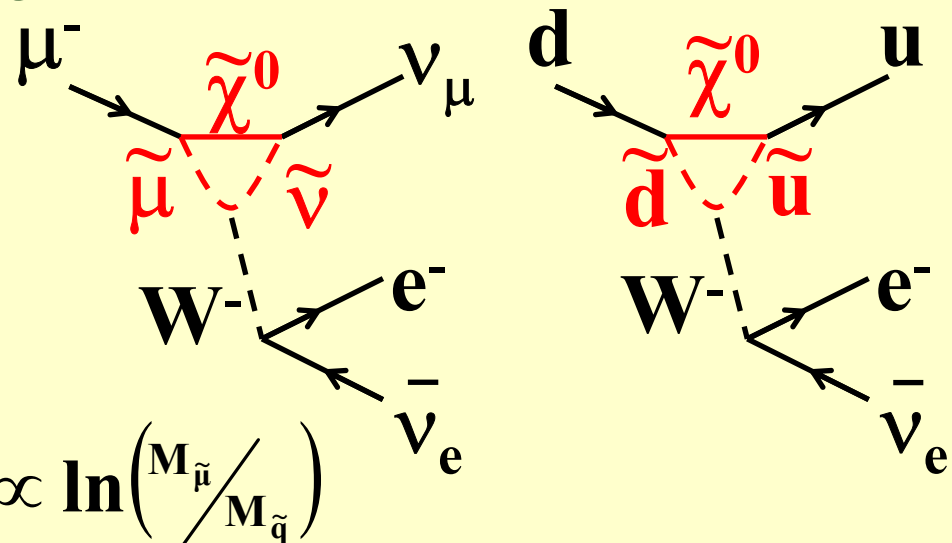
Kurylov&Ramsey-Musolf

Phys. Rev. Lett. 88, 071804 (2002)

- V_{ud} in Standard Model
(from μ vs. β -decay)



- Supersymmetric particles produce loop corrections



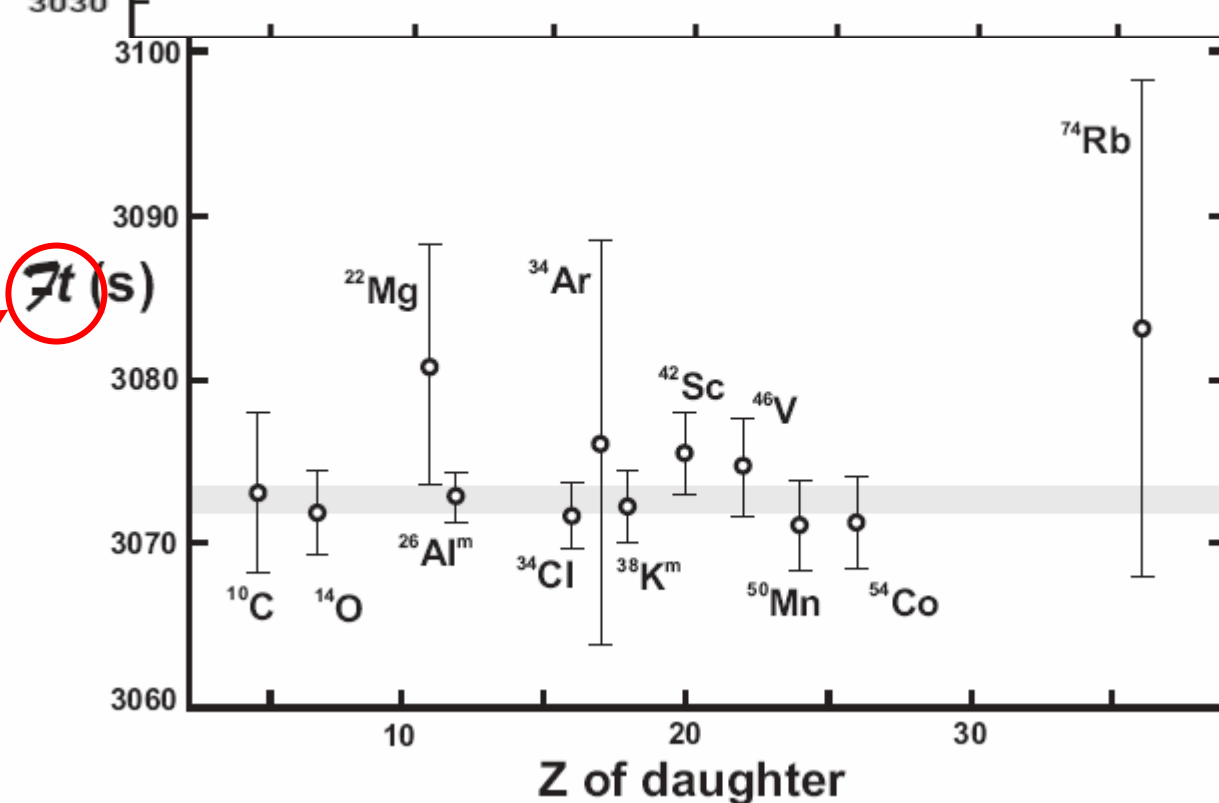
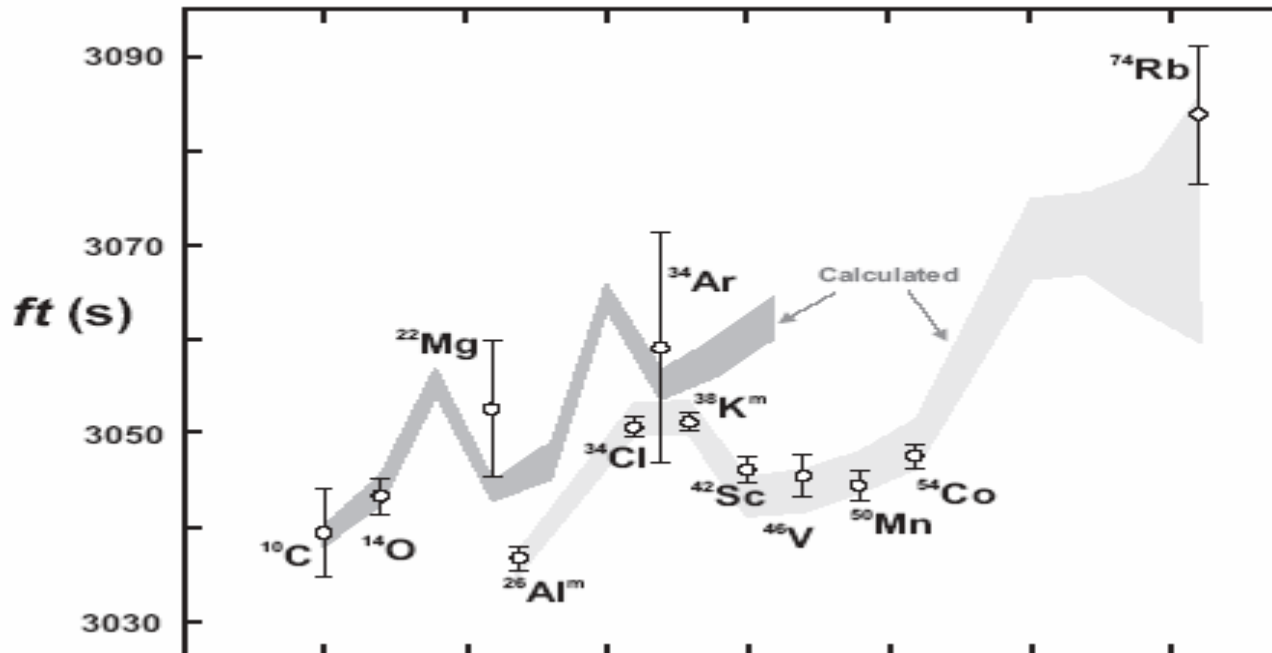
$$V_{ud}^{\text{super}} = V_{ud}^{\text{SM}} (1 - \Delta_{\text{loop}}) : \Delta_{\text{loop}} \propto \ln\left(\frac{M_{\tilde{\mu}}}{M_{\tilde{q}}}\right)$$

Nuclear β -Decay

$$0^+ - 0^+$$

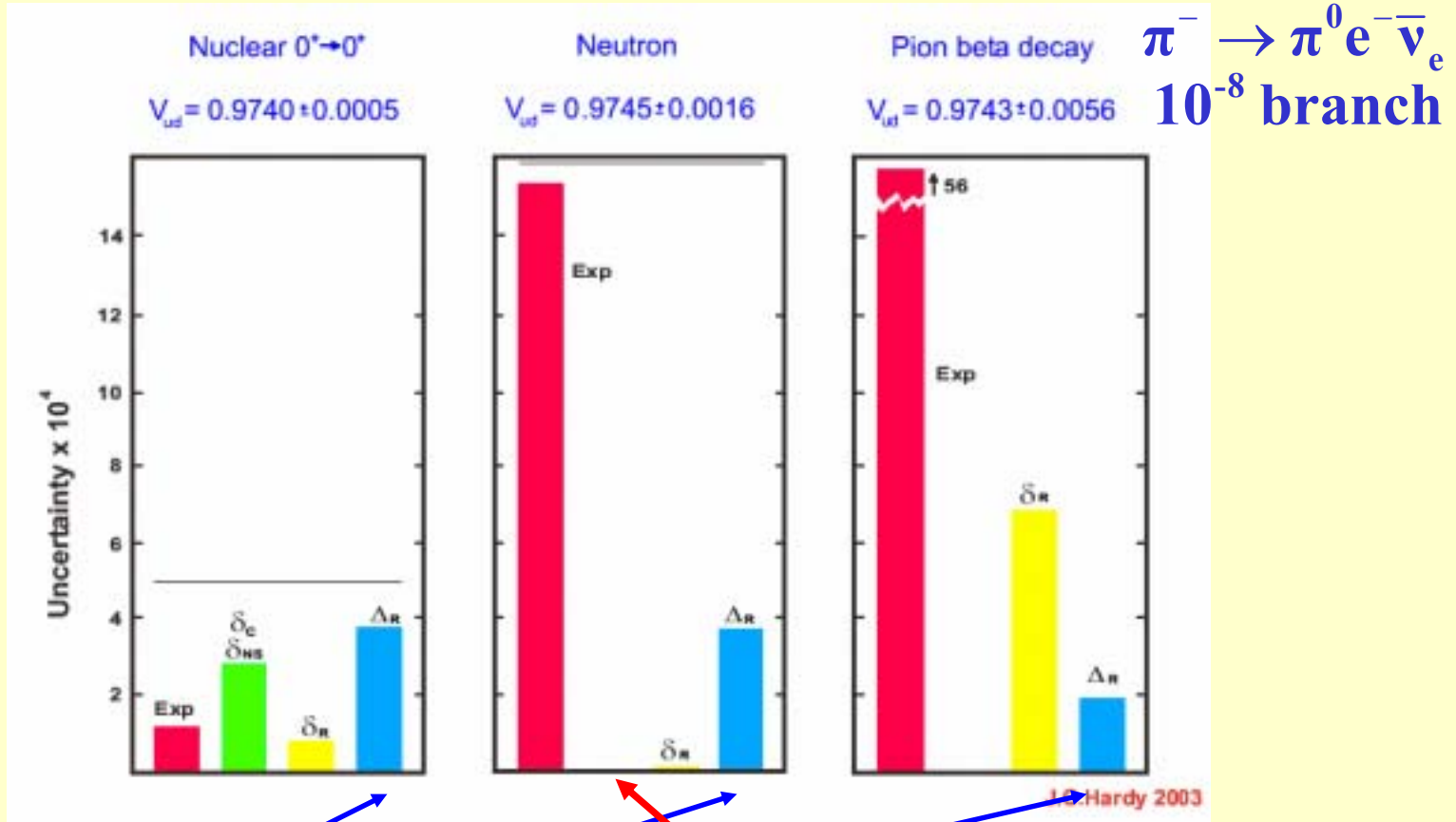
f = phase space
t = half-life

$$ft \propto \frac{1}{G_V^2}$$



Including
Nuclear
Corrections

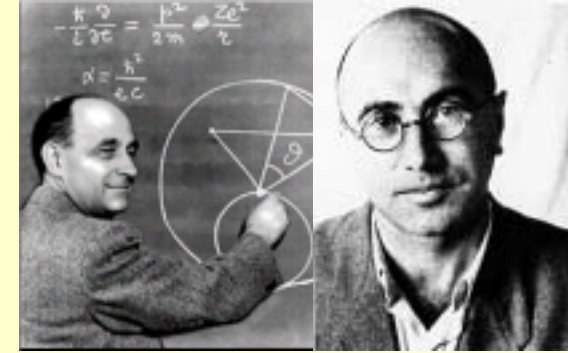
Uncertainties in V_{ud}



Electroweak corrections
(Z^0 and hadron loops)

Improved neutron
Experiments needed

Ultra-Cold Neutrons (UCN) (Fermi/Zeldovich)



- What are UCN ?

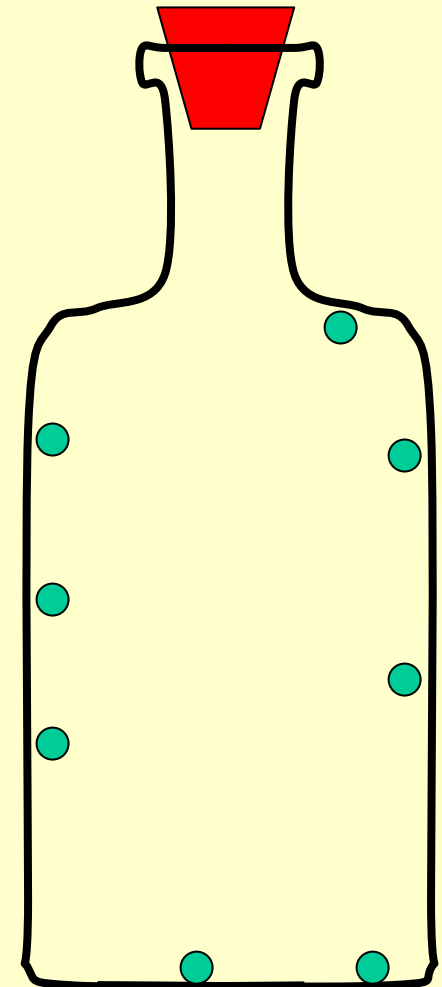
- Very slow neutrons

- $(v < 8 \text{ m/s} \rightarrow \lambda > 500 \text{ \AA})$

- that cannot penetrate into
certain materials



- Neutrons can be
trapped in bottles
or by magnetic
field



But... nuclear force is attractive at low energies. Where does repulsion come from?

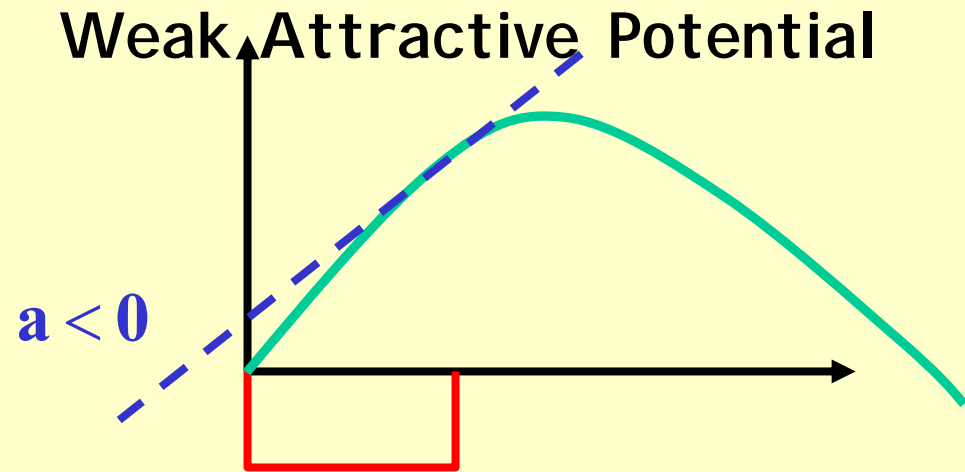
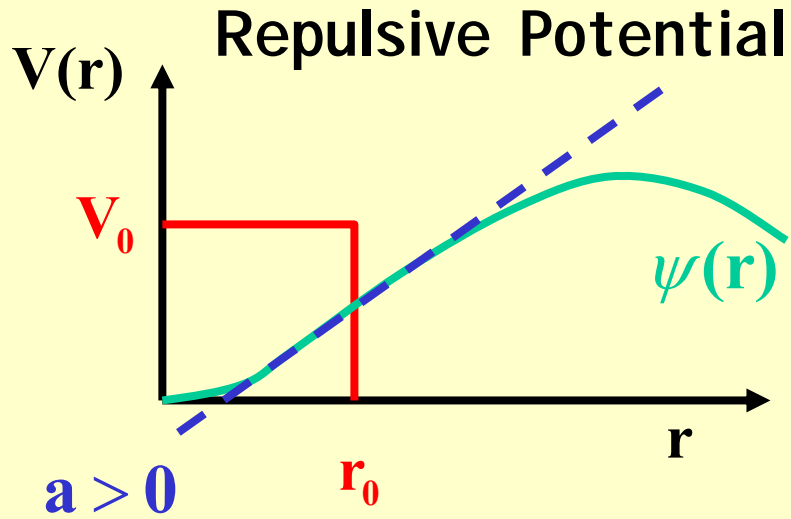
Recall (for short-range potential):

At low energies ($kr_0 \ll 1$; eg s-wave) elastic scattering determined solely by scattering length a

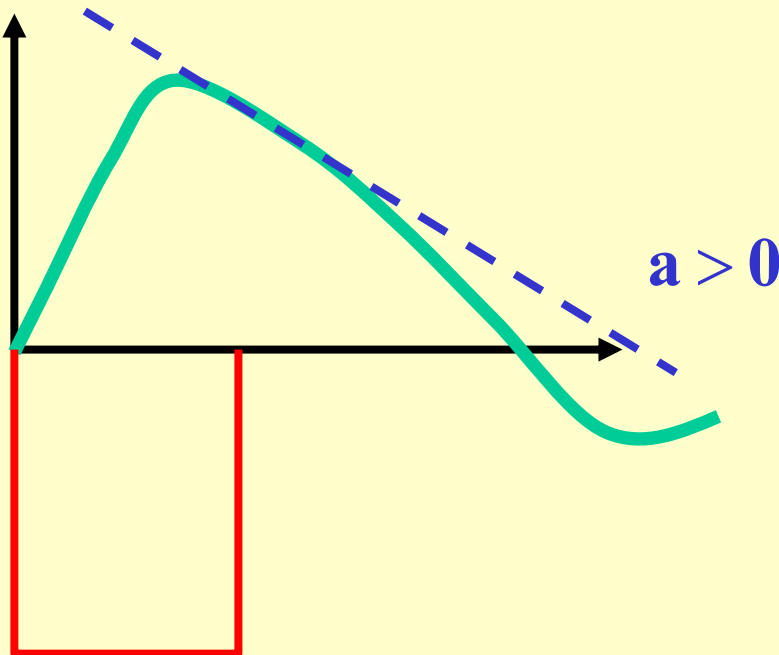
For $k \rightarrow 0$

$$\sigma_{\text{elas}} = 4\pi a^2$$

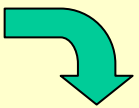
Scattering Length



Strong Attractive Potential



Thus many different V_0
and r_0 can give same a



eg.
$$V(\vec{r}) = \frac{a\hbar^2}{2m_n} (4\pi) \delta(\vec{r})$$

Fermi Pseudopotential

Then for many nuclei
in a solid:

$$\rightarrow V(\vec{r}) = \frac{4\pi\hbar^2}{2m_n} \sum_i a_i \delta(\vec{r} - \vec{r}_i')$$

And if $a_i = a$ and
 $1 \gg \text{atomic spacing}$:

$$\begin{aligned} \rightarrow V(\vec{r}) &= \frac{4\pi a \hbar^2}{2m_n} N_0 \int \frac{d^3r'}{V} \delta(\vec{r} - \vec{r}') \\ &= \frac{2\pi a \hbar^2 n_0}{m_n} \theta(\vec{r} \notin V) \equiv V_0 \theta(\vec{r} \notin V) \end{aligned}$$

Fermi Pseudopotential

Potential step analogous to index of refraction in optics

with

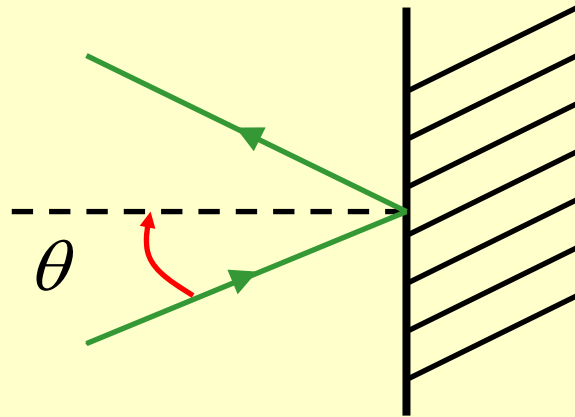
$$n = \sqrt{1 - \frac{V_0}{E_n}}$$

$$V_0 = \frac{2\pi a \hbar^2 n_0}{m_n}$$

Neutron kinetic energy

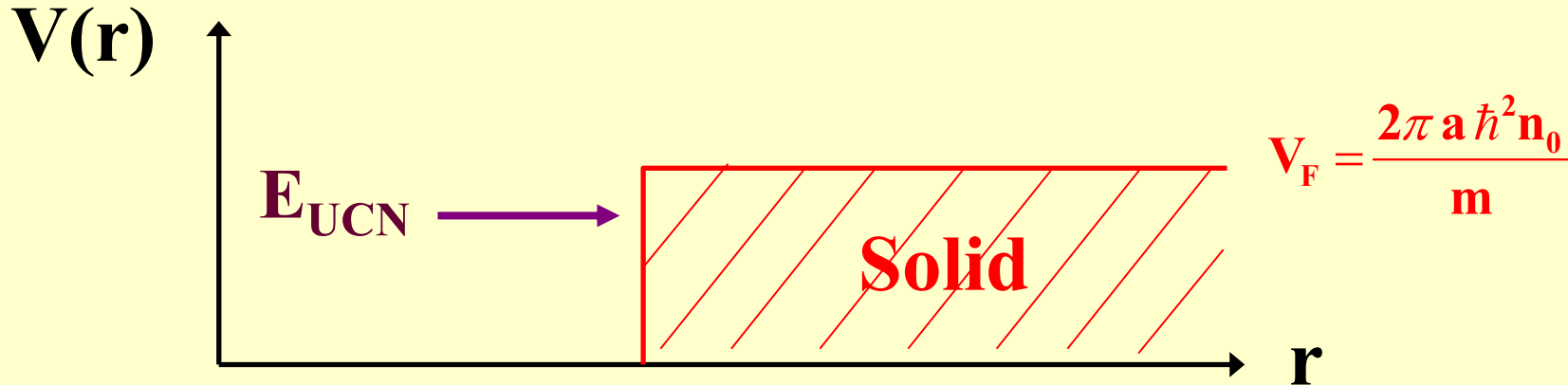
And if $a > 0$ can have total *external* reflection

$$n_{\text{out}} = 1$$



$$n_{\text{in}} = \sqrt{1 - \frac{V_0}{E}}$$

Fermi Pseudo-potential



The coherent nuclear potential can lead to repulsive pseudopotential (Fermi potential) for $a > 0$

For $E_{\text{UCN}} < V_F$, UCN are trapped

**Attractive potential can also lead to neutron absorption
but often $L_{\text{mfp}} \gg \lambda_n$ ($\sim 10^{-5}$ probability per bounce)**

Typical Fermi Potentials

Material	V_F (neV)
Al	54
^{58}Ni	350
Ti	- 48
Graphite	180
Stainless Steel	188
Diamond-like Carbon	282

neutron velocity
 $v_n \sim 8 \text{ m/s}$

Magnetic Bottles also possible

- B-field and neutron magnet moment produces a potential

$$V = -\vec{\mu} \cdot \vec{B} \quad (\text{Note: } \mu_n < 0)$$

- Thus can produce a 3D potential well at a B-field minimum

Ioffe Trap

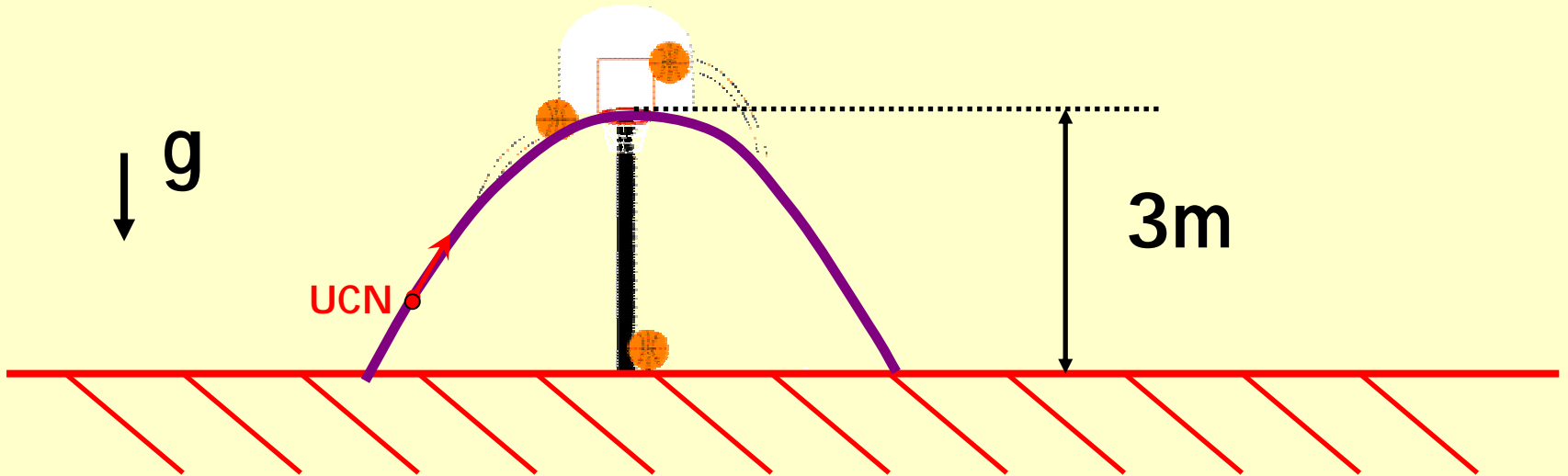


Traps one spin state:
“Low field seekers”
Spin anti-aligned with
magnetic field

For $v_n < 8$ m/s need $B < 6$ T

UCN Properties

- $v_{\text{UCN}} \leq 8 \text{ m/s}$ ($\approx 18 \text{ mi/hr}$)
- $\lambda_{\text{UCN}} \geq 500 \text{ \AA}$ (50 nm)
- " T_{UCN} " $\leq 4 \text{ mK}$
- Maximum height due to gravity : $\sim 3 \text{ m}$



How to make UCN?

- **Conventional Approach:**

- Start with neutrons from nuclear reactor core

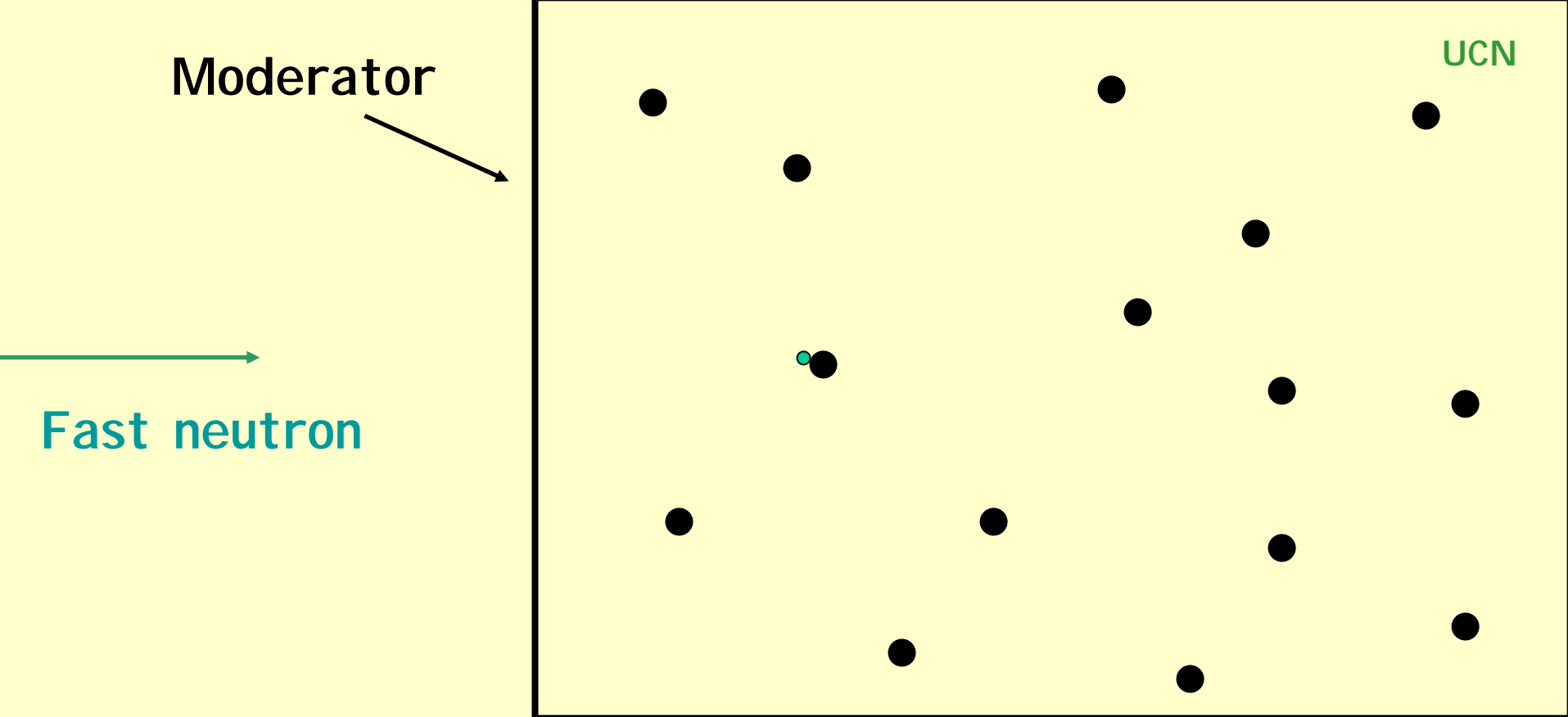
$$E_n \approx 5 - 10 \text{ MeV}$$

$$v_n \approx 4 \cdot 10^7 \text{ m/s}$$

- Use collisions with nuclei to slow down neutrons

**Some of neutron's energy lost
to nuclear recoil in each collision**

Gives a Maxwell-Boltzmann Distribution



After 20-100 collisions $E_n \sim 1/40$ eV (Room Temp)
 Slowing down takes $\sim 100 \mu\text{s}$

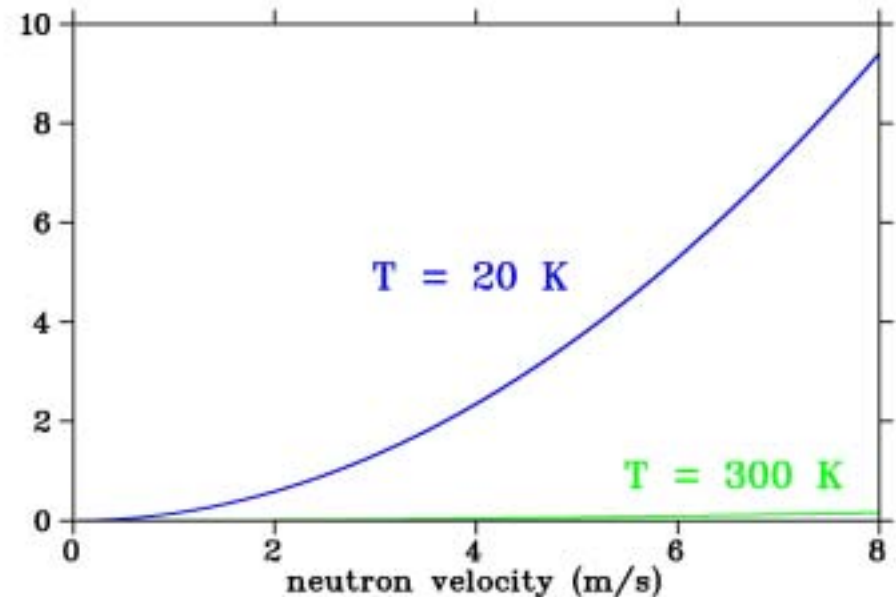
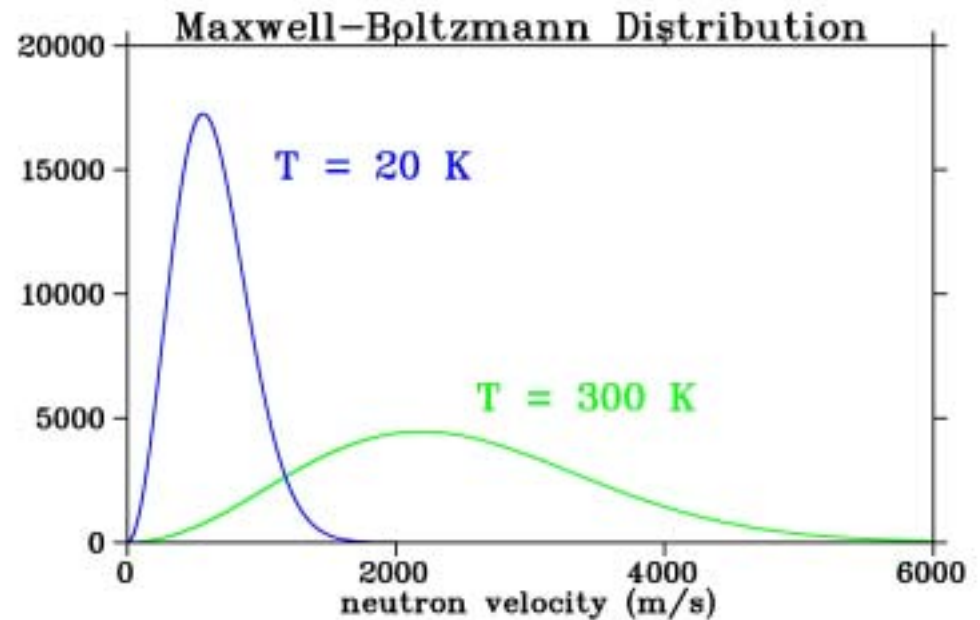
Maxwell-Boltzmann Distribution

$$P(E) \propto E^2 e^{-\frac{E}{kT}}$$

But...

- Only small fraction of neutron distribution is UCN

$\sim 10^{-8}$ for $T = 300\text{ K}$
 $\sim 10^{-6}$ for $T = 20\text{ K}$ (liquid H_2)

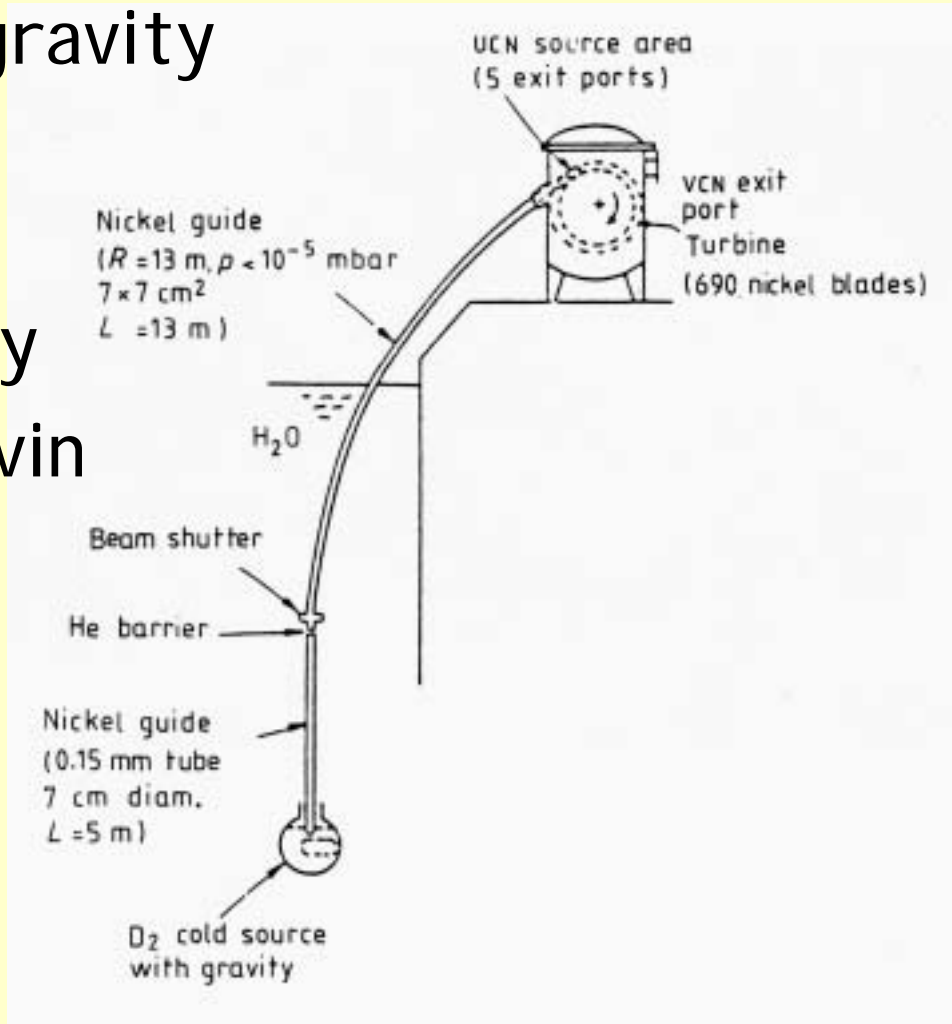


- Can improve some via gravity and moving turbines

- Previous record density at Institut Laue-Langevin (ILL) reactor in Grenoble

≈ 40 UCN/cm³ stored in bottle (1971)

Best vacuum on earth
 $\sim 10^4$ atoms/cm³



Can we make more?